## Notes on

### 1.63 Advanced Environmental Fluid Mechanics <br> Instructor: C. C. Mei, 2001 <br> ccmei@mit.edu, 16172532994

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2-5Stokes.tex

### 2.5 Stokes flow past a sphere

[Refs]
Lamb: Hydrodynamics
Acheson : Elementary Fluid Dynamics, p. 223 ff
One of the fundamental results in low Reynolds hydrodynamics is the Stokes solution for steady flow past a small sphere. The apllicatiuon range widely form the determination of electron charges to the physics of aerosols.

The continuity equation reads

$$
\begin{equation*}
\nabla \cdot \vec{q}=0 \tag{2.5.1}
\end{equation*}
$$

With inertia neglected, the approximate momentum equation is

$$
\begin{equation*}
0=-\frac{\nabla p}{\rho}+\nu \nabla^{2} \vec{q} \tag{2.5.2}
\end{equation*}
$$

Physically, the presssure gradient drives the flow by overcoming viscous resistence, but does affect the fluid inertia significantly.

Refering to Figure 2.5 for the spherical coordinate system $(r, \theta, \phi)$. Let the ambient velocity be upward and along the polar $(z)$ axis: $(u, v, w)=(0,0, W)$. Axial symmetry demands

$$
\frac{\partial}{\partial \phi}=0, \quad \text { and } \quad \vec{q}=\left(q_{r}(r, \theta), q_{\theta}(r, \theta), 0\right)
$$

Eq. (2.5.1) becomes

$$
\begin{equation*}
\frac{1}{r^{2}} \frac{\partial}{\partial r}\left(r^{2} q_{r}\right)+\frac{1}{r} \frac{\partial}{\partial \theta}\left(q_{\theta} \sin \theta\right)=0 \tag{2.5.3}
\end{equation*}
$$

As in the case of rectangular coordinates, we define the stream function $\psi$ to satisify the continuity equation (2.5.3) identically

$$
\begin{equation*}
q_{r}=\frac{1}{r^{2} \sin \theta} \frac{\partial \psi}{\partial \theta}, \quad q_{\theta}=-\frac{1}{r \sin \theta} \frac{\partial \psi}{\partial r} \tag{2.5.4}
\end{equation*}
$$

At infinity, the uniform velocity $W$ along $z$ axis can be decomposed into radial and polar components

$$
\begin{equation*}
q_{r}=W \cos \theta=\frac{1}{r^{2} \sin \theta} \frac{\partial \psi}{\partial \theta}, \quad q_{\theta}=-W \sin \theta=-\frac{1}{r \sin \theta} \frac{\partial \psi}{\partial r}, \quad r \sim \infty \tag{2.5.5}
\end{equation*}
$$



Figure 2.5.1: The spherical coordinates

The corresponding stream function at infinity follows by integration

$$
\begin{equation*}
\psi=\frac{W}{2} r^{2} \sin ^{2} \theta, \quad r \sim \infty \tag{2.5.6}
\end{equation*}
$$

Using the vector identity

$$
\begin{equation*}
\nabla \times(\nabla \times \vec{q})=\nabla(\nabla \cdot \vec{q})-\nabla^{2} \vec{q} \tag{2.5.7}
\end{equation*}
$$

and (2.5.1), we get

$$
\begin{equation*}
\nabla^{2} \vec{q}=-\nabla \times(\nabla \times \vec{q})=-\nabla \times \vec{\zeta} \tag{2.5.8}
\end{equation*}
$$

Taking the curl of (2.5.2) and using (2.5.8) we get

$$
\begin{equation*}
\nabla \times(\nabla \times \vec{\zeta})=0 \tag{2.5.9}
\end{equation*}
$$

After some straightforward algebra given in the Appendix, we can show that

$$
\begin{equation*}
\vec{q}=\nabla \times\left(\frac{\psi \vec{e}_{\phi}}{r \sin \theta}\right) \tag{2.5.10}
\end{equation*}
$$

and

$$
\begin{equation*}
\vec{\zeta}=\nabla \times \vec{q}=\nabla \times \nabla \times\left(\frac{\psi \vec{e}_{\phi}}{r \sin \theta}\right)=-\frac{\vec{e}_{\phi}}{r \sin \theta}\left(\frac{\partial^{2} \psi}{\partial r^{2}}+\frac{\sin \theta}{r^{2}} \frac{\partial}{\partial \theta}\left(\frac{1}{\sin \theta} \frac{\partial \psi}{\partial \theta}\right)\right) \tag{2.5.11}
\end{equation*}
$$

Now from (2.5.9)

$$
\nabla \times \nabla \times(\nabla \times \vec{q})=\nabla \times \nabla \times\left[\nabla \times\left(\nabla \times \frac{\psi \vec{e}_{\phi}}{r \sin \theta}\right)\right]=0
$$

hence, the momentum equation (2.5.9) becomes a scalar equation for $\psi$.

$$
\begin{equation*}
\left(\frac{\partial^{2}}{\partial r^{2}}+\frac{\sin \theta}{r^{2}} \frac{\partial}{\partial \theta}\left(\frac{1}{\sin \theta} \frac{\partial}{\partial \theta}\right)\right)^{2} \psi=0 \tag{2.5.12}
\end{equation*}
$$

The boundary conditions on the sphere are

$$
\begin{equation*}
q_{r}=0 \quad q_{\theta}=0 \quad \text { on } \quad r=a \tag{2.5.13}
\end{equation*}
$$

The boundary conditions at $\infty$ is

$$
\begin{equation*}
\psi \rightarrow \frac{W}{2} r^{2} \sin ^{2} \theta \tag{2.5.14}
\end{equation*}
$$

Let us try a solution of the form:

$$
\begin{equation*}
\psi(r, \theta)=f(r) \sin ^{2} \theta \tag{2.5.15}
\end{equation*}
$$

then $f$ is governed by the equi-dimensional differential equation:

$$
\begin{equation*}
\left[\frac{d^{2}}{d r^{2}}-\frac{2}{r^{2}}\right]^{2} f=0 \tag{2.5.16}
\end{equation*}
$$

whose solutions are of the form $f(r) \propto r^{n}$, It is easy to verify that $n=-1,1,2,4$ so that

$$
f(r)=\frac{A}{r}+B r+C r^{2}+D r^{4}
$$

or

$$
\psi=\sin ^{2} \theta\left[\frac{A}{r}+B r+C r^{2}+D r^{4}\right]
$$

To satisfy (2.5.14) we set $D=0, C=W / 2$. To satisfy (2.5.13) we use (2.5.4) to get

$$
q_{r}=0=\frac{W}{2}+\frac{A}{a^{3}}+\frac{B}{a}=0, \quad q_{\theta}=0=W-\frac{A}{a^{3}}+\frac{B}{a}=0
$$

Hence

$$
A=\frac{1}{4} W a^{3}, \quad B=-\frac{3}{4} W a
$$

Finally the stream function is

$$
\begin{equation*}
\psi=\frac{W}{2}\left[r^{2}+\frac{a^{3}}{2 r}-\frac{3 a r}{2}\right] \sin ^{2} \theta \tag{2.5.17}
\end{equation*}
$$

Inside the parentheses, the first term corresponds to the uniform flow, and the second term to the doublet; together they represent an inviscid flow past a sphere. The third term is called the Stokeslet, representing the viscous correction.

The velocity components in the fluid are: (cf. (2.5.4) :

$$
\begin{align*}
& q_{r}=W \cos \theta\left[1+\frac{a^{3}}{2 r^{3}}-\frac{3 a}{2 r}\right]  \tag{2.5.18}\\
& q_{\theta}=-W \sin \theta\left[1-\frac{a^{3}}{4 r^{3}}-\frac{3 a}{4 r}\right] \tag{2.5.19}
\end{align*}
$$

### 2.5.1 Physical Deductions

1. Streamlines: With respect to the the equator along $\theta=\pi / 2, \cos \theta$ and $q_{r}$ are odd while $\sin \theta$ and $q_{\theta}$ are even. Hence the streamlines (velocity vectors) are symmetric fore and aft.
2. Vorticity:

$$
\vec{\zeta}=\zeta_{\phi} \vec{e}_{\phi}\left(\frac{1}{r} \frac{\partial\left(r q_{\theta}\right)}{\partial r}-\frac{1}{r} \frac{\partial q_{r}}{\partial \theta}\right) \vec{e}_{\phi}=-\frac{3}{2} W a \frac{\sin \theta}{r^{2}} \vec{e}_{\phi}
$$

3. Pressure : From the $r$-component of momentum equation

$$
\frac{\partial p}{\partial r}=\frac{\mu W a}{r^{3}} \cos \theta(=-\mu \nabla \times(\nabla \times \vec{q}))
$$

Integrating with respect to $r$ from $r$ to $\infty$, we get

$$
\begin{equation*}
p=p_{\infty}-\frac{3}{2} \frac{\mu W a}{r^{3}} \cos \theta \tag{2.5.20}
\end{equation*}
$$

4. Stresses and strains:

$$
\frac{1}{2} e_{r r}=\frac{\partial q_{r}}{\partial r}=W \cos \theta\left(\frac{3 a}{2 r^{2}}-\frac{3 a^{3}}{2 r^{4}}\right)
$$

On the sphere, $r=a, e_{r r}=0$ hence $\sigma_{r r}=0$ and

$$
\begin{equation*}
\tau_{r r}=-p+\sigma_{r r}=-p_{\infty}+\frac{3}{2} \frac{\mu W}{a} \cos \theta \tag{2.5.21}
\end{equation*}
$$

On the other hand

$$
e_{r \theta}=r \frac{\partial}{\partial r}\left(\frac{q_{\theta}}{r}\right)+\frac{1}{r} \frac{\partial q_{r}}{\partial \theta}=-\frac{3}{2} \frac{W a^{3}}{r^{4}} \sin \theta
$$

Hence at $r=a$ :

$$
\begin{equation*}
\tau_{r \theta}=\sigma_{r \theta}=\mu e_{r \theta}=-\frac{3}{2} \frac{\mu W}{a} \sin \theta \tag{2.5.22}
\end{equation*}
$$

The resultant stress on the sphere is parallel to the $z$ axis.

$$
\Sigma_{z}=\tau_{r r} \cos \theta-\tau_{r \theta} \sin \theta=-p_{\infty} \cos \theta+\frac{3}{2} \frac{\mu W}{a}
$$

The constant part exerts a net drag in $z$ direction

$$
\begin{equation*}
D=\int_{o}^{2 \pi} a d \phi \int_{o}^{\pi} d \theta \sin \theta \Sigma_{z}==\frac{3}{2} \frac{\mu W}{a} 4 \pi a^{2}=6 \pi \mu W a \tag{2.5.23}
\end{equation*}
$$

This is the celebrated Stokes formula.
A drag coefficient can be defined as

$$
\begin{equation*}
C_{D}=\frac{D}{\frac{1}{2} \rho W^{2} \pi a^{2}}=\frac{6 \pi \mu W a}{\frac{1}{2} \rho W^{2} \pi a^{2}}=\frac{24}{\frac{\rho W(2 a)}{\mu}}=\frac{24}{R e_{d}} \tag{2.5.24}
\end{equation*}
$$

5. Fall velocity of a particle through a fluid. Equating the drag and the buoyant weight of the eparticle

$$
6 \pi \mu W_{o} a=\frac{4 \pi}{3} a^{3}\left(\rho_{s}-\rho_{f}\right) g
$$

hence

$$
W_{o}=\frac{2}{9} g\left(\frac{a^{2}}{\nu} \frac{\Delta \rho}{\rho_{f}}\right)=217.8\left(\frac{a^{2}}{\nu} \frac{\Delta \rho}{\rho_{f}}\right)
$$

in cgs units. For a sand grain in water,

$$
\begin{gather*}
\frac{\Delta \rho}{\rho_{f}}=\frac{2.5-1}{1}=1.5, \quad \nu=10^{-2} \mathrm{~cm}^{2} / \mathrm{s} \\
W_{o}=32,670 a^{2} \mathrm{~cm} / \mathrm{s} \tag{2.5.25}
\end{gather*}
$$

To have some quantitative ideas, let us consider two sand of two sizes :

$$
\begin{aligned}
& a=10^{-2} \mathrm{~cm}=10^{-4} \mathrm{~m}: \quad W_{o}=3.27 \mathrm{~cm} / s \\
& a=10^{-3} \mathrm{~cm}=10^{-5}=10 \mu m, \quad W_{o}=0.0327 \mathrm{~cm} / s=117 \mathrm{~cm} / \mathrm{hr}
\end{aligned}
$$

For a water droplet in air,

$$
\frac{\Delta \rho}{\rho_{f}}=\frac{1}{10^{-3}}=10^{3}, \quad \nu=0.15 \mathrm{~cm}^{2} / \mathrm{sec}
$$

then

$$
\begin{equation*}
W_{o}=\frac{(217.8) 10^{3}}{0.15} a^{2} \tag{2.5.26}
\end{equation*}
$$

in cgs units. If $a=10^{-3} \mathrm{~cm}=10 \mu \mathrm{~m}$, then $W_{o}=1.452 \mathrm{~cm} / \mathrm{sec}$.

## Details of derivation

Details of (2.5.10).

$$
\begin{gathered}
\nabla \times\left(\frac{\psi}{r \sin \theta} \vec{e}_{\phi}\right)=\frac{1}{r^{2} \sin \theta}\left|\begin{array}{ccc}
\vec{e}_{r} & \vec{e}_{\theta} & r \sin \theta \vec{e}_{\phi} \\
\frac{\partial}{\partial r} & \frac{\partial}{\partial \theta} & \frac{\partial}{\partial \phi} \\
0 & 0 & \psi
\end{array}\right| \\
=\vec{e}_{r}\left(\frac{1}{r^{2} \sin \theta} \frac{\partial \psi}{\partial \theta}\right)-\vec{e}_{\theta}\left(\frac{1}{r \sin \theta} \frac{\partial \psi}{\partial r}\right)
\end{gathered}
$$

Details of (2.5.11).

$$
\begin{aligned}
\nabla & \times \nabla \times \frac{\psi \vec{e}_{\phi}}{r \sin \theta}=\nabla \times \vec{q} \\
& =\frac{1}{r^{2} \sin \theta}\left|\begin{array}{ccc}
\vec{e}_{r} & r \vec{e}_{\theta} & r \sin \theta \vec{e}_{\phi} \\
\frac{\partial}{\partial r} & \frac{\partial}{\partial \theta} & \frac{\partial}{\partial \phi} \\
\frac{1}{r^{2} \sin \theta} \frac{\partial \psi}{\partial \theta} & \frac{-1}{\sin \theta} \frac{\partial \psi}{\partial r} & 0
\end{array}\right| \\
& =\frac{\vec{e}_{\theta}}{r \sin \theta}\left[\frac{\partial^{2} \phi}{\partial r^{2}}+\frac{\sin \theta}{r^{2}} \frac{\partial}{\partial \theta}\left(\frac{1}{\sin \theta} \frac{\partial \psi}{\partial \theta}\right)\right]
\end{aligned}
$$

