

Revisiting hadron structures in 3D

From electromagnetic to mechanical densities

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$$\mathcal{L}_{\text{QCD}} = \sum_q \bar{\psi}_q (i\gamma_\mu D^\mu - m_q) \psi_q - \frac{1}{2} \text{Tr}[G_{\mu\nu} G^{\mu\nu}]$$

Jefferson Lab Theory Seminar,
August 18, 2025, Zoom



A brief (and biased) history of the proton structure

- Proton magnetic moment (1930s)
- Elastic scattering of proton (1950s)
- Quark model (early 1960s)
- Chiral symmetry breaking (1960s)
- Deep inelastic scattering (late 1960s)
- Quantum chromodynamics (1970s)
-

Tremendous progress, but many puzzles remain. See, F. Gross and E. Klempt (eds.), *50 Years of quantum chromodynamics*, EPJC, 2023



Nobel prize 1943



Nobel prize 1951



Nobel prize 1969



Nobel prize 1990



Nobel prizes 1999 & 2004

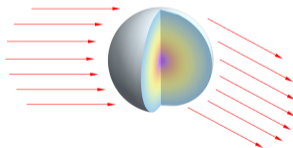


Nobel prize 2008

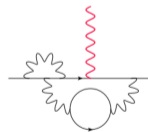
JLab, August 18, 2025

Electron structure in history

- Thomson discovered the electron in 1897 in cathode rays. Classical view of the electron: a spherical corpuscles of the size $r_e \sim 10^{-15}$ m moving in high speeds
- Minkowski (1908), and Einstein & Laub (1908) established the correct theory of macroscopic electromagnetism of moving bodies based on Einstein's electrodynamic theory of moving bodies (1905), i.e. special relativity [W. Pauli. Theory of Relativity (Oxford Univ. Press, 1958)]
 - Modern view of the electron: a pointlike particle $r_e \leq 10^{-24}$ m, first verified by Sam Ting in 1967
 - Relativity as a byproduct -- a right theory (indeed, a great theory) for a wrong problem
 - Modern applications: cosmology, black hole merger, inertial fusion, quark-gluon plasma (hydro), ...



Thomson Lorentz Abraham Einstein Minkowski Laub



Tomonaga Schwinger Feynman Ting

Minkowski-Einstein-Laub theory

$$\begin{array}{ccc} \partial_\alpha F^{\alpha\beta} = j^\beta & \xrightarrow{j^\beta = j_f^\beta + \partial_\alpha M^{\alpha\beta}} & \partial_\alpha H^{\alpha\beta} = j_f^\beta \\ \partial_\alpha \tilde{F}^{\alpha\beta} = 0 & & \partial_\alpha \tilde{F}^{\alpha\beta} = 0 \end{array}$$

where, j^α is the full current, and $j_f^\alpha = \varrho u^\alpha + j_{\text{Ohm}}^\alpha$ is the free current, and $H^{\alpha\beta} \equiv F^{\alpha\beta} - M^{\alpha\beta}$

- The theory appears identical to Maxwell's theory of macroscopic electromagnetism, but now **applicable to media in motion** -- rediscovered several times in applied physics!
- **Co-moving decomposition** of the medium polarization tensor:

$$M^{\alpha\beta} = u^\alpha \mathcal{P}^\beta - u^\beta \mathcal{P}^\alpha + \varepsilon^{\alpha\beta\kappa\lambda} u_\kappa \mathcal{M}_\lambda$$

where, u^α is the velocity vector, \mathcal{P}^α and \mathcal{M}^α are co-moving polarization & magnetization vectors

- Co-moving polarization & effective magnetic charge densities: $\varrho_{\text{pol}} = -\partial_\alpha \mathcal{P}^\alpha$, $\varrho_{\text{mag}} = -\partial_\alpha \mathcal{M}^\alpha$
- Dependence on the choice of u^α , e.g. Landau-Lifshitz vs. Eckart frames [Eckart:1940te]
- Need the microscopic theory to obtain the **full current** as developed by Max Born in 1909

[H. Minkowski Nachr. Ges. Wiss. Gött., Math.-Phys. Kl. 53 (1908);]

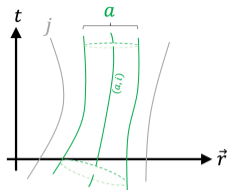
[H. Minkowski, Math. Ann. 68 (1910) 472;]

[A. Einstein, J. Laub, Annals Phys. 331 (1908) 532]

Classical many-body theory of Born

$$j^\mu(x) = \sum_{a,i} e_{ai} \int d\tau_{ai} \dot{X}_{ai}^\mu(\tau_{ai}) \delta^4(X_{ai}(\tau_{ai}) - x)$$

where a and i enumerate the "atoms" and their constituents, respectively.



- Born introduced a privileged worldline $X_a(\tau_a)$ for atom a and define $r_{ai} = X_{ai} - X_a$:

$$j^\mu(x) = \sum_{n=0}^{\infty} \sum_{a \in A} \int \tau_a \sum_{i \in a} e_{ai} \left[\dot{X}_a^\mu(\tau_a) + \dot{r}_{ai}^\mu(\tau_a) \right] (r_{ai} \cdot \partial)^n \delta^4(X_a(\tau_a) - x).$$

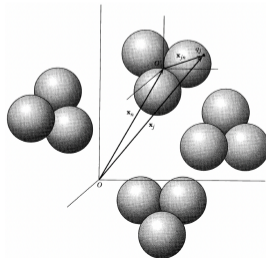
- **Multipole structure**, and the monopole term ($n = 0$) defines the free current

$$j_f^\mu(x) = \sum_a \int d\tau_a e_a \dot{X}_a^\mu(\tau_a) \delta^4(X_a(\tau_a) - x)$$

- If atomic motion \dot{X}_a negligible, factorization of the atomic distribution and **the internal density** within each atom

$$j^\mu(x) = \int d^3R \mathcal{J}^\mu(\vec{x} - \vec{R}) \rho_a(\vec{R}, t) + O(\dot{X}_a)$$

- Weyl quantization, Wigner-Newton position operator & particle localization



Quantum many-body theory and nuclear structure

- One-body density (OBD):

$$\rho(\vec{r}) = \langle \Psi | \sum_i e_i \delta^3(\vec{r} - \vec{r}_i) | \Psi \rangle$$

- Factorization of c.m. motion $|\Psi\rangle = |\Phi_{\text{cm}}\rangle |\psi_{\text{rel}}\rangle$ and translationally-invariant OBD:

$$\rho(\vec{r}) = \int d^3R \rho_{\text{cm}}(\vec{R}) \varrho(\vec{r} - \vec{R})$$

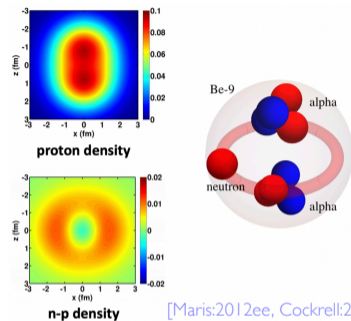
- Form factor: F.T. of OBD $\varrho(r)$

$$F(q^2) = \int d^3r e^{i\vec{q}\cdot\vec{r}} \varrho(\vec{r})$$

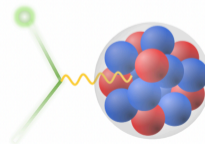
- Root-mean-square (r.m.s.) radius:

$$r^2 \equiv \int d^3r r^2 \varrho(\vec{r}) = -6F'(0)$$

- Elastic eA scattering $\sigma_{\text{el.}} = \sigma_{\text{Mott}} |F(\vec{q}^2)|^2$



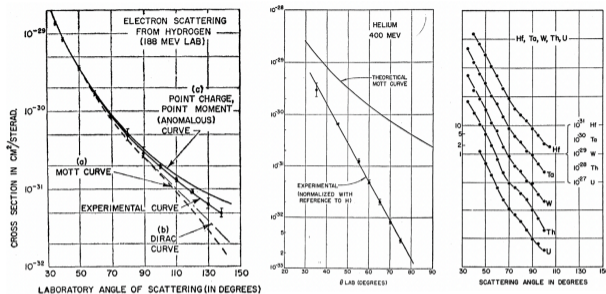
[Maris:2012ee, Cockrell:2012hf]



Nucleus is at rest (no recoil)



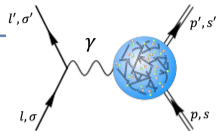
Hofstadter



	r	M	rM
pion	0.67 fm	0.14 GeV	0.5
charmonium	0.25--0.4 fm	3.0 GeV	3.8--6
proton	0.87 fm	0.94 GeV	4
nuclei	$1.3 A^{\frac{1}{3}}$ fm	$0.94A$ GeV	$6A^{\frac{4}{3}}$

- Hofstadter et al. systematically investigated nuclear structure using electron scattering
- For the proton, the nucleus of hydrogen, nucleus recoil is non-negligible: $E_\gamma \gtrsim M_p c^2$ and full relativistic description is needed.
 - Protons are intrinsically relativistic, since their radius $r_p \sim \lambda_C = M_p^{-1}$
 - In order to resolve the proton, the photon wavelength $\lambda_\gamma \ll r \Rightarrow E_\gamma \gg M_p$
 - Similar to other hadrons, e.g. neutron, pion

Proton form factors

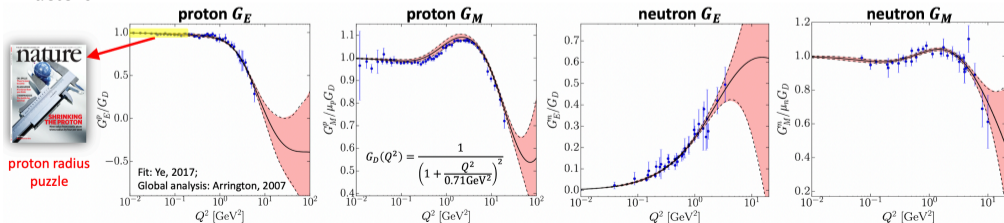


$$\frac{d\sigma_{\text{el.}}}{d\Omega} = \frac{d\sigma}{d\Omega}\bigg|_{\text{Mott}} \left[\frac{G_E^2(Q^2) + \tau G_M^2(Q^2)}{1 + \tau} + 2\tau \tan^2 \frac{\theta}{2} G_M^2(Q^2) \right]$$

where, $\tau = Q^2/4M^2$, and G_E and G_M parametrize the Lorentz covariant structures of the current hadronic matrix elements,

$$\begin{aligned} \langle p', s' | J^\mu(0) | p, s \rangle &= \bar{u}_{s'}(p') \left[\frac{P^\mu}{M} G_E(q^2) + \frac{i\varepsilon^{\mu\nu\rho\sigma} q_\nu P_\rho \gamma_\sigma \gamma_5}{2M^2} G_M(q^2) \right] u_s(p) \\ &= \bar{u}_{s'}(p') \left[\gamma^\mu F_1(q^2) + \frac{i\sigma^{\mu\nu} q_\nu}{2M} F_2(q^2) \right] u_s(p) \end{aligned}$$

Here, $P = \frac{1}{2}(p + p')$, $q = p' - p$. $G_{E,M}$ are called the Sachs form factors. $F_{1,2}$ are called the Dirac and Pauli form factors



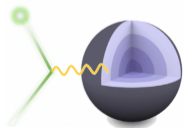
proton radius puzzle

The Sachs densities are defined as the F.T. of the hadronic matrix elements within the Breit frame

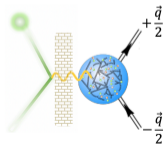
$$\rho_E^{(\text{Sach})}(\vec{r}) = \int \frac{d^3q}{(2\pi)^3 2E_q} e^{-i\vec{q}\cdot\vec{r}} \langle +\frac{1}{2}\vec{q} | J^0(0) | -\frac{1}{2}\vec{q} \rangle = \int \frac{d^3q}{(2\pi)^3} G_E(-\vec{q}^2) e^{-i\vec{q}\cdot\vec{r}}$$

- Ambiguities G_E vs F_1 vs $G_E/\sqrt{1+\tau}$ [Lorce:2020onh]
- Frame dependence: the proton is not at rest in the Breit frame. Densities in other frames?
- Lack of local probabilistic interpretation $J^0(x) = \bar{\Psi}\gamma^0\Psi \neq \sum_i e_i \delta^3(x - X_i)$ [Miller:2018ybm]
- Underlying assumption: proton as a rigid ball -- in contradiction with relativity [Jaffe:2020ebz]

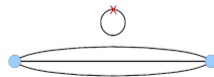
Is it possible to generalize the non-relativistic quantum many-body OBD to relativistic quantum theory?



$$\lambda_\gamma \sim r_{\text{nuc}} \gg \lambda_{\text{Comp}} = M_{\text{nuc}}^{-1}$$



$$\lambda_\gamma \sim r_N \sim \lambda_{\text{Comp}} = M_N^{-1}$$



Relativistic quantum many-body theory on the light front

- In relativistic quantum theory, the **position operator** does not exist*
- Fortunately, in light-front dynamics, transverse position operator exist, which defines the **OBD on the 2D transverse plane**:

$$\rho_T(\vec{r}_\perp) = \langle \Psi | \sum_i e_i \delta^2(r_\perp - r_{i\perp}) | \Psi \rangle$$

where, $\Sigma^+(\vec{r}_\perp) \equiv \sum_i e_i \delta^2(r_\perp - r_{i\perp})$ is the transverse charge density operator.

- Factorization of c.m. motion $|\Psi\rangle = |\Phi_{\text{cm}}\rangle |\psi_{\text{rel}}\rangle$ thanks to Galilean covariance in LFD
- Indeed, F.T. of $\rho_T(\vec{r}_\perp)$ is given by the celebrated **Drell-Yan-West formula** [Drell:1969km, West:1970av]

$$\rho_T(\vec{r}_\perp) \xrightarrow{\text{F.T.}} F(\vec{q}_\perp) = \sum_n \int [dx_i d^2k_{i\perp}]_n \sum_j e_j \psi_n^* (\{x_i, \vec{k}'_{i\perp}\}) \psi_n (\{x_i, \vec{k}_{i\perp}\})$$

where, Drell-Yan frame $q^+ = 0$ is equivalent to the longitudinal compactification $\Sigma^+ = (1/2) \int dx^- J^+(x)$

light-cone coordinates: $x^\pm = x^0 \pm x^3$,
 $\vec{x}_\perp = (x^1, x^2)$

$$\begin{aligned}\rho_T(\vec{r}_\perp) &= \int \frac{d^2 q_\perp}{(2\pi)^2 2P^+} e^{-i\vec{q}_\perp \cdot \vec{r}_\perp} \langle P + \frac{1}{2}q | J^+(0) | P - \frac{1}{2}q \rangle \Big|_{q^+=0} \\ &= \int \frac{d^2 q_\perp}{(2\pi)^2} e^{-i\vec{q}_\perp \cdot \vec{r}_\perp} F_1(\vec{q}_\perp^2)\end{aligned}$$

- Frame independent: boost invariance in light-front dynamics
- Local probabilistic interpretation: $J^+ \sim \sum_i \bar{q}_i \gamma^+ q_i \sim \sum_i e_i N_i$, vacuum suppressed
- Intrinsically relativistic and related to the forward generalized parton density $q(x, \vec{b}_\perp)$, i.e. what the probes "see" in high-energy collision experiments

[Burkardt:2000za]

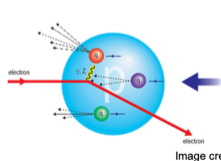
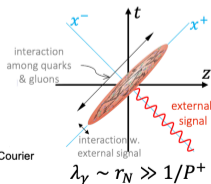
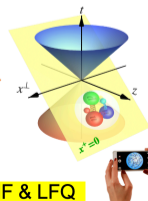


Image credit: CERN Courier

Experiment: DIS



Theory: IMF & LFQ



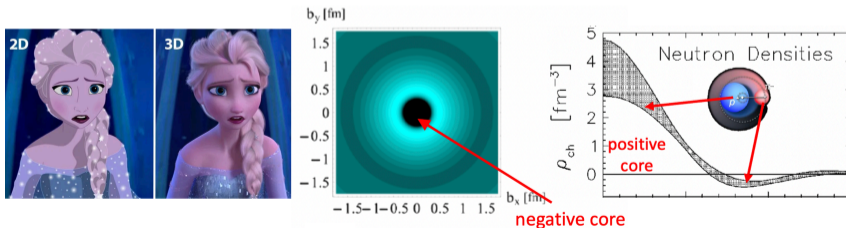
light-cone coordinates:

$$x^\pm = x^0 \pm x^3,$$

$$\vec{x}_\perp = (x^1, x^2)$$

What does the proton look like in 3D?

- Light-front densities are 2D $\xrightarrow{?}$ 3D [Panteleeva:2021iip]
- Light-front densities can be understood as equal-time densities in the infinite momentum frame which could be counter-intuitive [Lorce:2020onh]
- Physical densities associated with "bad" components, e.g. J^- , are not well explained
- Amplitude vs. quantum expectation value: what is truly probed by a semiclassical electromagnetic field is the quantum expectation value $j^\alpha(x) = \langle \Psi | J^\alpha(x) | \Psi \rangle$ where $|\Psi\rangle$ is a generic hadronic state

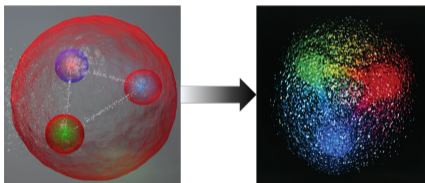


What does the proton look like in 3D?

resolving a non-relativistic particle: $r_{\text{hadron}} \gg \lambda_{\gamma} \gg \lambda_{\text{hadron}} \geq \lambda_C$

resolving a relativistic hadron: $\lambda_{\text{hadron}} \gtrsim r_{\text{hadron}} \sim \lambda_C \gg \lambda_{\gamma}$

where $\lambda_C = M^{-1}$ is the Compton wavelength, $\lambda_{\gamma} = Q^{-1}$ is the photon wavelength. λ_{hadron} is the de Broglie wavelength. r_{hadron} is the hadron radius.



- $\lambda_{\gamma} \ll \lambda_{\text{hadron}}$: the photon "sees" a de Broglie wave (medium)!
- The relevant theory is Minkowski-Einstein-Laub's *relativistic theory of macroscopic electromagnetism*
- Mass & spin decomposition as multi-fluid description

[de Broglie, 1924]

[Lorce:2017xzd]

Relativistic theory of macroscopic electromagnetism

$$j^\beta = j_f^\beta + \partial_\alpha M^{\alpha\beta} \Rightarrow \partial_\alpha H^{\alpha\beta} = j_f^\beta, \quad (H^{\alpha\beta} \equiv F^{\alpha\beta} - M^{\alpha\beta})$$

where, $j_f^\alpha = \varrho_f u^\alpha + j_{\text{Ohm}}^\alpha$ is the free current, consisting of the convective current and the Ohmic current. and $M^{\alpha\beta} = u^\alpha \mathcal{P}^\beta - u^\beta \mathcal{P}^\alpha + \varepsilon^{\alpha\beta\rho\sigma} u_\rho \mathcal{M}_\sigma$ is the medium polarization tensor.

- Our "medium" is non-dissipative, unmagnetized, unpolarized ($\alpha_s \gg \alpha_{\text{QED}}$)
- Densities (frame dependent): charge $\rho = j^0$, bound charge density $\rho_b = \nabla \cdot \vec{P}$, effective magnetic charge density $\rho_m = \nabla \cdot \vec{M}$, magnetization current $\vec{j}_m = \nabla \times \vec{M}$
- Co-moving densities -- densities measured in local rest frame, are Lorentz invariants:
 - Free charge density: $\varrho_f = u_\alpha j_f^\alpha$
 - Bound charge density: $\varrho_b = -\partial_\alpha \mathcal{P}^\alpha$
 - Total charge density: $\varrho = \varrho_f + \varrho_b$
 - Effective magnetic charge density: $\varrho_m = -\partial_\alpha \mathcal{M}^\alpha$

[Jackson, Classical Electrodynamics (John Wiley & Sons, 1999)]

[W. Pauli. Theory of Relativity (Oxford Univ. Press, 1958)]

[S.R. de Groot & L.G. Suttorp, Foundations of electrodynamics (North-Holland, 1972)]

Factorization

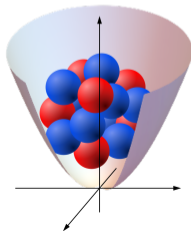
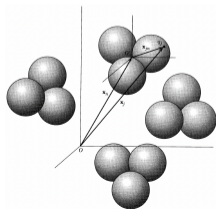
Consider media comprising of composite particles, e.g. atoms, molecules. The total current density factorizes into a convolution of an internal density and a convective current density:

$$j^\alpha(x) = \int d^3r \mathcal{J}^\alpha(\vec{r}) \rho_f(\vec{x} - \vec{r}, t)$$

Here, $\mathcal{J}^\alpha(\vec{x})$ is known as the atomic/molecular density.

Similar factorization formula for nuclear densities,

$$\rho(\vec{r}) = \int d^3R \rho_{\text{cm}}(\vec{R}) \rho_{\text{int}}(\vec{r} - \vec{R})$$



$$\begin{aligned}
 j^\mu(x) &\equiv \langle \Psi | J^\mu(x) | \Psi \rangle \\
 &= \frac{1}{M} \int \frac{d^3p}{(2\pi)^3 2p^0} \int \frac{d^3p'}{(2\pi)^3 2p'^0} \tilde{\Psi}_{s'}^*(p') \tilde{\Psi}_s(p) e^{iq \cdot x} \\
 &\quad \times \bar{u}_{s'}(p') \left[P^\mu G_E(q^2) + \frac{i}{2M} \varepsilon^{\mu\nu\rho\sigma} q_\nu P_\rho \gamma_\sigma \gamma_5 G_M(q^2) \right] u_s(p)
 \end{aligned}$$

where, $P = (p' + p)/2$ and $q = p' - p$. $\Psi_s(\vec{p}) = \langle p, s | \Psi \rangle$ is the momentum-space wavepacket, which is arbitrary, e.g. Gaussian $\tilde{\Psi}_s(\vec{p}) \propto \exp\left\{\frac{(\vec{p}-\vec{p}_0)^2}{4\sigma^2}\right\}$ and stationary planewave $\tilde{\Psi}_s(\vec{p}) \propto \delta^3(\vec{p})$ [Ernst:1960zza]

- Introduce the "coordinate space wave function" as the F.T. of $\tilde{\Psi}_s(\vec{p})$,

$$\Psi(x) = \sum_s \int \frac{d^3p}{(2\pi)^3 2p^0} \tilde{\Psi}_s(\vec{p}) u_s(p) e^{ip \cdot x}.$$

where $\Psi(x)$ satisfies the Dirac equation

- Conserved number current: $(f \vec{\partial} g \equiv f \partial g - \partial f g)$

$$n^\mu(x) \equiv \frac{1}{2M} \bar{\Psi}(x) i \vec{\partial}^\mu \Psi(x) \quad \Rightarrow \quad \partial_\mu n^\mu = 0$$

$$n^\alpha = n u^\alpha, \quad (u_\alpha u^\alpha = 1).$$

$n = n_\alpha u^\alpha$ is the proper number density

- The medium (wave) velocity u :

$$u^\alpha(x) \equiv \bar{\Psi}(x) \mathcal{U}^\alpha \Psi(x) = \frac{1}{\sqrt{4M^2 + \partial^2}} [\bar{\Psi} i \vec{\partial}^\alpha \Psi] = n^\alpha - \frac{1}{8M^2} \partial^2 n^\alpha + \frac{3}{16M^4} \partial^4 n^\alpha - \dots$$

- Consider the gradient of the velocity vector:

$$\partial_\alpha u_\beta = u_\alpha a_\beta + \Omega_{\alpha\beta} + \Sigma_{\alpha\beta} + \frac{1}{3} \theta \Delta_{\alpha\beta},$$

a is the 4-acceleration, $a_\alpha u^\alpha = 0$; $\Omega_{\alpha\beta}$ is the vorticity tensor; $\Sigma_{\alpha\beta}$ is the shear tensor; θ is the expansion scalar, $\theta = 0$ for hadrons, and $\Delta^{\alpha\beta} = g^{\alpha\beta} - u^\alpha u^\beta$ is the spatial metric tensor (spatial projector)



Electromagnetic structure: spin-0

Recall, in Minkowski-Einstein-Laub's theory,

$$j^\alpha = \rho_f u^\alpha + \partial_\beta M^{\alpha\beta}$$

- Free current of spin-0 particles:

$$j_f^\alpha = \mathcal{Q} n^\alpha = \mathcal{Q} \Phi^* i \vec{\partial}^\alpha \Phi \quad \Rightarrow \quad \rho_f = \mathcal{Q} n = \mathcal{Q} \Phi^* i \vec{\partial}_\alpha \mathcal{U}^\alpha \Phi$$

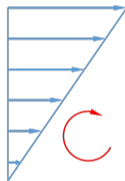
here, $\mathcal{Q} = F(0)$ is the charge number

- The polarization tensor can be constructed as,

$$\begin{aligned} M^{\alpha\beta}(x) &= \left\langle \frac{1}{M} \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot z} q^{[\alpha} i \vec{\partial}^{\beta]} \frac{F(q^2) - F(0)}{iq^2} \right\rangle \\ \Rightarrow \rho_b(x) &= \left\langle \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot z} \sqrt{1 - \frac{q^2}{4M^2}} [F(q^2) - F(0)] \right\rangle, \\ \rho_m(x) &= 0. \end{aligned}$$

where, the quantum average is defined as,

$$\langle \mathcal{O}(x) \rangle \equiv 2M \int d^3z \Phi^*(x-z) \mathcal{O}(z) \Phi(x-z)$$



Electromagnetic structure: spin-0

- Full electric density in the local rest frame (LRF):

$$\begin{aligned}\rho &= \rho_f + \rho_b = \langle \varrho(z) \rangle \\ &= 2M \int d^3z \Phi^*(x-z) \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot z} \sqrt{1 - \frac{q^2}{4M^2}} F(q^2) \Phi(x-z)\end{aligned}$$

- Full density is a convolution of the wave density $\propto \Phi^* \Phi$, i.e. the c.m. density and the hadronic density $\varrho(z)$, i.e. the intrinsic distribution
- The full current can also be written as,

$$j^\alpha = \langle \varrho \mathcal{U}^\alpha \rangle$$

corresponds to a new current decomposition in relativistic hydrodynamics (for spin-0):

$$j^\alpha = \rho u^\alpha$$

Electromagnetic structure: spin-1/2

Generalizing the current decomposition to polarized relativistic matter:

- Electromagnetic decomposition:

$$j^\alpha = \rho u^\alpha + \partial_\beta M^{\alpha\beta}$$

- Special case: Minkowski-Einstein-Laub decomposition:

$$j^\alpha = \rho_f u^\alpha + \partial_\beta M^{\alpha\beta}$$

- Bemfica-Disconzi-Noronha-Kovtun (BDNK) formalism: [\[Kovtun:2016lfw, Hernandez:2017mch, Bemfica:2019knx\]](#)

$$j^\alpha = \mathcal{N} u^\alpha + \mathcal{J}^\alpha \quad (\mathcal{J}^\alpha \equiv \Delta^\alpha_\beta j^\beta)$$

where the relation between the e.m. and the 3+1 decompositions is,

$$\mathcal{N} = \rho + \partial_\alpha \mathcal{P}^\alpha + \mathcal{P}_\alpha a^\alpha + \mathcal{M}_\alpha \Omega^\alpha, \quad \mathcal{J}^\alpha = u_\beta \dot{\mathcal{P}}^{[\beta} u^{\alpha]} + \varepsilon^{\alpha\beta\rho\sigma} u_\rho \partial_\beta \mathcal{M}_\sigma$$

where, $\Omega^\alpha = \frac{1}{2} \varepsilon^{\alpha\beta\rho\sigma} u_\beta \Omega_{\rho\sigma}$ is the vorticity vector, and $\dot{a} = u_\alpha \partial^\alpha a$

No dissipation (no entropy production), no gradient expansion, no local thermal equilibrium assumption

Electromagnetic decomposition

Hadron matrix elements expressed using Dirac-Pauli form factors:

$$(F_{1+2} = F_1 + F_2)$$

$$\begin{aligned}\langle p', s' | J^\mu(0) | p, s \rangle &= \bar{u}_{s'}(p') \left[\frac{P^\mu}{M} F_1(q^2) + \frac{i\sigma^{\mu\nu} q_\nu}{2M} F_{1+2}(q^2) \right] u_s(p), \\ \Rightarrow j^\alpha &= \left\langle \varrho \mathcal{U}^\alpha + \frac{1}{2M} \sigma^{\alpha\beta} \partial_\beta \mathcal{F}_{1+2} \right\rangle\end{aligned}$$

where, the quantum average for spin-1/2 particles is,

$$\langle \mathcal{O}(x) \rangle = \int d^3z \bar{\Psi}(x-z) \mathcal{O}(z) \Psi(x-z)$$

- One can identify the F.T. of the Dirac form factor as the hadronic electric charge density,

$$\varrho(z) = \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot z} \sqrt{1 - \frac{q^2}{4M^2}} F_1(q^2)$$

- Identify the F.T. of the F_{1+2} as the polarization (magnetic dipole) density,

$$\mathcal{M}^{\alpha\beta} = \frac{1}{2M} \sigma^{\alpha\beta} \mathcal{F}_{1+2}, \quad \mathcal{F}_{1+2}(z) = \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot z} F_{1+2}(q^2)$$



3+1 decomposition

Hadron matrix elements expressed using Sachs form factors:

$$\begin{aligned}\langle p', s' | J^\mu(0) | p, s \rangle &= \bar{u}_{s'}(p') \left[\frac{P^\mu}{M} G_E(q^2) + \frac{1}{4M^2} i \varepsilon^{\mu\nu\rho\sigma} q_\nu P_\rho \gamma_\sigma \gamma_5 G_M(q^2) \right] u_s(p), \\ \Rightarrow j^\alpha &= \left\langle \mathcal{N} \mathcal{U}^\alpha + \frac{1}{2M} \varepsilon^{\alpha\beta\rho\sigma} \mathcal{U}_\rho \gamma_\sigma \gamma_5 \partial_\beta \mathcal{G}_M \right\rangle\end{aligned}$$

- One can identify the F.T. of the Sachs form factor G_E as the hadronic temporal charge density,

$$\mathcal{N}(z) = \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot z} \sqrt{1 - \frac{q^2}{4M^2}} G_E(q^2)$$

- Identify the F.T. of the Sachs form factor G_M as the magnetization density,

$$\mathcal{M}^\alpha = \frac{\gamma^\alpha \gamma_5}{2M} \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot z} \sqrt{1 - \frac{q^2}{4M^2}} G_M(q^2) \propto \mathcal{J}_A^\alpha$$



$$\rho(x) = \int d^3z \bar{\Psi}(x-z) \left\{ \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot z} \sqrt{1 - \frac{q^2}{4M^2}} F_{\text{ch}}(q^2) \right\} \Psi(x-z) \Big|_{z^0=0}$$

The **hadronic part** is *not factorizable* due to the dependence of $\vec{P} = (-i/2)\vec{\nabla}_x$ in $q^2 = (q^0)^2 - \vec{q}^2$, where $q^0 = \sqrt{(\vec{P} + \frac{1}{2}\vec{q})^2 + M^2} - \sqrt{(\vec{P} - \frac{1}{2}\vec{q})^2 + M^2}$

- Taylor expansion around $\vec{P} = 0$: multipole series,

$$\rho(\vec{x}) = \sum_{n=0}^{\infty} \frac{(-i)^n}{2^n n!} \rho_n^{i_1 i_2 \dots i_n}(\vec{x}) \vec{\nabla}^{i_1} \vec{\nabla}^{i_2} \dots \vec{\nabla}^{i_n}$$

- Monopole density gives the **Breit-frame distribution (Sachs distribution)**

$$\rho(r) = \int \frac{d^3q}{(2\pi)^3} \sqrt{1 + \frac{\vec{q}^2}{4M^2}} F_{\text{ch}}(-\vec{q}^2) e^{-i\vec{q} \cdot \vec{r}} + O(\max\{r_{\text{hadron}}, \lambda_C\}/\lambda_{\text{hadron}})$$

- High-multipole moments exist due to Lorentz distortion
- No special frame or non-relativistic approximation is taken
- Convergence of the multipole series: $\lambda_{\text{hadron}} \gg \max\{r_{\text{hadron}}, \lambda_C\}$ -- sufficiently *delocalized* wavepacket

Is the multipole expansion unique? No! \rightarrow Alternative: Taylor (Laurent) expansion around $1/|\vec{P}| = 0$

- Sufficient to take $P_z \rightarrow \infty \Rightarrow |\vec{P}| = \sqrt{\vec{P}_\perp^2 + P_z^2} \rightarrow \infty$
- Monopole density gives the **2D light-front distribution**

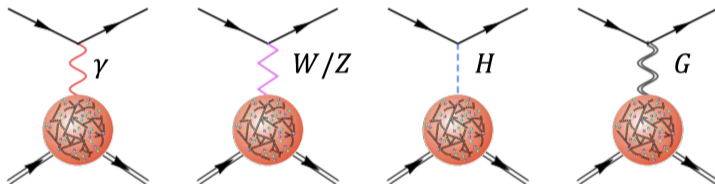
$$\rho_{\text{Mon}}(r) = \delta(r_\parallel) \int \frac{d^2 q_\perp}{(2\pi)^3} \sqrt{1 + \frac{\vec{q}_\perp^2}{4M^2}} F_{\text{ch}}(-\vec{q}_\perp^2) e^{-i\vec{q}_\perp \cdot \vec{r}_\perp}$$

- No special frame, e.g. Drell-Yan $q^+ = 0$ frame, is chosen
- Relativistic, suitable also for massless hadrons (in contrast to the Sachs distribution)
- Convergence of the multipole series: $|\vec{P}| \gg \lambda_{\text{hadron}}^{-1} \gg \{M, r_{\text{hadron}}^{-1}\}$ -- sufficiently *localized* z -direction
- Consistent with analyses from the (2+2)D phase space approach and the light-front quantization approach. And we have retained full Lorentz covariance [Lorce:2020onh, Freese:2021mzg]
- In the infinite momentum frame (IMF), the current components form a hierarchy: $j^+ \gg \vec{j}_\perp \gg j^-$

Hadronic energy-momentum tensor

$$H = \int d^3x T^{00}(x) \quad \Rightarrow \quad t^{\alpha\beta}(x) = \langle \Psi | T^{\alpha\beta}(x) | \Psi \rangle$$

Hadronic energy-momentum tensor encodes the energy-stress densities inside hadrons



Hadronic matrix elements and gravitational form factors (GFFs):

[Kobzarev:1962wt, Pagels:1966zza]

$$\langle p', s' | T_i^{\mu\nu}(0) | p, s \rangle = \frac{1}{M} \bar{u}_{s'}(p') \left[P^\mu P^\nu A_i(q^2) + \frac{1}{2} i P^{\{\mu} \sigma^{\nu\} \rho} q_\rho J_i(q^2) + \frac{1}{4} (q^\mu q^\nu - g^{\mu\nu} q^2) D_i(q^2) + g^{\mu\nu} \bar{c}_i(q^2) \right] u_s(p)$$

$$\mathcal{T}_{ss'}^{\alpha\beta}(\vec{r}_\perp; P) = \int \frac{d^2 q_\perp}{(2\pi)^2 2P^+} e^{-i\vec{q}_\perp \cdot \vec{r}_\perp} \langle P + \frac{1}{2}q, s' | T^{\alpha\beta}(0) | P - \frac{1}{2}q, s \rangle \Big|_{q^+=0}$$

$$\mathcal{P}^+(r_\perp) \equiv \mathcal{T}^{++}(r_\perp; P) = P^+ \mathcal{A}(r_\perp),$$

$$\int d^3x T^{+\mu}(x) = P^\mu \quad \Rightarrow \quad \mathcal{P}_\perp^i(r_\perp) \equiv \mathcal{T}_{ss}^{+i}(r_\perp; P) = P_\perp^i \mathcal{A}(r_\perp) + (\nabla \times \vec{\mathcal{S}})^i, \quad (i = 1, 2),$$

$$\mathcal{P}^-(r_\perp) \equiv \mathcal{T}_{ss}^{+-}(r_\perp; P) = \frac{P_\perp^2 \mathcal{A}(r_\perp) + \vec{P}_\perp \cdot (\nabla \times \vec{\mathcal{S}}) + \mathcal{M}^2(r_\perp)}{P^+}$$

- $\mathcal{A}(r_\perp)$ can be interpreted as the number density (cf. mass density, tensor density, long. momentum density)
- $\mathcal{M}^2(r_\perp)$ can be interpreted as the invariant mass squared density
- $\vec{\mathcal{S}}(r_\perp)$ can be interpreted as the spin current density

$$P^- = \frac{P_\perp^2 + M^2}{P^+}$$

$$\mathcal{A}(r_\perp) = \int \frac{d^2 q_\perp}{(2\pi)^2} e^{-i\vec{q}_\perp \cdot \vec{r}_\perp} A(q_\perp^2),$$

$$\mathcal{M}^2(r_\perp) = \int \frac{d^2 q_\perp}{(2\pi)^2} e^{-i\vec{q}_\perp \cdot \vec{r}_\perp} \left[(M^2 + \frac{1}{4}q_\perp^2) A(q_\perp^2) + \frac{1}{2}q_\perp^2 D(q_\perp) \right],$$

$$\vec{\mathcal{S}}(r_\perp) = 2\vec{s} \int \frac{d^2 q_\perp}{(2\pi)^2} e^{-i\vec{q}_\perp \cdot \vec{r}_\perp} J(q_\perp^2)$$

Frame independent!

Relativistic quantum many-body rep'n on the light front

- Drell-Yan-West formula: transverse charge density as exact OBD:

[Drell:1969km, West:1970av]

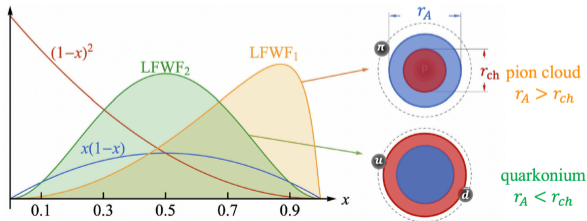
$$\rho_{\text{ch}}(r_{\perp}) = \sum_n \int [dx_i d^2 r_{i\perp}]_n \left| \tilde{\psi}_n(\{x_i, \vec{r}_{i\perp}\}) \right|^2 \sum_j e_j \delta^2(r_{\perp} - r_{j\perp}) \equiv \left\langle \sum_j e_j \delta^2(r_{\perp} - r_{j\perp}) \right\rangle$$

- Brodsky-Hwang-Ma-Schmidt formula: number density $\mathcal{A}(r_{\perp})$ as exact OBD :

[Brodsky:2000ii]

$$\mathcal{A}(r_{\perp}) = \left\langle \sum_j x_j \delta^2(r_{\perp} - r_{j\perp}) \right\rangle$$

Number density $\mathcal{A}(r_{\perp})$ mainly samples the valence partons $x_j \sim O(1)$; wee parton $x_j \ll 1$ contributions suppressed



Wave function representation of D -term

International Journal of Modern Physics A | Vol. 33, No. 26, 1830025 (2018)

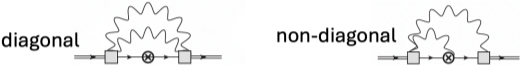
| Reviews

Forces inside hadrons: Pressure, surface tension, mechanical radius, and all that

Maxim V. Polyakov and Peter Schweitzer

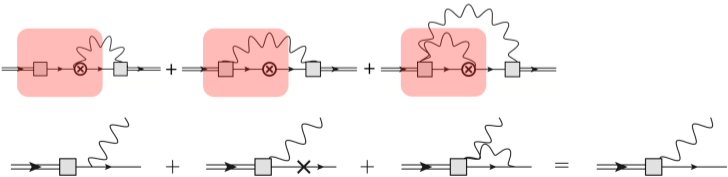
<https://doi.org/10.1142/S0217751X18300259> | Cited by: 212 (Source: Crossref)

\hat{T}_{++} of the EMT. Being related to the stress tensor \hat{T}_{ij} the form factor $D(t)$ naturally “mixes” good and bad light-front components and is described in terms of transitions between different Fock state components in overlap representation. As a quantity intrinsically nondiagonal in a Fock space, it is difficult to study the D -term in approaches based on light-front wave functions. This is due to the rela-



- In an explicit model calculation (scalar model in 3+1D), we showed that all non-diagonal contributions add up to a diagonal contribution

[Cao:2023ohj]



$$t^{12} = \frac{1}{2} \left\langle \sum_j e^{-i\vec{q}_\perp \cdot \vec{r}_{j\perp}} \frac{i\vec{\nabla}_{j\perp}^1 i\vec{\nabla}_{j\perp}^2 - q_\perp^1 q_\perp^2}{x_j} \right\rangle,$$

$$t^{+-} = 2 \left\langle \underbrace{\sum_j e^{i\vec{r}_{j\perp} \cdot \vec{q}_\perp} \frac{-\frac{1}{4}\vec{\nabla}_{j\perp}^2 + m_j^2 - \frac{1}{4}q_\perp^2}{x_j}}_{\text{kinetic part}} + \underbrace{V e^{i\vec{r}_{N\perp} \cdot \vec{q}_\perp}}_{\text{potential part}} \right\rangle$$

where, $V = M^2 - \sum_j \frac{-\vec{\nabla}_{j\perp}^2 + m_j^2}{x_j}$, and the quantum average is defined as,

$$\langle O \rangle \equiv \sum_n \int [dx_i d^2 r_{i\perp}]_n \tilde{\psi}_n^* (\{x_i, \vec{r}_{i\perp}\}) O_n \tilde{\psi}_n (\{x_i, \vec{r}_{i\perp}\})$$

- The off-shell factors $e^{i\vec{r}_{j\perp} \cdot \vec{q}_\perp} \xrightarrow{\text{F.T.}} \delta^2(r_\perp - r_{j\perp})$ indicate the location of the graviton coupling
- Large contributions from the wee parton -- condensates?
- Use scalar model as an example

$$t^{\alpha\beta} = e u^\alpha u^\beta - p \Delta^{\alpha\beta} + \frac{1}{2} \partial_\sigma (u^{\{\alpha} s^{\beta\}\sigma}) + \pi^{\alpha\beta} - g^{\alpha\beta} \Lambda + \text{dissipative terms}$$

where, u^α is the medium velocity with $u_\alpha u^\alpha = 1$, $\Delta^{\alpha\beta} = g^{\alpha\beta} - u^\alpha u^\beta$ is the spatial metric tensor.
 $a^{\{\mu} b^{\nu\}} = a^\mu b^\nu + a^\nu b^\mu$.

- $e(x)$ -- proper energy density, i.e. energy density measured in local rest frame (LRF)
- $c^{\alpha\beta} = \pi^{\alpha\beta} - p \Delta^{\alpha\beta}$ -- Cauchy stress tensor, consisting of a traceless shear tensor and a normal pressure $p(x)$.
- $\pi^{\alpha\beta}(x)$ -- shear tensor, dissipative in fluids but non-dissipative in solids
- $s^{\alpha\beta}(x)$ -- spin tensor, recently proposed by Fukushima et. al. in relativistic spin hydrodynamics
- Λ -- cosmological constant term, non-conserving, representing an external pressure [Teryaev:2013qba, Teryaev:2016edw, Liu:2023cse]
- Coupled multifluids vs Interacting multifluids

[Fukushima:2020ucl, cf. Li:2020eon]

$$T_{\text{LRF}}^{\mu\nu} = \begin{pmatrix} \boxed{T^{00}} & T^{01} & T^{02} & T^{03} \\ T^{10} & \boxed{T^{11}} & T^{12} & T^{13} \\ T^{20} & T^{21} & \boxed{T^{22}} & T^{23} \\ T^{30} & T^{31} & T^{32} & \boxed{T^{33}} \end{pmatrix}$$

energy density energy flux
momentum density momentum flux isotropic pressure

- It can be shown that the quantum expectation value of the EMT tensor can be written as,

$$\langle \Psi | T^{\alpha\beta}(x) | \Psi \rangle = \langle \mathcal{E} \mathcal{U}^\alpha \mathcal{U}^\beta - \mathcal{P} \Delta^{\alpha\beta} + \frac{1}{2} \partial_\rho (\mathcal{U}^{\{\alpha} \mathcal{S}^{\beta\} \rho}) + \Pi^{\alpha\beta} - g^{\alpha\beta} \Lambda \rangle_\Psi$$

where,

$$\langle \mathcal{O}(x) \rangle_\Psi = \int d^3z \bar{\Psi}(z) \mathcal{O}(x-z) \Psi(z) \Big|_{x^0=z^0},$$

is a convolution with the wavepacket $\Psi(x)$.

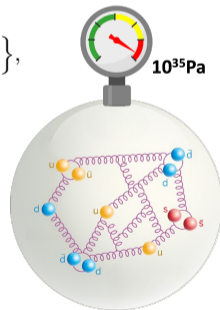
$$\mathcal{E}(x) = M \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot x} \left\{ \left(1 - \frac{q^2}{4M^2} \right) A(q^2) + \frac{q^2}{4M^2} \left[2J(q^2) - D(q^2) \right] \right\},$$

$$\mathcal{P}(x) = \frac{1}{6M} \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot x} q^2 D(q^2),$$

$$\mathcal{S}^{\alpha\beta}(x) = \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot x} \left\{ i\sigma^{\alpha\beta} \sqrt{1 - \frac{q^2}{4M^2}} - \frac{U^{[\alpha} q^{\beta]}}{2M} \right\} J(q^2),$$

$$\Pi^{\alpha\beta}(x) = \frac{1}{4M} \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot x} \left(q^\alpha q^\beta - \frac{q^2}{3} \Delta^{\alpha\beta} \right) D(q^2),$$

$$\Lambda = -M^2 \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot x} \bar{c}(q^2)$$



$$e(x) = \int d^3z \bar{\Psi}(x-z) \left\{ M \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot x} \left\{ \left(1 - \frac{q^2}{4M^2}\right) A(q^2) + \frac{q^2}{4M^2} [2J(q^2) - D(q^2)] \right\} \right\} \Psi(x-z) \Big|_{z^0=0}$$

The **hadronic part** is *not factorizable* due to the dependence of $\vec{P} = (-i/2)\vec{\nabla}_x$ in $q^2 = (q^0)^2 - \vec{q}^2$, where $q^0 = \sqrt{(\vec{P} + \frac{1}{2}\vec{q})^2 + M^2} - \sqrt{(\vec{P} - \frac{1}{2}\vec{q})^2 + M^2}$

- Taylor expansion around $\vec{P} = 0$: multipole series,

$$\mathcal{E}(\vec{r}) = \sum_{n=0}^{\infty} \frac{(-i)^n}{2^n n!} \mathcal{E}_n^{i_1 i_2 \dots i_n}(\vec{r}) \vec{\nabla}^{i_1} \vec{\nabla}^{i_2} \dots \vec{\nabla}^{i_n}$$

- Monopole density gives the Breit-frame distribution (Sachs distribution)

$$\mathcal{E}_0(\vec{r}) = M \int \frac{d^3q}{(2\pi)^3} e^{-i\vec{q} \cdot \vec{r}} \left\{ \left(1 + \frac{\vec{q}^2}{4M^2}\right) A(-\vec{q}^2) - \frac{\vec{q}^2}{4M^2} [2J(-\vec{q}^2) - D(-\vec{q}^2)] \right\}$$

- High-multipole moments exist due to Lorentz distortion
- Note that for spin-0, the monopole density \mathcal{E}_0 differs from the Breit-frame distribution given by Polyakov & Schweitzer

Is the multipole expansion unique? No! \rightarrow Alternative: Taylor (Laurent) expansion around $1/|\vec{P}| = 0$

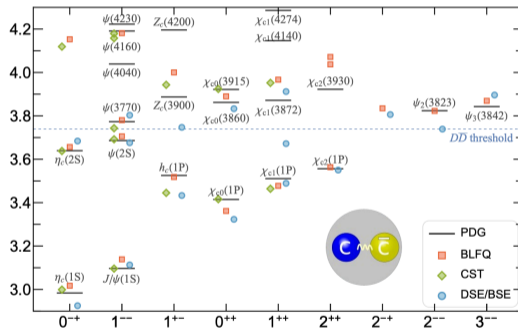
- Sufficient to take $P_z \rightarrow \infty \Rightarrow |\vec{P}| = \sqrt{\vec{P}_\perp^2 + P_z^2} \rightarrow \infty$
- Monopole density gives the **2D light-front distribution**

$$\mathcal{E}_0(x) = \delta(x_\parallel) M \int \frac{d^2 q_\perp}{(2\pi)^2} e^{-i\vec{q}_\perp \cdot \vec{x}_\perp} \left[\left(1 + \frac{\vec{q}_\perp^2}{4M^2} \right) A(-\vec{q}_\perp^2) - \frac{\vec{q}_\perp^2}{4M^2} \left(2J(-\vec{q}_\perp^2) - D(-\vec{q}_\perp^2) \right) \right].$$

- No special frame (e.g. Drell-Yan $q^+ = 0$ frame) is chosen
- Relativistic, suitable also for massless hadrons (in contrast to the Sachs distribution)
- Convergence of the multipole series: $|\vec{P}| \gg \lambda_{\text{hadron}}^{-1} \gg \{M, r_{\text{hadron}}^{-1}\}$ -- sufficiently *localized* z -direction
- In the infinite momentum frame (IMF), components of the EMT form a hierarchy:

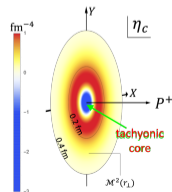
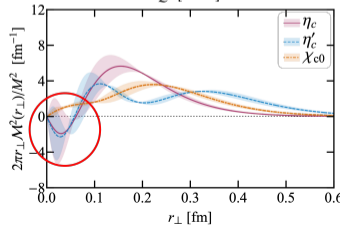
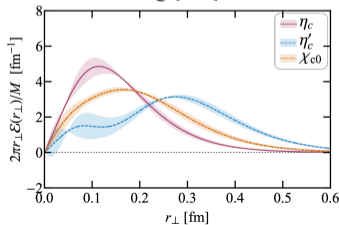
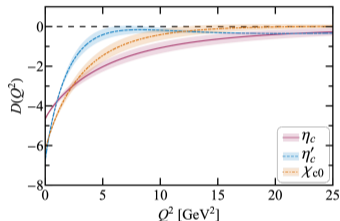
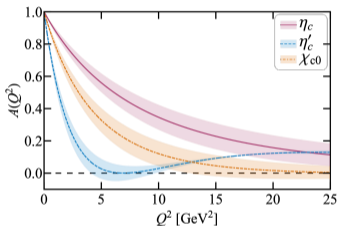
$$\underbrace{\mathcal{T}^{++} \sim P_z^2}_{\text{best}}, \quad \underbrace{\mathcal{T}^{+i} \sim P_z^1}_{\text{good}}, \quad \underbrace{\mathcal{T}^{+-} \sim \mathcal{T}^{ij} \sim P_z^0}_{\text{bad}}, \quad \underbrace{\mathcal{T}^{-i} \sim P_z^{-1}}_{\text{worse}}, \quad \underbrace{\mathcal{T}^{--} \sim P_z^{-2}}_{\text{worst}}$$

$$H = \sum_i \frac{p_{i\perp}^2 + m_i^2}{x_i} + U_i^{(0)} + \sum_{ij} \alpha_s U_{ij}^{(1)} + \sum_{ij} V_{ij}^{\text{QCD}} - U_{ij}$$



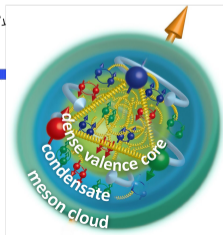
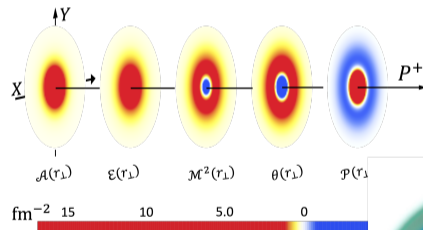
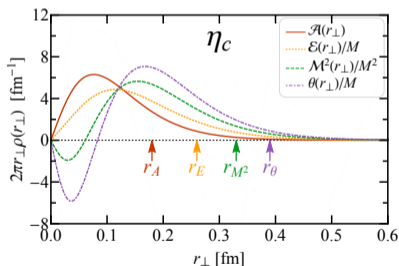
- Basis light-front quantization: two free parameters (m_c, κ), rms deviation: 30 MeV
- Good agreement with the PDG data for both the masses and decay constants and radiative widths

- Obtained GFF D from both T^{+-} (via localizing P^-) and T^{12} , results consistent with each other
- Energy density $\mathcal{E}(r_\perp)$ is positive. However, $\mathcal{M}^2(r_\perp)$ is negative at small r_\perp : tachyonic core within hadrons?



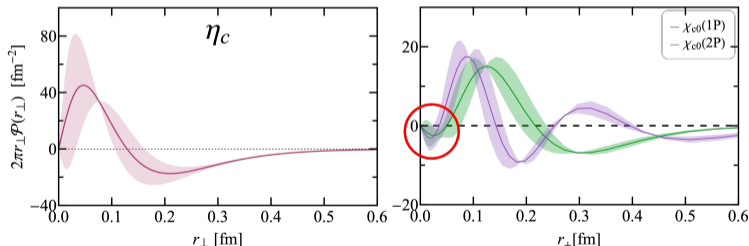
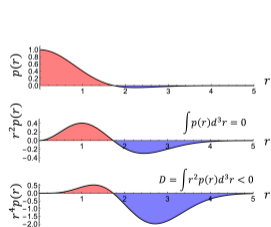
- Number density $\mathcal{A}(r_\perp)$, energy density $\mathcal{E}(r_\perp)$, invariant mass squared density $\mathcal{M}^2(r_\perp)$ and trace scalar density $\theta(r_\perp)$
- Onion like structure -- true for all hadrons with negative D

$$r_A < r_E < r_{M^2} < r_\theta < r_D$$



$$D = \int d^3r r^2 \mathcal{P}(r) \stackrel{???}{<} 0$$

- Speculation: a mechanically stable system must have a repulsive core and an attractive edge
- We find that while η_c has a repulsive core, χ_{c0} has an attractive cores, and both have negative D !



Summary

- Hadrons are unique relativistic systems $r \sim \lambda_C$ which enables us to view them as de Broglie waves. Minkowski-Einstein-Laub's relativistic theory of macroscopic electromagnetism provides the relevant concepts & tools for describing their electromagnetic structure in 3D
- We showed the multipole structures of relativistic currents (first proposed by Born) -- Sachs & light-front distributions can be understood as monopole densities
- I also discussed the relativistic quantum many-body description of form factors as 2D Fourier transform (F.T.) of one-body densities (OBDs) on the light front
- I extended the concepts to the hadronic energy-momentum tensor and discussed the mechanical properties of hadrons
- So far, I focused on the interpretations. Are there any physical consequences of the interpretations? (Gravity, soft-photon, energy conditions, ...)

Thank you!

A scene featuring three Stormtroopers in white armor. The trooper on the left is aiming a blaster upwards, with a red laser beam visible. The trooper in the center is crouching and aiming a blaster. The trooper on the right is aiming a blaster forward, with a bright red light effect at the muzzle. The background is a futuristic, metallic environment with blue and red lighting.

I think we're gonna need

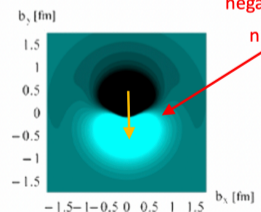
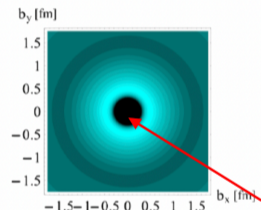
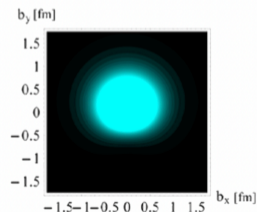
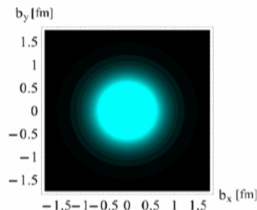
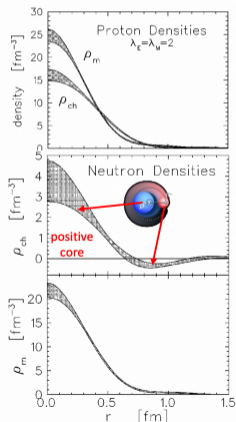
Backup

$$\rho_{\text{ch}}^{\text{Sachs}}(r) = \int \frac{d^3q}{(2\pi)^3} e^{i\vec{q}\cdot\vec{r}} \frac{\langle +\frac{1}{2}\vec{q}_{\perp}, s | J^0(0) | -\frac{1}{2}\vec{q}_{\perp}, s \rangle}{2E}$$

$$\rho_{\text{ch}}^{\text{LF}}(\vec{b}_{\perp}; \vec{S}) = \int \frac{d^2q_{\perp}}{(2\pi)^2} e^{i\vec{q}_{\perp}\cdot\vec{b}_{\perp}} \frac{\langle P^+, +\frac{1}{2}\vec{q}_{\perp}, \vec{S} | J^+(0) | P^+, -\frac{1}{2}\vec{q}_{\perp}, \vec{S} \rangle}{2P^+}$$

proton

neutron



negative core

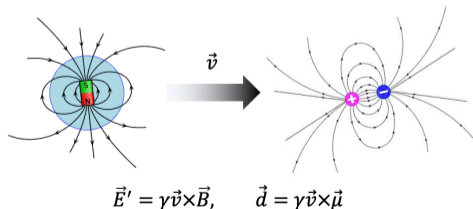
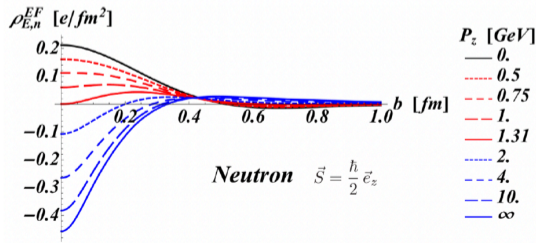
nEDM

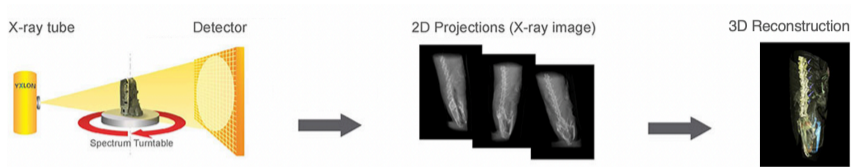
Lorcé interpreted the F.T. as the Weyl function of the current operator,

$$\rho_W(\vec{r}, \vec{P}) = \int \frac{d^3q}{(2\pi)^3} e^{i\vec{q}\cdot\vec{r}} \frac{\langle \vec{P} + \frac{1}{2}\vec{q} | J^0 | \vec{P} - \frac{1}{2}\vec{q} \rangle}{2P^0}$$

- Weyl quantization $O \leftrightarrow O_W(\vec{r}, \vec{p})$, eps. the Wigner function $\varrho \leftrightarrow W(\vec{r}, \vec{p})$
- Sachs distribution: $\vec{P} = 0$; Light-front distribution: $\vec{P} \rightarrow \infty$ (IMF)
- Quasi-probabilistic
- Still need a special frame, $q^0 = 0$ (elastic frame)

$$\overline{O}_\Psi = \int d^3r \int \frac{d^3p}{(2\pi)^3} \times O_W(\vec{r}, \vec{p}) W(\vec{r}, \vec{p})$$





- An Abel-Radon signal at angle \hat{n} and distance s is obtained by the line integral along the path of the X ray.

$$\mathcal{R}f(s, \hat{n}) = \int d^3r f(\vec{r}) \delta(s - \hat{n} \cdot \vec{r})$$

- Inversion problem
- Panteleeva & Polyakov showed with some reasonable assumptions the 3D Sachs and the 2D light-front distributions can be related by the Abel-Radon transformation

[Panteleeva:2021iip]

Example: Gaussian wavepacket

- Normalization of the momentum space wave function:

$$\langle \Psi | \Psi \rangle = 1, \quad \langle p' | p \rangle = 2p^0 (2\pi)^3 \delta^3(p - p') \quad \Rightarrow \quad \int \frac{d^3p}{(2\pi)^3 2p^0} \tilde{\Psi}^*(\vec{p}) \tilde{\Psi}(\vec{p}) = 1$$

where, $p^0 = \sqrt{\vec{p}^2 + M^2}$ is the on-shell energy.

- Gaussian wavepacket:

$$\tilde{\Psi}(\vec{p}) = N_\sigma e^{-\frac{\vec{p}^2}{2\sigma^2}}$$

where, σ is the width of the Gaussian, and $N_\sigma = 4\pi^{\frac{3}{4}} / [\sigma U^{-\frac{1}{2}}(\frac{1}{2}, 0, M^2/\sigma^2)]$. $U(a, b, c)$ is the confluent hypergeometric function of the second kind. Plane wave limit: $\sigma \rightarrow 0$ and localization limit $\sigma \rightarrow \infty$

- Normalization of the coordinate-space wave function,

$$2M \int d^3x |\Psi(x)|^2 = \frac{M}{\sigma} \frac{U(1, \frac{1}{2}, \frac{M^2}{\sigma^2})}{U(\frac{1}{2}, 0, \frac{M^2}{\sigma^2})} = \begin{cases} 1 - \frac{3\sigma^2}{4M^2} + O(\frac{\sigma^4}{M^4}) & (\sigma \rightarrow 0), \\ \frac{\sqrt{\pi}M}{\sigma} - \frac{\pi M^2}{\sigma^2} + O(\frac{M^3}{\sigma^3}) & (\sigma \rightarrow \infty) \end{cases}$$

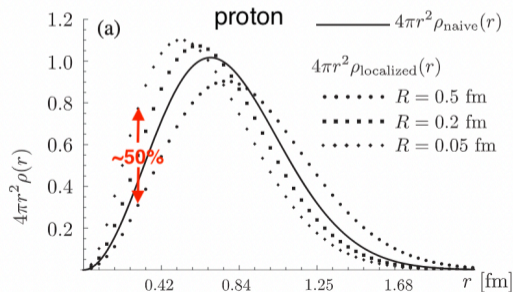
The wave function cannot be normalized in the localization limit or the massless limit.

- Consider a Gaussian wavepacket:

$$r_{\Psi}^2 \equiv \int d^3x \vec{x}^2 \langle \Psi | J^0(x) | \Psi \rangle = 6F'(0) + 3R_{\Psi}^2$$

where, R_{Ψ} is the size of the wavepacket.

- Sachs distribution is valid (i.e. $r_{\Psi}^2 \approx 6F'(0)$) if, $r_{\text{hadron}} \gg Q_{\text{max}}^{-1} \gg R_{\Psi} \gg M^{-1}$
- However, for hadrons $r_{\text{hadron}} M \sim 1$
- Factorization approach works for $R_{\Psi} \gg r_{\text{hadron}}$

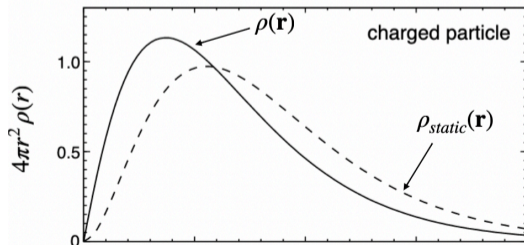
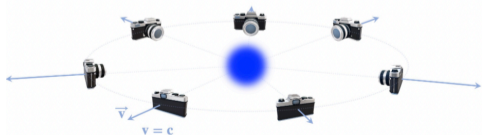


	r	M	rM
pion	0.67 fm	0.14 GeV	0.5
charmonium	0.25--0.4 fm	3.0 GeV	3.8--6
proton	0.87 fm	0.94 GeV	4
nuclei	$1.3 A^{\frac{1}{3}} \text{ fm}$	$0.94A \text{ GeV}$	$6A^{\frac{4}{3}}$

Epelbaum et al. proposed to use sharply localized wavepacket $R_\Psi \rightarrow 0$ with spherical symmetry,

$$\rho(\vec{r}) = \int \frac{d^3}{(2\pi)^3} e^{-i\vec{q}\cdot\vec{r}} \int_{-1}^{+1} d\alpha \frac{1}{2} F_{\text{ch}} \left[- (1 - \alpha^2) \vec{q}^2 \right]$$

- Angle-averaged light-front density
- Particle localization problem in QFTs
- $R_\Psi \rightarrow 0$ limit is not well defined, e.g. $R_x = 2R_y \rightarrow 0$ gives different results



$$\begin{aligned}
 j^\alpha(x) &= \int \frac{d^3P}{(2\pi)^3} \int \frac{d^3q}{(2\pi)^3} \tilde{\Phi}^*(\vec{P} + \frac{1}{2}\vec{q}) \tilde{\Phi}(\vec{P} - \frac{1}{2}\vec{q}) \frac{P^\alpha}{2p^0 p'^0} F_{\text{ch}}(q^2) e^{iq \cdot x}, \\
 &= \int d^3x_1 \Phi^*(\vec{x}_1, t) \int d^3x_2 i\partial^\alpha \Phi(\vec{x}_2, t) \int \frac{d^3P}{(2\pi)^3} e^{i\vec{P} \cdot (\vec{x}_1 - \vec{x}_2)} \int \frac{d^3q}{(2\pi)^3} F_{\text{ch}}(q^2) e^{-i\vec{q} \cdot (\vec{x} - \frac{\vec{x}_1 + \vec{x}_2}{2})} \\
 &\quad - \int d^3x_1 i\partial^\alpha \Phi^*(\vec{x}_1, t) \int d^3x_2 \Phi(\vec{x}_2, t) \int \frac{d^3P}{(2\pi)^3} e^{i\vec{P} \cdot (\vec{x}_1 - \vec{x}_2)} \int \frac{d^3q}{(2\pi)^3} F_{\text{ch}}(q^2) e^{-i\vec{q} \cdot (\vec{x} - \frac{\vec{x}_1 + \vec{x}_2}{2})}
 \end{aligned}$$

The above integral is not factorizable because q^0 is not independent of \vec{P} :

$$q^0 = \sqrt{(\vec{P} + \frac{1}{2}\vec{q})^2 + M^2} - \sqrt{(\vec{P} - \frac{1}{2}\vec{q})^2 + M^2}$$

However, we can evaluate the d^3P integration, provided all $\vec{P} \rightarrow (-i/2)\vec{\nabla}$ inside q^0 :

$$j^\alpha(x) = \int d^3r \Phi(\vec{r}, t) i\vec{\partial}_r^\alpha D(\vec{x} - \vec{r}; \frac{-i}{2}\vec{\nabla}_r) \Phi(\vec{r}, t)$$

where, the intrinsic density is identified as,

$$D(\vec{x}; \vec{P}) = \int \frac{d^3q}{(2\pi)^3} F_{\text{ch}}(q^2) e^{-i\vec{q} \cdot \vec{x}}$$

In Drell-Yan frame ($\omega \cdot q = 0$):

$$\langle p' | J^\mu(0) | p \rangle = 2P^\alpha F(-q^2) + \frac{M^2 \omega^\mu}{\omega \cdot P} S(-q^2),$$

where, $P = (p' + p)/2$, $q = p' - p$. $\omega^\mu = (\omega^+, \omega^-, \vec{\omega}_\perp) = (0, 2, 0)$ is a null vector indicating the orientation of the quantization surface.

- Emergence of spurious form factors S due to the violation of dynamical Lorentz symmetries in practical calculations, which usually contain uncanceled divergences
- EM current is automatically conserved in the Drell-Yan frame $q^+ = 0$
- Identify J^+ , \vec{J}_\perp as the good currents that are free of spurious form factors or divergence

In Drell-Yan frame ($\omega \cdot q = 0$):

$$\begin{aligned} \langle p' | T_i^{\alpha\beta}(0) | p \rangle = & 2P^\alpha P^\beta A_i(-q^2) + \frac{1}{2}(q^\alpha q^\beta - q^2 g^{\alpha\beta}) D_i(-q^2) + 2M^2 g^{\alpha\beta} \bar{c}_i(-q^2) \\ & + \frac{M^4 \omega^\alpha \omega^\beta}{(\omega \cdot P)^2} S_{1i}(-q^2) + (V^\alpha V^\beta + q^\alpha q^\beta) S_{2i}(-q^2), \end{aligned}$$

where, $P = (p' + p)/2$, $q = p' - p$. $\omega^\mu = (\omega^+, \omega^-, \vec{\omega}_\perp) = (0, 2, 0)$ is a null vector indicating the orientation of the quantization surface. Vector V^α is defined as $V^\alpha = \varepsilon^{\alpha\beta\rho\sigma} P_\beta q_\rho \omega_\sigma / (\omega \cdot P)$.

- Emergence of spurious form factors $S_{1,2}$ due to the violation of dynamical Lorentz symmetries in practical calculations, which usually contain uncanceled divergences
- Identify T^{++} , T^{+i} , T^{12} , T^{+-} as the good currents that are free of spurious form factors or divergence

$$t_i^{++} = 2(P^+)^2 A_i(q_\perp^2),$$

$$t_i^{12} = \frac{1}{2} q_\perp^1 q_\perp^2 D_i(q_\perp^2),$$

$$t_i^{+-} = 2(M^2 + \frac{1}{4} q_\perp^2) A_i(q_\perp^2) + q_\perp^2 D_i(q_\perp^2) + 4M^2 \bar{c}_i(q_\perp^2)$$

$$t_i^{--} = 2 \left(\frac{M^2 + \frac{1}{4} q_\perp^2}{P^+} \right)^2 A_i(q_\perp^2) + \frac{4M^4}{(P^+)^2} S_{1i}(q_\perp^2).$$

$$t_i^{11} + t_i^{22} = -\frac{1}{2} q_\perp^2 D_i(q_\perp^2) - 4M^2 \bar{c}_i(q_\perp^2) + 2q_\perp^2 S_{2i}(q_\perp^2).$$