# ANALYSIS OF THE HDO ABSORPTION SPECTRUM BETWEEN 9600－10200 $\mathrm{cm}^{-1^{*}}$ 

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#### Abstract

The absorption spectrum of $\mathrm{HD}^{16} \mathrm{O}$ was recorded at a resolution of $0.02 \mathrm{~cm}^{-1}$ with a Bruker IFS 120HR Fourier Transform Spectrometer in the region of $9600-10200 \mathrm{~cm}^{-1}$ ．As far as we know it is the first time to record and analyze the HDO spectrum in this region which was assigned as the $\nu_{1}+2 \nu_{3}$ and $2 \nu_{2}+2 \nu_{3}$ bands．With the strong resonance interactions between these two bands considered，the spectroscopic parameters were optimized by the nonlinear least squares method．


Keywords：Vib－rotational spectrum，HDO

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## I．INTRODUCTION

The investigation of high resolution of $\mathrm{H}_{2} \mathrm{O}$ and its major isotopomers is strongly required to inter－ pret the atmospheric spectra and calculations of the absorption due to minor atmospheric constituents in the atmospheric transmittance windows．As an asym－ metric isotopic modification of water，the spectrum of HDO is studied and gives additional information to determine the intramolecular potential function of the water molecule．Since the high stretching vibrational states of HDO have the delocalized nature，it has been a focus of significant attention as the molecule to elu－ cidate photofragmentation and chemical reaction dy－ namics of vibrational quantum state selection ${ }^{[1]}$ ．More high resolution spectroscopic studies are required to realize the goal of the bond selective chemistry reac－ tion．

The first extended study of the HDO infrared spectrum was carried out by Benedict et al．${ }^{[2]}$ from 2400 to $8000 \mathrm{~cm}^{-1}$ ．The most extensive study was fulfilled by Toth and co－authors，${ }^{[3]}$ which covers the region from 600 to $7700 \mathrm{~cm}^{-1}$ ．The $3 \nu_{2}$ and $\nu_{1}+\nu_{2}$ bands which in the region of $3900-4400 \mathrm{~cm}^{-1}$ were in－ vestigated by Perrin et al．${ }^{[4]}$ The $\nu_{2}$ and $\nu_{1} / 2 \nu_{2}$ bands were analyzed by Guelachvili ${ }^{[5]}$ and Papineau et al．，${ }^{[6]}$ respectively．Ohshima et al．${ }^{[7]}$ recorded the spectrum of HDO from 6380 to $6600 \mathrm{~cm}^{-1}$ ．More spectra ${ }^{[8]}$ have been analyzed with the wave number higher than 8500 $\mathrm{cm}^{-1}$ ．But in the region of $9600-10200 \mathrm{~cm}^{-1}$ ，there has been no discussion on it yet．

The rotational energy levels of the ground vibra－ tional state and spectroscopic constants of the Wat－ son A－reduced Hamiltonian have been reported by Johns ${ }^{[9]}$ and Papineau et al．，${ }^{[6]}$ respectively．

We present here the first detailed high resolution Fourier transform spectroscopic observation and the analysis of the HDO spectrum in the region of $9600-$ $10200 \mathrm{~cm}^{-1}$ ．The observed（022）and（102）states of HDO interact with each other through both the Fermi and the Coriolis interactions．With these resonances considered，a good agreement between the experimen－ tal data and the calculated levels is achieved．

## II．EXPERIMENTAL DETAILS

The sample of $\mathrm{D}_{2}^{16} \mathrm{O}$ was purchased from Bei－ jing Chemical Industry，Ltd．（China）．The stated abundance of deuterium was $99.8 \%$ ．The sample of HDO was prepared by mixing the $\mathrm{D}_{2} \mathrm{O}$ and $\mathrm{H}_{2} \mathrm{O}$ equally．The spectrum was recorded at room tempera－ ture with the Bruker IFS 120HR Fourier transform in－ terferometer，which was equipped with a multipass gas cell with adjustable path length，a tungsten source，a $\mathrm{CaF}_{2}$ beamsplitter，and a Si diode detector．The un－ apodized resolution was $0.02 \mathrm{~cm}^{-1}$ ，and the apodiza－ tion function was Blackman－Harris 3－Term．The sam－ ple pressure was 1500 Pa ，which was measured by a manometer with a stated accuracy of 20 Pa ．The to－ tal path length was 105 m ．

Since there are many absorption lines due to the $\mathrm{H}_{2} \mathrm{O}$ and $\mathrm{D}_{2} \mathrm{O}$ in the region under the study，we use

[^0]the spectrum of "pure" $\mathrm{D}_{2} \mathrm{O}$, which was also recorded before in our laboratory, and the $\mathrm{H}_{2} \mathrm{O}$ lines listed in the Hitran 96 database in this region to identify the absorption lines of HDO. The frequency was also calibrated with these $\mathrm{H}_{2} \mathrm{O}$ lines in the Hitran96 database. The estimated precision in positions of unblended lines was $2 \times 10^{-3} \mathrm{~cm}^{-1}$.

## III. ROTATIONAL ANALYSIS

The HDO is a prolate asymmetric top molecule of the $C_{s}$ type symmetry. As a consequence, any of its vibrational-rotational band contains absorption lines arising from transitions of the two types (both A and B) .

In the region of $9600-10200 \mathrm{~cm}^{-1}$, the origins of the $\nu_{1}+2 \nu_{3}$ and $2 \nu_{2}+2 \nu_{3}$ bands are predicted by Jensen et al. ${ }^{[10]}$ Since these two bands are located closely, strong resonance interactions exist. So the Hamiltonian model which we used in our theoretical rotational analysis should take into account two interacting vibrational states:
usual Watson's operators ${ }^{[11]}$ :

$$
\begin{align*}
H^{v v}= & E^{v}+\left[A^{v}-\frac{1}{2}\left(B^{v}+C^{v}\right)\right] J_{z}^{2} \\
& +\frac{1}{2}\left(B^{v}+C^{v}\right) J^{2}+\frac{1}{2}\left(B^{v}-C^{v}\right) J_{x y}^{2} \\
& -\Delta_{K}^{v} J_{z}^{4}-\Delta_{J K}^{v} J_{z}^{2} J^{2}-\Delta_{J}^{v} J^{4} \\
& -\delta_{K}^{v}\left[J_{z}^{2}, J_{x y}^{2}\right]_{+}-2 \delta_{J}^{v} J^{2} J_{x y}^{2} \\
& +H_{K}^{v} J_{z}^{6}+H_{K J}^{v} J_{z}^{4} J^{2}+H_{J K}^{v} J_{z}^{2} J^{4}+H_{J}^{v} J^{6} \\
& +\left[J_{x y}^{2}, h_{K}^{v} J_{z}^{4}+h_{J K}^{v} J^{2} J_{z}^{2}+h_{J}^{v} J^{4}\right]_{+} \\
& +L_{K}^{v} J_{z}^{8}+L_{K K J}^{v} J_{z}^{6} J^{2}+\ldots+P_{K}^{v} J_{z}^{10}+\ldots \tag{2}
\end{align*}
$$

where $|1\rangle=(022),|2\rangle=(102)$. The $H^{v v^{\prime}}\left(v \neq v^{\prime}\right)$ account for the operators of resonance interactions. In this case, resonance operators $H^{v v^{\prime}}$ should be written in the form of the sum of two terms

$$
\begin{equation*}
H^{v v^{\prime}}=H_{F}^{v v^{\prime}}+H_{C}^{v v^{\prime}} \tag{3}
\end{equation*}
$$

which describes resonance interactions of the Fermi type:

$$
\begin{align*}
H_{F}^{v v^{\prime}}= & F_{0}^{v v^{\prime}}+F_{K}^{v v^{\prime}} J_{z}^{2}+F_{J}^{v v^{\prime}} J^{2} \\
& +F_{K K}^{v v^{\prime}} J_{z}^{4}+\ldots+F_{x y}^{v v^{\prime}} J_{x y}^{2}+\ldots \tag{4}
\end{align*}
$$

and Coriolis type:

$$
\begin{align*}
H_{C}^{v v^{\prime}}= & C_{y}\left(i J_{y}\right)+C_{y K}^{v v^{\prime}}\left[i J_{y}, J_{z}^{2}\right]_{+}+\ldots  \tag{5}\\
& +C_{x z}^{v v^{\prime}}\left[J_{x}, J_{z}\right]_{+}+C_{x z J}^{v v^{\prime}}\left[J_{x}, J_{z}\right]_{+} J^{2}+. . \tag{5}
\end{align*}
$$

Here the diagonal operators $H^{v v}(v=1,2)$ are the

Fig.1. Parts of the spectrum of the (102) and (022) bands of HDO. The label $J_{K a^{\prime} K c^{\prime}}^{\prime}-J_{K a^{\prime \prime} K c^{\prime \prime}}^{\prime \prime}\left(n_{1} n_{2} n_{3}\right)$ denotes the transitions from the ground state level $J_{K a^{\prime \prime} K c^{\prime \prime}}^{\prime \prime}$ to the upper state level $J_{K a^{\prime} K c^{\prime}}^{\prime}$ of the upper band $\left(n_{1} n_{2} n_{3}\right)$. Lines marked by " $\mathrm{H}_{2} \mathrm{O}$ " are belonged to the absorption of water.

In Eqs.[2]-[5] the following notations are used: $J_{x y}^{2}=J_{x}^{2}-J_{y}^{2}, J^{2}=\sum_{\alpha} J_{\alpha}^{2}(\alpha=x, y$ and $z)$ and $[A, B]_{+}=A B+B A$. In our fit, the Hamiltonian in Eq. 1 is transformed by a contact transformation to a form in which the leading term in Eq. 4 is set to zero, i.e., $F_{0}^{v v^{\prime}}=0$.

Parts of the spectrum of these two bands are illustrated in Fig.2. The expression $J_{K a^{\prime} K c^{\prime}}^{\prime}-J_{K a^{\prime \prime} K c^{\prime \prime}}^{\prime \prime}$ $\left(n_{1} n_{2} n_{3}\right)$ denotes the transitions from the ground state level $J_{K a^{\prime \prime} K c^{\prime \prime}}^{\prime \prime}$ to the upper state level $J_{K a^{\prime} K c^{\prime}}^{\prime}$ of the upper band ( $n_{1} n_{2} n_{3}$ ).

The transitions were assigned by means of the ground state combination differences (GSCD) method. The ground state rotational energies were calculated with the basis parameters taken from Ref.[6].

As a result of our analysis, about 650 transitions were assigned. And 149 upper energy levels were obtained with $J^{\prime}$ up to 11 and $K a^{\prime}$ up to 6 , which are listed in Table 1. The spectroscopic parameters listed in Table 2 are obtained by nonlinear least squares fitting.

Table 1. Energy Levels $\left(\mathrm{cm}^{-1}\right)$ of the (022) and (102) Normal Mode States of HDO.

| $J$ | $K a$ | Kc | (022) |  |  | (102) |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | Obs. | Cal. | Obs. - Cal. | Obs. | Cal. | Obs. - Cal. |
| 0 | 0 | 0 | 9934.772 | 9934.777 | -0.005 | 9967.014 | 9967.011 | 0.003 |
| 1 | 0 | 1 | 9950.060 | 9950.066 | -0.006 | 9982.235 | 9982.237 | -0.002 |
| 1 | 1 | 1 | 9964.563 | 9964.565 | -0.001 | 9995.908 | 9995.902 | 0.007 |
| 1 | 1 | 0 | 9967.617 | 9967.618 | -0.001 | 9998.878 | 9998.876 | 0.002 |
| 2 | 0 | 2 | 9980.186 | 9980.189 | -0.003 | 10012.248 | 10012.255 | -0.008 |
| 2 | 2 | 0 | 10043.207 | 10043.206 | 0.001 | 10075.149 | 10075.147 | 0.002 |
| 2 | 1 | 2 | 9992.028 | 9992.026 | 0.002 | 10023.416 | 10023.407 | 0.008 |
| 2 | 2 | 1 | 10042.784 | 10042.786 | -0.001 | 10074.712 | 10074.710 | 0.002 |
| 2 | 1 | 1 | 10001.161 | 10001.157 | 0.005 | 10032.331 | 10032.333 | -0.001 |
| 3 | 0 | 3 | 10024.302 | 10024.304 | -0.002 | 10056.249 | 10056.261 | -0.012 |
| 3 | 2 | 1 | 10090.337 | 10090.334 | 0.003 | 10122.754 | 10122.757 | -0.003 |
| 3 | 1 | 3 | 10032.953 | 10032.953 | 0.001 | 10064.418 | 10064.411 | 0.008 |
| 3 | 3 | 1 | 10160.565 | 10160.558 | 0.007 | 10202.138 | 10202.143 | -0.005 |
| 3 | 2 | 2 | 10088.302 | 10088.305 | -0.003 | 10120.633 | 10120.632 | 0.001 |
| 3 | 1 | 2 | 10051.100 | 10051.092 | 0.009 | 10082.226 | 10082.227 | -0.001 |
| 3 | 3 | 0 | 10160.602 | 10160.597 | 0.005 | 10202.180 | 10202.182 | -0.003 |
| 4 | 0 | 4 | 10081.411 | 10081.413 | -0.002 | 10113.294 | 10113.306 | -0.012 |
| 4 | 2 | 2 | 10154.357 | 10154.350 | 0.007 | 10187.539 | 10187.545 | -0.006 |
| 4 | 4 | 0 | 10316.904 | 10316.905 | -0.000 | 10379.055 | 10379.061 | -0.006 |
| 4 | 1 | 4 | 10087.080 | 10087.078 | 0.002 | 10118.660 | 10118.654 | 0.006 |
| 4 | 3 | 2 | 10221.537 | 10221.541 | -0.004 | 10264.379 | 10264.378 | 0.001 |
| 4 | 2 | 3 | 10147.722 | 10147.720 | 0.002 | 10181.515 | 10181.517 | -0.002 |
| 4 | 4 | 1 | 10316.904 | 10316.902 | 0.002 | 10379.058 | 10379.058 | 0.000 |
| 4 | 1 | 3 | 10116.937 | 10116.928 | 0.008 | 10149.066 | 10149.069 | -0.003 |
| 4 | 3 | 1 | 10221.807 | 10221.807 | -0.000 | 10264.649 | 10264.648 | 0.001 |
| 5 | 0 | 5 | 10150.742 | 10150.745 | -0.003 | 10182.654 | 10182.663 | -0.008 |
| 5 | 2 | 3 | 10235.581 | 10235.577 | 0.003 | 10269.949 | 10269.954 | -0.005 |
| 5 | 4 | 1 | 10392.821 | 10392.820 | 0.001 | 10456.934 | 10456.925 | 0.008 |
| 5 | 1 | 5 | 10154.125 | 10154.123 | 0.002 | 10185.868 | 10185.864 | 0.004 |
| 5 | 3 | 3 | 10297.842 | 10297.851 | -0.008 | 10342.311 | 10342.311 | 0.000 |
| 5 | 5 | 1 |  |  |  | 10602.665 | 10602.665 | -0.000 |
| 5 | 2 | 4 | 10223.120 | 10223.123 | -0.003 | 10257.072 | 10257.075 | -0.003 |
| 5 | 4 | 2 | 10392.794 | 10392.794 | 0.000 | 10456.906 | 10456.900 | 0.006 |

Table 1. (Continued)

| 5 | 1 | 4 | 10197.942 | 10197.938 | 0.005 | 10229.862 | 10229.851 | 0.011 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 5 | 3 | 2 | 10298.879 | 10298.882 | -0.003 | 10343.366 | 10343.361 | 0.005 |
| 5 | 5 | 0 |  |  |  | 10602.665 | 10602.665 | 0.000 |
| 6 | 0 | 6 | 10231.957 | 10231.960 | -0.003 | 10263.998 | 10264.003 | -0.005 |
| 6 | 2 | 4 | 10333.798 | 10333.804 | -0.006 | 10369.910 | 10369.918 | -0.008 |
| 6 | 4 | 2 | 10484.177 | 10484.175 | 0.002 | 10550.648 | 10550.648 | -0.000 |
| 6 | 1 | 6 | 10233.837 | 10233.836 | 0.002 | 10265.797 | 10265.793 | 0.004 |
| 6 | 3 | 4 | 10389.394 | 10389.404 | -0.011 |  |  |  |
| 6 | 6 | 0 |  |  |  | 10870.109 | 10870.110 | -0.001 |
| 6 | 5 | 2 |  |  |  | 10695.876 | 10695.873 | 0.002 |
| 6 | 2 | 5 | 10311.933 | 10311.935 | -0.002 | 10346.953 | 10346.958 | -0.005 |
| 6 | 4 | 3 | 10484.046 | 10484.046 | -0.000 | 10550.526 | 10550.525 | 0.001 |
| 6 | 1 | 5 | 10293.152 | 10293.150 | 0.002 | 10325.957 | 10325.939 | 0.019 |
| 6 | 3 | 3 | 10392.337 | 10392.338 | -0.001 | 10438.882 | 10438.879 | 0.003 |
| 6 | 6 | 1 |  |  |  | 10870.108 | 10870.110 | $-0.001$ |
| 6 | 5 | 1 |  |  |  | 10695.879 | 10695.875 | 0.004 |
| 7 | 0 | 7 | 10325.022 | 10325.025 | -0.003 | 10357.284 | 10357.285 | -0.001 |
| 7 | 2 | 5 | 10486.891 | 10486.896 | -0.005 | 10448.380 | 10448.395 | -0.015 |
| 7 | 4 | 3 | 10591.173 | 10591.171 | 0.002 | 10660.403 | 10660.422 | -0.019 |
| 7 | 1 | 7 | 10326.018 | 10326.016 | 0.002 | 10358.240 | 10358.240 | -0.000 |
| 7 | 3 | 5 | 10496.043 | 10496.053 | -0.010 | 10544.920 | 10544.900 | 0.020 |
| 7 | 5 | 3 |  |  |  | 10804.777 | 10804.772 | 0.005 |
| 7 | 2 | 6 | 10414.398 | 10414.396 | 0.002 | 10450.775 | 10450.780 | -0.005 |
| 7 | 4 | 4 | 10590.709 | 10590.714 | -0.005 | 10659.969 | 10659.984 | -0.015 |
| 7 | 1 | 6 | 10401.465 | 10401.461 | 0.004 | 10435.688 | 10435.671 | 0.017 |
| 7 | 3 | 4 | 10502.773 | 10502.763 | 0.011 | 10551.851 | 10551.857 | -0.006 |
| 7 | 5 | 2 |  |  |  | 10804.779 | 10804.784 | -0.005 |
| 8 | 0 | 8 | 10430.011 | 10430.012 | $-0.001$ | 10462.566 | 10462.564 | 0.002 |
| 8 | 2 | 6 | 10616.810 | 10616.826 | -0.016 | 10578.456 | 10578.471 | -0.015 |
| 8 | 1 | 8 | 10430.521 | 10430.517 | 0.004 | 10463.058 | 10463.058 | 0.000 |
| 8 | 3 | 6 | 10620.776 | 10620.788 | -0.013 | 10669.088 | 10669.087 | 0.001 |
| 8 | 4 | 4 |  |  |  | 10786.538 | 10786.517 | 0.020 |
| 8 | 2 | 7 | 10530.130 | 10530.127 | 0.003 | 10568.139 | 10568.144 | -0.004 |
| 8 | 1 | 7 | 10521.902 | 10521.895 | 0.007 | 10558.039 | 10558.035 | 0.004 |
| 8 | 3 | 5 | 10630.602 | 10630.571 | 0.031 | 10682.814 | 10682.805 | 0.009 |
| 8 | 4 | 5 |  |  |  | 10785.255 | 10785.265 | -0.010 |
| 9 | 0 | 9 | 10546.987 | 10546.988 | -0.001 | 10579.895 | 10579.892 | 0.003 |
| 9 | 2 | 7 | 10752.794 | 10752.788 | 0.006 | 10722.992 | 10722.998 | -0.006 |
| 9 | 1 | 9 | 10547.240 | 10547.239 | 0.001 | 10580.140 | 10580.142 | -0.003 |
| 9 | 3 | 7 |  |  |  | 10808.054 | 10808.053 | 0.001 |
| 9 | 2 | 8 | 10658.773 | 10658.770 | 0.003 | 10698.671 | 10698.673 | -0.002 |
| 9 | 1 | 8 | 10653.887 | 10653.879 | 0.008 | 10692.326 | 10692.328 | -0.002 |
| 10 | 0 | 10 | 10675.989 | 10675.990 | -0.001 | 10709.297 | 10709.296 | 0.001 |
| 10 | 1 | 10 | 10676.113 | 10676.113 | 0.000 | 10709.418 | 10709.421 | -0.003 |
| 10 | 2 | 9 | 10800.021 | 10800.022 | -0.001 | 10842.034 | 10842.035 | -0.001 |
| 10 | 1 | 9 | 10797.272 | 10797.271 | 0.000 | 10838.244 | 10838.250 | -0.006 |
| 11 | 2 | 10 | 10953.646 | 10953.644 | 0.002 | 10997.965 | 10997.962 | 0.003 |
| 11 | 1 | 10 | 10952.159 | 10952.161 | -0.002 |  |  |  |

Table 2. Spectroscopic Parameters of the (022) and (102) Vibrational States of HDO Molecule (in $\left.\mathrm{cm}^{-1}\right)^{a}$.

| Parameter | (022) | (102) |
| :---: | :---: | :---: |
| $E$ | 9934.7756(31) | 9967.0114(33) |
| A | $23.8225(24)$ | 22.6201(22) |
| $B$ | 9.2185(12) | 9.10905(43) |
| C | $6.1222(13)$ | $6.11087(24)$ |
| $\Delta_{K} \times 10^{2}$ | 4.850(37) | -0.296(30) |
| $\Delta_{J K} \times 10^{2}$ | 0.3216(46) | -0.1901(80) |
| $\Delta_{J} \times 10^{2}$ | 0.04491(28) | 0.04391(37) |
| $\delta_{K} \times 10^{2}$ | 0.4508(39) | $0.1027(50)$ |
| $\delta_{J} \times 10^{2}$ | $0.01979(16)$ | $0.01354(20)$ |
| $H_{K} \times 10^{5}$ | $32.8(21)$ | 12.6(14) |
| $H_{K J} \times 10^{5}$ | $-7.88(23)$ | 10.51(76) |
| $H_{J K} \times 10^{5}$ | $0.24{ }^{\text {b }}$ | $0.24{ }^{\text {b }}$ |
| $H_{J} \times 10^{5}$ | $0.0039^{\text {b }}$ | $0.0039{ }^{\text {b }}$ |
| $h_{K} \times 10^{5}$ | 14.48(31) | $-16.73(50)$ |
| $h_{J K} \times 10^{5}$ | $0.10^{\text {b }}$ | $0.10^{\text {b }}$ |
| $h_{J} \times 10^{5}$ | $0.0021{ }^{\text {b }}$ | $0.0021{ }^{\text {b }}$ |
| $L_{K} \times 10^{7}$ | 98.6(31) | -64.0(20) |
| $L_{K K J} \times 10^{7}$ | $0.86{ }^{\text {b }}$ | -19.3(20) |
| $l_{K} \times 10^{7}$ | $-1.9{ }^{\text {b }}$ | $-1.9{ }^{\text {b }}$ |
| $P_{K} \times 10^{9}$ | $0.58{ }^{\text {b }}$ | $0.58{ }^{\text {b }}$ |
| Fermi Type Interactions |  |  |
| $F_{K}^{022-102}$ | $1.7503(22)$ |  |
| $F_{J}^{022-102}$ | 0.08225(69) |  |
| $F_{x y}^{022-102}$ | 0.12400(83) |  |
| $F_{K K}^{022-102}$ | $0.005050(44)$ |  |
| Coriolis Type Interactions |  |  |
| $C_{y}^{022-102}$ | -0.925(21) |  |
| $C_{x z}^{022-102}$ | -0.575(10) |  |

${ }^{\text {a }}$ Values in parentheses are the $1 \sigma$ statistical confidence intervals in the last significant digit. Parameters presented without confidence intervals were fixed to their initial values (see text and footnotes below for details).
${ }^{\mathrm{b}}$ Fixed to the values of corresponding parameters of the ground vibrational state from Ref.[6].

A statistical analysis of the residuals shows a satisfactory agreement between the observed and calculated energies. For 149 levels which we used to fit the (022) and (102) states,

$$
\begin{aligned}
\delta & \leq 0.005 & & 67.8 \% \text { of the levels } \\
0.005 & \leq \delta \leq 0.01 & & 21.5 \% \text { of the levels } \\
0.01 & \leq \delta \leq 0.02 & & 8.7 \% \text { of the levels } \\
0.02 & \leq \delta \leq 0.04 & & 2.0 \% \text { of the levels }
\end{aligned}
$$

with $\delta=\left|E_{\text {obs. }}-E_{\text {cal. }}\right|\left(\right.$ in $\left.\mathrm{cm}^{-1}\right)$. We can see that $98 \%$ of the energy levels are reproduced by our fit with errors less than $0.02 \mathrm{~cm}^{-1}$, which is the unapodized resolution of the spectrum.

In our fit, the transformed Hamiltonian we used leads to a mixing of the rotational and coupling constants, so that the initial vib-rotational and resonance parameters cannot be determined well, and the upper state energy levels with high $K a$ quantum number are not able to be predicted well.

To resolve this problem, first the low J energy levels with $K a=0,1$ are assigned, then these levels are used to fit the interaction parameters. After these interaction parameters are obtained, the higher J and Ka energy levels are predicted well. Then more lines are assigned and more levels are involved to fit more high-order parameters.

Those high-order parameters, such as $H_{J k}, H_{K}$, $h_{J k}, h_{J}$, and $P_{K}$, cannot be significantly determined and are fixed to their ground state values.

## IV. DISCUSSION

The highly excited vibrational states (102) and (022) of $\mathrm{HD}^{16} \mathrm{O}$ have been investigated. Line assignments and energy levels are determined and rotational constants of the Watson A - reduced Hamiltonian are obtained. We can see that the obtained parameters have successfully reproduced the upper state levels.

The Fermi and the Coriolis type resonances are significant since the difference of the two band origins is small. It is also shown that there exist resonances in the $\left(n_{1}, n_{2}, n_{3}\right)$ and ( $\left.n_{1}-1, n_{2}+2, n_{3}\right)$ bands because the wave numbers of $\nu_{1}$ and $2 \nu_{2}$ are close to each other $\left(\nu_{1}=2723.6795 \mathrm{~cm} .^{-1}\right.$ and $2 \nu_{2}=2792.0115$ $\left.\mathrm{cm}^{-1}\right) .{ }^{[6]}$ Ohshima et al. ${ }^{[7]}$ got the interaction parameters $F_{0}, F_{K}$, and $C_{y}$ by studying the (101) and (021) bands. Perrin et al. ${ }^{[4]}$ recorded the $3 \nu_{2}$ and $\nu_{1}+\nu_{2}$ bands and got the interaction parameters $F_{0}, F_{K}, F_{J}$, $F_{x y}, C_{y}$ and $C_{x z}$.


Fig.2. The rotational constant A of the (100), (101), (102), (020), (021) and (022) bands. The dots which denotes the bands with the same $v_{1}$ and $v_{2}$ quantum number are linked by one line.

In our present work, the interaction parameters $F_{K}, F_{J}, F_{x y}, F_{K K}, C_{x z}$ and $C_{y}$ are fitted, while the parameter $F_{0}$ is not included in the fitting and is set to zero. If we compare the rotational constant A with those obtained by Papineau et al. ${ }^{[6]}$ and Ohshima et al. ${ }^{[7]}$, we can see that the change of the value of the rotational constant A does not obey the $\alpha$ - relations (Fig.2): the lines linking different dots which have the same $v_{1}$ and $v_{2}$ quantum number are not straight.

The reason is that the contact transformed Hamiltonian leads to a mixing of the rotational and coupling constants. The vib-rotational parameters got via this method are the effective values. Because different $F_{0}$ initial values mean different contact transformations of the Hamiltonian, we suggest the parameter $F_{0}$ should be considered in the global fitting rather than in this case.

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