# CRDS absorption spectrum of ${ }^{17} \mathrm{O}$ enriched water vapor in the $12,277-12,894 \mathrm{~cm}^{-1}$ range 

 Gu-Liang Liu ${ }^{\text {a }}$, Shui-Ming Hu ${ }^{\text {a,e }}$<br>${ }^{\text {a }}$ Hefei National Laboratory for Physical Sciences at Microscale, University of Science and Technology of China, Hefei 230026 China<br>${ }^{\mathrm{b}}$ Harvard-Smithsonian Center for Astrophysics, Atomic and Molecular Physics Division, Cambridge, MA, USA<br>${ }^{\text {c }}$ Laboratory of Theoretical Spectroscopy, V.E. Zuev Institute of Atmospheric Optics SB RAS, 1, Akademician Zuev square, Tomsk 634021, Russia<br>${ }^{\text {d Climate and Environmental Physics Laboratory, Ural Federal University, 19, Mira av.,Yekaterinburg 620002, Russia }}$<br>${ }^{e}$ CAS Center for Excellence in Quantum Information and Quantum Physics, University of Science and Technology of China, Hefei 230026, China

## A R T I C L E I N F O

## Article history:

Received 3 September 2018
Revised 3 October 2018
Accepted 4 October 2018
Available online 5 October 2018

## Keywords:

Cavity ring-down spectroscopy
Water vapor
$\mathrm{H}_{2}{ }^{17} \mathrm{O}$
$\mathrm{HD}^{17} \mathrm{O}$
Vibration-rotational assignment


#### Abstract

The absorption spectrum of water vapor enriched in oxygen-17 has been recorded by a continuous-wave cavity ring-down spectrometer (CW-CRDS) at room temperature in the $12,277-12,894 \mathrm{~cm}^{-1}$ region. The typical sensitivity of $2 \times 10^{-10} \mathrm{~cm}^{-1}$ allowed the detection of water vapor transitions with intensities larger than $1 \times 10^{-28} \mathrm{~cm}^{-1} /\left(\right.$ molecule $\left.\mathrm{cm}^{-2}\right)$. In total, more than 3400 lines were observed in the recorded spectra with a typical frequency accuracy of $0.002 \mathrm{~cm}^{-1}$. In which more than 3400 transitions were rovibrationally assigned to five water isotopic species $\left(\mathrm{H}_{2}{ }^{17} \mathrm{O}, \mathrm{H}_{2}{ }^{16} \mathrm{O}, \mathrm{H}_{2}{ }^{18} \mathrm{O}, \mathrm{HD}^{16} \mathrm{O}\right.$, and $\left.\mathrm{HD}^{17} \mathrm{O}\right)$ using the known empirical energy levels combined with the calculated variational line lists based on the results of Partridge and Schwenke. About 1300 transitions of $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ were assigned, leading to the determination of new energy levels of 17 vibration states. Those assigned $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions are involved with the upper vibration states of (013), (023), (032), (051), (061), (070), (080), (090), (112), (131), (150), (160), (211), (230), (301), (310), and (320). The maximum values of the rotational numbers $J$ and $K_{\mathrm{a}}$ were 16 and 9 , respectively. The $5 v_{1}$ band of $\mathrm{HD}^{17} \mathrm{O}$ was assigned for the first time as well. The line positions and energy levels determined in this work were compared with literature, and it shows a significant improvement of the line positions of $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ and $\mathrm{HD}^{17} \mathrm{O}$ in the $12,277-12,894 \mathrm{~cm}^{-1}$ region.


© 2018 Elsevier Ltd. All rights reserved.

## 1. Introduction

The present work is a subsequent part of the systematic study regarding the absorption spectrum of some important atmospheric molecules below $1 \mu \mathrm{~m}$ by continuous-wave cavity ring-down spectroscopy [1-8]. In this paper, we focus on the absorption of ${ }^{17} \mathrm{O}$ substituted water vapor between 12,277 and $12,894 \mathrm{~cm}^{-1}$. An IUPAC task group (IUPAC-TG) has performed an exhaustive review for the nine isotopologues of water [9-11]. The spectra of the main isotopic species $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ have been recorded by Fourier-transform spectroscopy (FTS) [12-16] and intra-cavity laser absorption spectroscopy (ICLAS) [17]. In total, about 1400 transitions of $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ in this region have been reported in the literature. Note that more

[^0]than 60 of them have been classified as "wrong" transitions by the IUPAC task group [9].

As for the isotopic species $\mathrm{H}_{2}{ }^{18} \mathrm{O}$, transitions in the $12,277-$ $12,894 \mathrm{~cm}^{-1}$ region were observed from the absorption spectra of ${ }^{18} \mathrm{O}$ enriched water vapor by FTS $[18,19]$ and ICLAS $[20,21]$. Some of those transitions were observed by Tolchenov et al. from FTS analysis of "natural" water vapor as well [14]. Overall, more than $820 \mathrm{H}_{2}{ }^{18} \mathrm{O}$ transitions have been reported by previous work in this region so far. The absorption spectrum of water vapor enriched by ${ }^{17} \mathrm{O}$ was reported by Tanaka et al. [19] in the region of interest. The ${ }^{17} \mathrm{O}$ enrichment was about $5 \%$ in their sample. Few additional $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ lines were reported in ICLAS studies of ${ }^{18} \mathrm{O}$ enriched samples $[20,21]$ and FTS spectrum of "natural" water samples [14] as well. In conclusion, More than $1810 \mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions have been obtained from various experimental measurements up to $12,740 \mathrm{~cm}^{-1}$ but most of them were reported in the $7443-7921 \mathrm{~cm}^{-1}$ region by Mondelain et al. [22]. Fig. 1 shows an overview spectrum of the $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions in the investigated region of this work. As shown in this figure, the $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ line positions from current HITRAN2016 in this region were mostly calcu-


Fig. 1. Overview of the $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions in the $12,277-12,894 \mathrm{~cm}^{-1}$ region for line intensities in the natural abundance at 296 K . The grey squares are corresponding to the calculations by Partridge and Schwenke [31,32] while the red circles are the current HIRTRAN2016 [24]. The blue triangles represent the experimental observations in this work, and black cross for the FTS results from Ref. [19].
lated from the IUPAC energy levels $[10,23,24]$. The $\mathrm{HD}^{16} \mathrm{O}$ absorption spectra between 12,200 and $12,900 \mathrm{~cm}^{-1}$ were studied by different techniques: photo-acoustic [25], Fourier-transform intracavity laser absorption spectroscopy [26], ICLAS [27,28], and FTS [29]. More than 900 transitions have been reported in these studies. As for the $\mathrm{HD}^{17} \mathrm{O}$ species, there are no experimental results for transitions above $12,000 \mathrm{~cm}^{-1}$ to the best of our knowledge.

In the present work, we report the spectra of ${ }^{17} \mathrm{O}$ enriched water vapor absorption using CW-CRDS covering the spectral region from 12,277 to $12,894 \mathrm{~cm}^{-1}$. The description of the experimental set up will be presented in Section 2 and the rotation-vibrational analysis of the recorded spectra will be presented in Section 3. In final conclusion, the obtained results for line positions related to the isotopic species $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ and $\mathrm{HD}^{17} \mathrm{O}$ will be provided and compared to the energy levels from literature.

## 2. Experimental details and line list preparation

The detail configuration of the continuous-wave cavity ringdown spectrometer based on a Ti:Sapphire laser has been described in Refs. [1-5]. Briefly, the structure of the setup is as follows: a beam from a CW tunable Ti:Sapphire laser (Coherent MBR 110 ) is coupled into a $65.5-\mathrm{cm}$-long resonance cavity. The high reflectivity of $99.995 \%$ cavity mirrors (Layertec GmbH) are installed in a high-vacuum chamber and can be precisely adjusted by a set of step motors (New Focus) using a controller outside the chamber. The ring-down signal is detected by a photodiode and recorded by an analog-digital converter (ADLink PCI 9228) installed on a personal computer. A nonlinear least-square fitting program is applied to derive the ring-down time $\tau$. The sample absorption coefficient $\alpha$ at the frequency $v$ is obtained using the following equation:
$\alpha(v)=\frac{1}{c \tau(v)}-\frac{1}{c \tau_{0}}$
where $c$ is the speed of light and $\tau_{0}$ is the ring-down time of the empty cavity. The achieved sensitivity without averaging varies from 0.5 to $5 \times 10^{-10} \mathrm{~cm}^{-1}$ depending on the reflectivity of the mirrors.

The enriched ${ }^{17} \mathrm{O}$ water vapor sample with a pressure of 7.5 Torr was filled into the CRDS cell without further purification. The enriched ${ }^{17} \mathrm{O}$ water sample was purchased from Icon Services, and the stated isotopic concentration for the oxygen atom is $60.8 \%$ of


Fig. 2. (a)The overview of the recorded spectrum between 12,277 and $12,894 \mathrm{~cm}^{-1}$ in the upper panel; (b) and (c) an example of the $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ line assignments around $12,700 \mathrm{~cm}^{-1}$ as shown in the lower panel. The marks " 171 new" correspond to the lines due to transitions with previously unknown upper energy levels of $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions, while the regular $\mathrm{H}_{2}{ }^{16} \mathrm{O}, \mathrm{H}_{2}{ }^{18} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions are marked as " 161 ", " 181 " and " 171 "; (d) the residuals between the calculated and observed spectra in the range of $12,703-12,709 \mathrm{~cm}^{-1}$.
${ }^{17} \mathrm{O}$. The spectra were recorded at room temperature ( $296 \pm 1 \mathrm{~K}$ ). Since there was a continuous exchange between water molecules in the gas phase and those absorbed on the walls of the sample cell, the sample pressure was continuously monitored with a capacitance gauge (full range $100 \mathrm{kPa}, 0.5 \%$ accuracy) during the measurement. $\mathrm{H}_{2}{ }^{16} \mathrm{O}, \mathrm{H}_{2}{ }^{17} \mathrm{O}, \mathrm{H}_{2}{ }^{18} \mathrm{O}, \mathrm{HD}^{16} \mathrm{O}$ and $\mathrm{HD}^{17} \mathrm{O}$ water isotopologues were found in the sample with the assignment of the recorded spectrum. The concentrations of three main species were estimated and listed in Table 1 together with the respective natural abundances of water isotopologues. The spectrum calibration was based on the reading given by a calibrated lambda-meter (Burleigh WA-1500) with 60 MHz accuracy.

An overview of the CW-CRDS spectra is presented in Fig. 2(a). A small part of the recorded spectrum around $12,706 \mathrm{~cm}^{-1}$ is given in Fig. 2(b) and (c). These pictures illustrate the transition assignments and demonstrate the detective capability of transitions with $4-5$ orders of dynamic range on the absorption.

## 3. Spectrum analysis and the line assignments

The complete line list was obtained using the computer code SpectraPlot [30]. The high quality of the recorded spectrum allows the determination of more than 3400 absorption lines in the intensity range of $4 \times 10^{-29}$ to $1.5 \times 10^{-24} \mathrm{~cm}^{-1} /\left(\right.$ molecule $\left.\mathrm{cm}^{-2}\right)$. The line parameters were determined through the multi-line fit-

Table 1
Overview information of assigned transitions and isotopologue abundances of studied water sample

| Isotopologue | Number of transitions | Range $\left(\mathrm{cm}^{-1}\right)$ | Estimated abundance | Natural abundance [24] |
| :--- | :--- | :--- | :--- | :--- |
| $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ | 1590 | $12,278-12,891$ | 0.540 | 0.997317 |
| $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ | 1283 | $12277-12,893$ | 0.454 | $0.371884 \mathrm{E}-3$ |
| $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ | 411 | $12,277-12,794$ | 0.006 | $0.199983 \mathrm{E}-2$ |
| $\mathrm{HD}^{16} \mathrm{O}$ | 93 | $12,559-12,877$ |  | $0.310693 \mathrm{E}-3$ |
| $\mathrm{HD}^{17} \mathrm{O}$ | 51 | $12,570-12,853$ |  | $0.115853 \mathrm{E}-6$ |
| Unassigned | 191 (lines) | $12,279-12,850$ |  |  |
| Total | 3428 |  |  |  |

ting procedure provided by SpectraPlot. The Voigt profile was applied in the fitting procedure of each line. The Gaussian width was fixed at the Doppler width. The pressure-broadening line width as well as the line position and intensity were obtained by the fitting for each transition. For instance, Fig. 2(d) gives the residuals between the observed and simulated spectra around $12,706 \mathrm{~cm}^{-1}$.

In total, $262 \mathrm{H}_{2}{ }^{16} \mathrm{O}$ line positions from Ref. [12] were used for the calibration of the recorded spectrum which yields a standard deviation of $0.0013 \mathrm{~cm}^{-1}$. The absolute accuracy of the line position in this work is better than $2 \times 10^{-3} \mathrm{~cm}^{-1}$ for those not very weak and isolated lines.

The complete line list of all 3412 lines is given as Supplementary Material I to this paper. The list includes experimental lines parameters (position, intensity and line width) of all recorded lines. Empirical line positions (if possible) as well as variational line intensities are given for all assigned lines.

The transitions of five water isotopologues $\left(\mathrm{H}_{2}{ }^{17} \mathrm{O}, \mathrm{H}_{2}{ }^{16} \mathrm{O}\right.$, $\mathrm{H}_{2}{ }^{18} \mathrm{O}, \mathrm{HD}^{16} \mathrm{O}$, and $\mathrm{HD}^{17} \mathrm{O}$ ) were assigned from the recorded spectrum. General information about all assigned transitions is given in Table 1. The assignment procedure was based on the variational line lists of five isotopologues mentioned above calculated by Tashkun (Tomsk, Russia) using the results of Partridge and Schwenke $[31,32]$. These line lists are available online at the system "Spectroscopy of Atmospheric Gases", http://spectra.iao.ru [33]. During the first step of the so-called "trivial assignment", variational line positions of $\mathrm{H}_{2}{ }^{17} \mathrm{O}, \mathrm{H}_{2}{ }^{16} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ molecules were replaced by data calculated from previously reported empirical energy levels $[9,10]$. This allows us to assign the lines by directly comparing experimental line list with the obtained line transitions. In the next step, variational line positions and intensities were compared with the left unassigned lines after "trivial assignment". Table 2 contains band-by-band information about assigned lines of non-deuterated species. The second step of the line assignments was to compare the experimental line list with variational calculations [33] for molecules $\mathrm{H}_{2}{ }^{17} \mathrm{O}, \mathrm{H}_{2}{ }^{16} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{18} \mathrm{O}$. This procedure allows us to assign about 800 transitions of $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ corresponding to previously unknown upper energy levels.

The comparison of previously reported $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions to those assigned in this study is shown in Fig. 1. The line positions and intensities of the transitions reported in Refs. [14,19-21,24] are given by red open circles (current HITRAN database) and dark crosses (FTS data from Tanaka et al. [19]). As shown in Fig. 1, only a few transitions weaker than $1 \times 10^{-28} \mathrm{~cm}^{-1} /\left(\right.$ molecule $\left.\mathrm{cm}^{-2}\right)$ which is equivalent to $1 \times 10^{-25} \mathrm{~cm}^{-1} /\left(\right.$ molecule $\left.\mathrm{cm}^{-2}\right)$ in pure $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ were observed in the published work [14, 19-21].

The lines for isotopic species $\mathrm{HD}^{16} \mathrm{O}$ were identified by "trivial assignment" using the results of Campargue et al. [27]. Ninety-four transitions mainly of the $5 v_{1}$ band were observed between 12,559 and $12,877 \mathrm{~cm}^{-1}$. The line positions of assigned transitions are in very good agreement with the previous work [27].

Because a water sample with more than $45 \%{ }^{17} \mathrm{O}$ enrichment was used in the measurement in this work, fifty-one transitions of the strongest band $\left(5 v_{1}\right)$ with maximum values of rotational numbers $J=8$ and $K_{a}=4$ were assigned for the $\mathrm{HD}^{17} \mathrm{O}$ isotopologue in


Fig. 3. $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ line position differences between HITRAN2016 and new data obtained in this work.
this region. Note that this is the first assignment of $\mathrm{HD}^{17} \mathrm{O}$ absorption lines above $8710 \mathrm{~cm}^{-1}$.

In total, 3221 lines were assigned to 3428 transitions of five isotopic water species as mentioned above. The $5 v_{1}$ band of $\mathrm{HD}^{17} \mathrm{O}$ was observed for the first time. There were still about two hundred lines with the intensity values between $4 \times 10^{-29}$ and $4.5 \times 10^{-27} \mathrm{~cm} /$ molecule left unassigned as shown in Table 1. In this work, a total of 1792 ro-vibrational energy levels of five different water isotopologues were determined, in which 613 levels were new or corrected. The complete list of the energy levels is given in Supplementary Material II attached to this paper.

## 4. Results and discussions

## 4.1. $\mathrm{H}_{2}{ }^{16} \mathrm{O}$

About 1600 transitions of 16 cold and 4 hot bands were assigned in this work for $\mathrm{H}_{2}{ }^{16} \mathrm{O}$. In general, the comparison of the observed data against the HITRAN2016 line list [24] shows good agreement. The line position differences are below $0.005 \mathrm{~cm}^{-1}$ for 1157 transitions and less than $0.01 \mathrm{~cm}^{-1}$ for 1307 ones. While there were still about 160 transitions which showed a significant difference between our results and the HITRAN2016 database. The $\left|v^{\text {HIT16 }}-v^{\text {TW }}\right|$ values for these transitions are between $0.01 \mathrm{~cm}^{-1}$ and $0.79 \mathrm{~cm}^{-1}$. Note that there are two types of $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ line positions included in the HITRAN2016 database for the spectral range of $12,277-12,894 \mathrm{~cm}^{-1}$. The first type is the so-called "empirical line positions" ( 2840 of 4103 transitions). These values were calculated from empirical energy levels recommended by the IUPAC task group [9]. A general overview of the comparison between our results and HITRAN2016 "empirical" line positions is shown on the upper panel of Fig. 3 (red open circles). The average dif-

Table 2
Band by band statistic information of $\mathrm{H}_{2}{ }^{17} \mathrm{O}, \mathrm{H}_{2}{ }^{16} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ assigned transitions.

| Band | $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ |  |  | $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ |  |  | $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | NT | $J_{\text {max }} K_{\mathrm{a} \text { max }}$ | Region ( $\mathrm{cm}^{-1}$ ) | NT | $J_{\text {max }} K_{\mathrm{a} \text { max }}$ | Region ( $\mathrm{cm}^{-1}$ ) | NT | $J_{\text {max }} K_{\mathrm{a} \text { max }}$ | Region ( $\mathrm{cm}^{-1}$ ) |
| $2 v_{1}+v_{2}+v_{3}$ | 232 | 168 | 12,278-12,855 | 176 | 148 | 12,287-12,858 | 49 | 136 | 12,284-12,678 |
| $3 v_{1}+v_{2}$ | 156 | 139 | 12,279-12,868 | 171 | 139 | 12,284-12,863 | 31 | 106 | 12,279-12,667 |
| $3 \nu_{2}+2 \nu_{3}$ | 110 | 148 | 12,280-12,858 | 66 | 147 | 12,279-12,705 | 9 | 96 | 12,297-12,681 |
| $\nu_{1}+3 \nu_{2}+v_{3}$ | 80 | 128 | 12,283-12,882 | 41 | 127 | 12,281-12,759 | 7 | 95 | 12,307-12,489 |
| $\nu_{2}+3 \nu_{3}$ | 432 | 168 | 12,280-12,891 | 390 | 167 | 12,277-12,893 | 170 | 126 | 12,277-12,783 |
| $\nu_{1}+6 \nu_{2}$ | 14 | 121 | 12,287-12,716 | 6 | 122 | 12,317-12,719 |  |  |  |
| $\nu_{1}+\nu_{2}+2 \nu_{3}$ | 345 | 149 | 12,288-12,885 | 283 | 138 | 12,278-12,889 | 143 | 107 | 12,279-12,794 |
| $4 \nu_{1}+v_{2}$ |  |  |  | 1 | 62 | $12,325$ |  |  |  |
| $8 v_{2}$ | 12 | 124 | 12,298-12,823 | 7 | 103 | 12,309-12,630 | 1 | 102 | 12,284 |
| $3 v_{1}+v_{3}$ | 17 | 84 | 12,281-12,392 | 15 | 83 | 12,283-12,427 |  |  |  |
| $2 v_{1}+3 v_{2}$ | 57 | 129 | 12,283-12,817 | 43 | 128 | 12,294-12,852 |  |  |  |
| $4 \nu_{1}$ | 4 | 84 | 12,304-12,340 |  |  |  |  |  |  |
| $2 v_{2}+3 v_{3}-v_{2}$ | 50 | 95 | 12,311-12,659 | 30 | 83 | 12,317-12,613 |  |  |  |
| $2 \nu_{1}+2 \nu_{2}+\nu_{3}-\nu_{2}$ | 1 | 63 | 12,323 |  |  |  |  |  |  |
| $9 v_{2}$ | 6 | 121 | 12,324-12,891 | 3 | 50 | 12,408-12,631 |  |  |  |
| $7 \nu_{2}$ | 5 | 128 | 12,332-12,761 | 6 | 128 | 12,332-12,700 |  |  |  |
| $5 v_{2}+v_{3}$ | 18 | 128 | 12,334-12,733 | 9 | 128 | 12,333-12,796 |  |  |  |
| $3 v_{1}+2 v_{2}$ |  |  |  | 3 | 62 | 12,341-12,438 |  |  |  |
| $\nu_{1}+2 \nu_{2}+2 \nu_{3}-v_{2}$ | 14 | 65 | 12,284-12,614 |  |  |  |  |  |  |
| $\nu_{1}+5 \nu_{2}$ | 7 | 128 | 12,455-12,772 | 5 | 118 | 12,589-12,735 |  |  |  |
| $6 v_{2}+v_{3}$ | 7 | 92 | 12,551-12,860 | 7 | 82 | 12,536-12,762 | 1 | 81 | 12,336 |
| $v_{1}+3 \nu_{3}-v_{2}$ | 23 | 63 | 12,619-12,826 | 21 | 73 | 12,579-12,804 |  |  |  |
| Total | 1590 | 169 | 12,278-12,891 | 1283 | 149 | 12,277-12,893 | 411 | 127 | 12,277-12,783 |

NT - Ro-vibrational transition number of given band. $J_{\text {max }} K_{\mathrm{a}} \max$ - maximum value of rotational numbers $J$ and $K_{\mathrm{a}}$.
ference between our line positions ( $v^{\mathrm{TW}}$ ) and this part of the HITRAN2016 data ( $\nu^{\text {HIT16 }}$ (Empirical)) is $-0.00094 \mathrm{~cm}^{-1}$ with a root mean square (RMS) difference of $0.00048 \mathrm{~cm}^{-1}$. The biggest differences are corresponding to the transitions with upper levels (211) $12{ }_{66}$ ( $\nu^{\text {HIT16 }}$ (Empirical) $-\nu^{\text {TW }}=0.1743 ; 0.1716 ; 0.1714 \mathrm{~cm}^{-1}$ ). Note that the $\left(V_{1}, V_{2}, V_{3}\right) J K_{a} K_{c}$ is according to the domination fraction relative to the vibrational and rotational numbers. The second part of the HITRAN2016 line list came from the variational calculation by Bubukina et al. [34] (1263 transitions). The general overview of the differences between our line positions and the variational values of HITRAN2016 is shown on the lower panel of Fig. 3 (blue open triangles) as well. The average difference between our line positions ( $\nu^{\mathrm{TW}}$ ) and this part of the HITRAN2016 data ( $\nu^{\mathrm{HIT16}}($ Variational $)$ ) is $-0.021 \mathrm{~cm}^{-1}$ with the RMS difference of $0.017 \mathrm{~cm}^{-1}$. The biggest differences are corresponding to the (160) $2_{02-(000)} 2_{11}\left(v^{\mathrm{HIT16}}(\right.$ Variational $\left.)-v^{\mathrm{TW}}=-0.7942 \mathrm{~cm}^{-1}\right)$ and (090) $1_{01}-(000) 2_{12}\left(v^{\text {HIT16 }}(\right.$ Variational $\left.)-v^{\mathrm{TW}}=-0.7116 \mathrm{~cm}^{-1}\right)$ transitions. Note that 162 transitions with the discrepancies larger than $0.01 \mathrm{~cm}^{-1}$ were provided in the Supplementary Material III.

Some examples of significant discrepancies between $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ observed line positions and those of the HITRAN2016 database are shown in Fig. 4. The empirical line position of the (131) $113_{3} 8^{-10}$ 19 transitions in the HITRAN database is shifted by $-0.0204 \mathrm{~cm}^{-1}$ compared to the observed position at $12,375.8826 \mathrm{~cm}^{-1}$ (upper panel of Fig. 4). The second line observed at $12,377.3308 \mathrm{~cm}^{-1}$ corresponds to the (310) $66_{6}-7{ }_{35}$ transition. The variational line position in the HITRAN list is shifted by $-0.0410 \mathrm{~cm}^{-1}$ compared to their position in our recorded spectrum. The line intensities of these two lines are $1.24 \times 10^{-27}$ and $1.78 \times 10^{-27} \mathrm{~cm}^{-1} /($ molecule $\mathrm{cm}^{-2}$ ), respectively. As shown in the lower panel of Fig. 4, the line positions of two transitions in the (013) band from the HITRAN database between 12,715 and $12,716 \mathrm{~cm}^{-1}$ are considered to be incorrect. The HITRAN's positions of the $854^{-7} 53$ ( $12,714.9394 \mathrm{~cm}^{-1}$ ) and $85_{3}-75_{2}\left(12,715.8754 \mathrm{~cm}^{-1}\right)$ transitions are from variational calculation [34]. The corresponding lines are shifted by about $0.050 \mathrm{~cm}^{-1}$ comparing to the observation.

We also made the comparison of the energy levels determined from the $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ line list given in this work against those reported


Fig. 4. Examples of $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ line position deviated from HITRAN2016 around $12,376 \mathrm{~cm}^{-1}$ (left panel) and $12,715 \mathrm{~cm}^{-1}$ (right panel). Observed and calculated spectra are given with blue dotted lines and red full lines, respectively.
by IUPAC-TG [9] and Ref. [34]. The average difference between our energies and the data from literature [9] is $-0.0004 \mathrm{~cm}^{-1}$ with the RMS difference of $0.0006 \mathrm{~cm}^{-1}$. The maximum difference is $0.408 \mathrm{~cm}^{-1}$ corresponding to the (230) 954 energy level (see Table 3).

Based on the assigned $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ transitions, we are able to obtain 73 new rotational energy levels of twelve vibrational states. These energies are listed in Table 3. In addition, Table 3 contains 72 rotation-vibration energies which differ more than $0.010 \mathrm{~cm}^{-1}$ from the values published by IUPAC-TG [9] and Ref. [34].

## 4.2. $\mathrm{H}_{2}{ }^{18} \mathrm{O}$

Four hundred and eleven transitions of $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ were assigned in the recorded spectrum. In general, we found a very good agreement between our observations and the HITRAN2016 data. According to Ref. [24], the $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ transitions of the HITRAN2016 database were mainly from calculations using empirical energy levels of IUPAC-TG [10] with few updates by Kyuberis recently
[40] and the variational calculation of Bubukina et al. [34]. Average difference between 395 our line positions and those from HITRAN is $-0.0001 \mathrm{~cm}^{-1}$ with the RMS difference of $0.0002 \mathrm{~cm}^{-1}$. The differences $\left(\nu^{\mathrm{HIT16}}-\nu^{\mathrm{TW}}\right)$ are bigger than $0.0 \mathrm{~cm}^{-1}$ only for eleven transitions. The maximum difference is $0.0269 \mathrm{~cm}^{-1}$ for the (013) $10{ }_{110}{ }^{-9} 1_{19}$ transition. The line position of this transition ( $12,657.0674 \mathrm{~cm}^{-1}$ ) yields the upper energy level $E_{(013)} 101$ $10=13,573.3605 \mathrm{~cm}^{-1}$ instead of $13,573.38763 \mathrm{~cm}^{-1}$ in Ref. [24]. Note that the (013) $10{ }_{110}$ level was obtained in Ref. [10] from two transitions, and it has a poor accuracy (grade B+) according to the IUPAC-TG ranking. Only one new energy $E=13,534.0028 \mathrm{~cm}^{-1}$ of the (013) $7_{53}$ level has been derived from previously observed (013) $7_{53^{-6}}{ }_{52}$ transition at $12,653.9266 \mathrm{~cm}^{-1}$.

## 4.3. $\mathrm{H}_{2}{ }^{17} \mathrm{O}$

More than 1280 transitions of $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ were assigned in the recorded spectrum (see Tables 1 and 2). More than 1050 transitions were observed for the first time. Graphical comparison of new data against those published in Ref. [19] is shown on the lower panel of Fig. 3 (blue open circle). The line position differences are from $-0.062 \mathrm{~cm}^{-1}$ for the transitions of the (310) 725 upper energy level up to $0.410 \mathrm{~cm}^{-1}$ for the transitions of the (080) $5{ }_{3} 3$ upper energy level. Mean difference between 458 our line positions and the data of Ref. [24] is $0.0051 \mathrm{~cm}^{-1}$ with the root mean square difference of $0.0019 \mathrm{~cm}^{-1}$. It is interesting to note that only four experimental line positions of Ref. [19] are included in the HITRAN2016 database. In addition, the discrepancy between our line position of the (131) $7_{52}-6{ }_{33}$ transition and the value from Ref. [19] is over $0.05 \mathrm{~cm}^{-1}$. There are about five hundred transitions in the HITRAN2016 database from Lodi and Tennyson [23]. This last paper is a compilation of partly "empirical" line positions calculated from energy levels [10], partly from the variational calculation of Shirin et al. [35].

Three of four line positions labeled in HITRAN2016 as the data of Ref. [19] differ from our observations by more than $0.01 \mathrm{~cm}^{-1}$. These lines are shown in Fig. 5. The lines (112) $202^{-2} 11$ at $12,361.190 \mathrm{~cm}^{-1}$ and (131) $75^{-6} 33$ at $12,362.708 \mathrm{~cm}^{-1}$ (right panel) and (112) $32_{1-3} 12$ at $12,424.192 \mathrm{~cm}^{-1}$ (left panel) are shifted comparing to our results by $-0.028,-0.058$ and $-0.014 \mathrm{~cm}^{-1}$, respectively. Once again, all three line positions in HITRAN2016 labeled as observations from Ref. [19]. Two calculated lines located at $12,423.498$ and $12,423.674 \mathrm{~cm}^{-1}$ correspond to the (112) $8{ }_{54^{-7}} 43$ and (013) $6{ }_{16^{-6}} 15$ transitions from Ref. [23]. They are shifted by 0.020 and $0.033 \mathrm{~cm}^{-1}$, respectively.

Only 499 of 1283 assigned $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions can be derived using energy levels given by IUPAC-TG [10]. The differences between observed line positions and calculated values using the empirical energy levels [10] are within $0.005 \mathrm{~cm}^{-1}$ for 345 transitions. Maximum discrepancy of $0.062 \mathrm{~cm}^{-1}$ is found for the transitions of the (310) 725 upper level.


Fig. 5. Examples of $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ line position deviated from HITRAN2016 around $12,362 \mathrm{~cm}^{-1}$ (left panel) and $12,424 \mathrm{~cm}^{-1}$ (right panel). Observed and calculated spectra are given in blue dotted lines and red full lines, respectively.

Based on the assigned $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ transitions, we were able to obtain 467 new and corrected rotational energy levels of 17 vibrational states. These energies are listed in Table 4. As for $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ energies (see Table 3), Table 4 contains 121 rotation-vibration energies which differ by more than $0.010 \mathrm{~cm}^{-1}$ from the values published by IUPAC-TG [10] and Ref. [23]. Corrected energy levels are given in italic.

### 4.4. Unambiguous ro-vibrational assignments

The difficulties of unambiguous assignment of vibrationrotation energy levels of the water molecule have been discussed in numerous studies on the absorption spectrum analysis of water in the near infrared range (see, for example, Refs. [36-39]). Therefore, as given in Ref. [10], for instance, some levels of the (310) state were re-assigned to the (112) and (230) states (see Table 4 and table footnotes). Our labeling gives more consistent sets of the energy levels for each of these states. The energies increase smoothly with the increasing of the rotational numbers $J$ and $K_{a}$. Due to the small number of assigned energy levels for the (051), (150), (061), (160), (070), (080), and (090) states (see Table 4), the identification of these states should be considered as preliminary assignments. Their vibrational labels could be verified in future studies.

As mentioned before, the variational line list [33] based on the results of Partridge and Schwenke [31,32] was used in the spectrum assignment procedure. In some cases, two or more of the different frequency values given in this list were assigned to the same transition of highly excited upper levels. The reason is that the procedure of domination fraction relative to the vibrational and rotational numbers may not work properly

Table 3
Corrected and new energy levels of $\mathrm{H}_{2}{ }^{16} \mathrm{O}$.

| Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | $E$ | $\Delta E$ | Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | $E$ | $\Delta E$ |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| 013 | 770 | $13,900.4575$ | -0.0307 | 112 | 1074 | $14,444.0991$ | 0.1718 |
| 013 | 771 | $13,900.4574$ | -0.0310 | 112 | 1166 | $14,530.1983$ | 0.0104 |
| 013 | 863 | $13,924.0045$ | -0.0569 | 131 | 881 | $14,042.0550$ | 0.0271 |
| 013 | 871 | $14,090.6844$ | 0.0129 | 131 | 880 | $14,042.0597$ | 0.0317 |
| 013 | 872 | $14,090.6830$ | 0.0112 | 131 | 972 | $13,985.9298$ | -0.0763 |
| 013 | 880 | $14,275.3371$ | 0.0188 | 131 | 982 | $14,264.8177$ | 0.0195 |
| 013 | 881 | $14,275.3373$ | 0.0171 | 131 | 981 | $14,264.8482$ | 0.0498 |
| 013 | 954 | $13,990.7161$ | -0.0119 | 131 | 1138 | $13,668.9012$ | 0.0217 |
| 013 | 963 | $14,206.6816$ | 0.0154 | 131 | 1166 | $14,174.8601$ | 0.0308 |
|  |  |  |  |  |  |  | (continued on next page) |

Table 3 (continued)

| Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | $E$ | $\Delta E$ | Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | $E$ | $\Delta E$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 013 | 1028 | 13,959.0555 | 0.0103 | 131 | 12210 | 13,819.7335 | 0.0267 |
| 013 | 1038 | 13,959.9420 | -0.0199 | 160 | 202 | 12,451.1602 | 0.0370 |
| 013 | 1056 | 14,222.7949 | 0.0139 | 211 | 1064 | 14,004.8612 | -0.0682 |
| 013 | 1148 | 14,341.1244 | -0.0761 | 211 | 1074 | 14,135.1942 | -0.0156 |
| 013 | 1166 | 14,634.9649 | 0.0142 | 211 | 1165 | 14,264.5049 | 0.0116 |
| 013 | 12112 | 14,042.6584 | -0.0159 | 211 | 1266 | 14,454.5671 | -0.1719 |
| 013 | 12211 | 14,266.0266 | 0.0162 | 211 | 14311 | 14,800.1394 | 0.0146 |
| 013 | 13013 | 14,275.2042 | 0.0261 | 211 | 16115 | 14,964.3214 | 0.0194 |
| 013 | 14114 | 14,527.7918 | 0.0370 | 230 | 651 | 12,856.8443 | -0.0169 |
| 023 | 532 | 14,587.1213 | 0.0110 | 230 | 744 | 12,826.4054 | 0.0134 |
| 023 | 734 | 14,916.4463 | -0.0108 | 230 | 862 | 13,439.8219 | 0.3723 |
| 032 | 936 | 13,343.3215 | 0.0388 | 230 | 863 | 13,439.8022 | 0.3418 |
| 032 | 945 | 13,459.9914 | 0.0171 | 230 | 954 | 13,428.6727 | -0.4082 |
| 032 | 972 | 14,075.1198 | 0.0220 | 230 | 964 | 13,653.6622 | -0.0174 |
| 032 | 973 | 14,075.1161 | 0.0193 | 230 | 990 | 14,368.8595 | 0.0144 |
| 032 | 13211 | 14,301.2304 | -0.0348 | 230 | 991 | 14,368.8597 | 0.0146 |
| 051 | 973 | 13,800.9975 | 0.0101 | 230 | 1047 | 13,462.7160 | -0.0186 |
| 061 | 000 | 12,586.2299 | -0.0251 | 230 | 1129 | 13,499.3326 | 0.0455 |
| 061 | 514 | 13,065.2492 | 0.0241 | 301 | 422 | 14,123.5240 | -0.0182 |
| 061 | 918 | 13,819.5725 | -0.0385 | 310 | 660 | 13,194.0261 | -0.0629 |
| 070 | 1147 | 13,778.6369 | 0.0111 | 310 | 1065 | 13,999.5313 | -0.0128 |
| 080 | 532 | 12,877.7365 | 0.1482 | 310 | 1091 | 14,789.6155 | 0.0124 |
| 080 | 927 | 13,074.2001 | 0.3581 | 310 | 1092 | 14,789.6158 | 0.0131 |
| 080 | 936 | 13,587.8400 | 0.0375 | 310 | 1129 | 13,773.8357 | 0.1634 |
| 090 | 505 | 12,880.1436 | -0.1288 | 310 | 1138 | 13,896.2910 | -0.0139 |
| 112 | 1019 | 13,647.2372 | 0.0157 | 310 | 1148 | 13,929.0993 | 0.0115 |
| 112 | 1073 | 14,444.0778 | 0.1509 | 310 | 1174 | 14,704.2776 | 0.0114 |
| 013 | 1047 | 14,083.0981 | new | 131 | 1175 | 14,488.1742 | new |
| 013 | 1065 | 14,374.8268 | new | 131 | 1267 | 14,544.3937 | new |
| 013 | 1074 | 14,542.2188 | new | 150 | 972 | 13,657.5607 | new |
| 013 | 1073 | 14,542.2560 | new | 150 | 1073 | 14,227.2902 | new |
| 013 | 1249 | 14,618.1022 | new | 150 | 1175 | 14,490.6470 | new |
| 013 | 13410 | 14,911.1967 | new | 150 | 1267 | 14,269.9478 | new |
| 013 | 14212 | 15,021.7203 | new | 211 | 1083 | 14,384.2162 | new |
| 013 | 14312 | 15,023.8963 | new | 211 | 1174 | 14,395.2518 | new |
| 013 | 16016 | 15,082.5187 | new | 211 | 1248 | 14,289.0323 | new |
| 032 | 753 | 13,225.2783 | new | 211 | 13310 | 14,488.8723 | new |
| 032 | 1055 | 14,083.0999 | new | 211 | 1349 | 14,604.4466 | new |
| 032 | 1074 | 14,314.9461 | new | 230 | 1055 | 13,668.9738 | new |
| 032 | 1083 | 14,668.3462 | new | 230 | 1065 | 13,890.6045 | new |
| 032 | 1082 | 14,668.3415 | new | 230 | 1083 | 14,429.8106 | new |
| 032 | 1139 | 13,752.9477 | new | 230 | 1082 | 14,429.8059 | new |
| 032 | 1156 | 14,397.8638 | new | 230 | 1092 | 14,608.2577 | new |
| 032 | 13311 | 14,306.5068 | new | 230 | 1091 | 14,608.2581 | new |
| 051 | 1055 | 13,458.4487 | new | 230 | 1148 | 13,716.8708 | new |
| 051 | 1184 | 14,646.6417 | new | 230 | 1156 | 13,932.0577 | new |
| 051 | 1183 | 14,646.7035 | new | 230 | 1174 | 14,398.5027 | new |
| 070 | 1248 | 14,065.1252 | new | 230 | 1248 | 14,050.2814 | new |
| 080 | 1038 | 13,801.0016 | new | 230 | 1257 | 14,221.7456 | new |
| 080 | 1249 | 14,822.0585 | new | 230 | 1267 | 14,539.7546 | new |
| 090 | 101 | 12,404.4517 | new | 310 | 1028 | 13,531.6531 | new |
| 090 | 515 | 13,337.6642 | new | 310 | 1037 | 13,629.5498 | new |
| 090 | 12112 | 14,287.8202 | new | 310 | 1046 | 13,718.3963 | new |
| 112 | 991 | 14,426.3376 | new | 310 | 1074 | 14,137.5711 | new |
| 112 | 990 | 14,426.3381 | new | 310 | 1175 | 14,704.3657 | new |
| 112 | 1148 | 14,195.1471 | new | 310 | 1184 | 14,586.5261 | new |
| 112 | 1147 | 14,246.4807 | new | 310 | 1183 | 14,586.5306 | new |
| 112 | 1157 | 14,343.6825 | new | 310 | 1193 | 15,053.2043 | new |
| 112 | 1156 | 14,371.4599 | new | 310 | 1192 | 15,053.2036 | new |
| 112 | 1248 | 14,530.7872 | new | 310 | 1276 | 14,987.3372 | new |
| 112 | 1257 | 14,656.3786 | new | 310 | 13211 | 14,318.1639 | new |
| 131 | 881 | 14,042.0550 | new | 310 | 13310 | 14,460.7870 | new |
| 131 | 1156 | 13,963.0297 | new | 310 | 1367 | 15,115.9713 | new |
| 131 | 1165 | 14,170.4470 | new |  |  |  |  |

Notes: Vib $J K_{\mathrm{a}} K_{\mathrm{c}}$ - vibration and rotation assignment of the energy level; $E$ - energy values obtained in this study $\left(\mathrm{cm}^{-1}\right) ; \Delta E$ - energy difference between this work and the literature values from IUPAC [9] and Ref. [34] ( $\mathrm{cm}^{-1}$ ).
for strongly perturbed levels. As a consequence, two (or more) strongly interacting energy levels are attributed to the same quantum numbers. For instance, two sets of frequencies and intensities ( $12,689.1548 \mathrm{~cm}^{-1}, 2.712 \times 10^{-27} \mathrm{~cm}^{-1} /\left(\right.$ molecule $\left.\mathrm{cm}^{-2}\right)$ and $12,693.8815 \mathrm{~cm}^{-1}, 2.994 \times 10^{-27} \mathrm{~cm}^{-1} /\left(\right.$ moleculecm $\left.{ }^{-2}\right)$ ) are as-
signed for the transition (013) $121_{12}$ - (000) $111_{11}$ in the calculation [33]. In this case, one of them was assigned to the transition mentioned above, and the rest needs to be re-assigned. It also means that the second transition remains uncertain in the vibrational numbers ( $V_{1}, V_{2}, V_{3}$ ) and possibly for the rotational numbers
$K_{a}$ and $K_{c}$ as well. It is similar for the transition (013) $12{ }_{111}-(000)$ $111_{10}$ - 12,716.2038 cm ${ }^{-1}, 2.931 \times 10^{-27} \mathrm{~cm}^{-1} /\left(\right.$ moleculecm $\left.{ }^{-2}\right)$ and $12,719.2580 \mathrm{~cm}^{-1}, 2.747 \times 10^{-27} \mathrm{~cm}^{-1} /\left(\right.$ moleculecm $\left.{ }^{-2}\right)$. As can be seen from Table 4, this problem is not resolved yet for 45 levels.

Please notice here the problem of ambiguous identification of the variational lists is not a specific problem of the SP lists of Ref.
[33]. This is a common issue for variational calculations of polyatomic molecules. Regarding the list of water molecules, the predictions from UCL (including recent lists [40]) prefer not to give the complete vibrational and rotational numbers for many transitions. Since then, 174 of $493 \mathrm{H}_{2}{ }^{17} \mathrm{O}$ HITRAN2016 transitions came from Ref. [23] still remain incomplete vibration-rotation assignment.

Table 4
New and corrected energy levels of $\mathrm{H}_{2}{ }^{17} \mathrm{O}$.

| Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | Term value | Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | Term value | Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | Term value |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 013 | 432 | 12,910.1443 | 051 | 880 | 14,005.2276 | 211 | 964 | 13,741.6260 |
| 013 | 532 | 13,035.5408 | 051 | 982 | 14,111.8630 | 211 | 963 | 13,741.7825 |
| 013 | 541 | 13,131.6372 | 051 | 981 | 14,111.8979 | 211 | 973 | 13,868.4890 |
| 013 | 550 | 13,250.3432 | 051 | 1055 | 13,429.0726 | 211 | 972 | 13,868.5439 |
| 013 | 616 | 12,965.6374 | 051 | 1266 | 14,268.9936 | 211 | 1047 | 13,666.1685 |
| 013 | 625 | 13,075.0392 | 061 | 220 | 12,895.9398 | 211 | 1046 | 13,703.5819 |
| 013 | 643 | 13,276.4175 | 061 | 514 | 13,038.6745 | 211 | 1056 | 13,809.0832 |
| 013 | 642 | 13,277.5624 | 061 | 615 | 13,193.9789 | 211 | 1055 | 13,815.2518 |
| 013 | 652 | 13,392.3102 | 061 | 716 | 13,372.1051 | 211 | 1065 | 13,977.3995 |
| 013 | 651 | 13,392.3734 | 061 | 817 | 13,572.1936 | 211 | 1064 | 13,977.8366 |
| 013 | 661 | 13,538.1966 | 070 | 881 | 13,934.0333 | 211 | 1074 | 14,105.6197 |
| 013 | 660 | 13,538.1989 | 070 | 880 | 13,934.0335 | 211 | 1073 | 14,105.6362 |
| 013 | 707 | 13,097.7199 | 070 | 982 | 14,152.8366 | 211 | 11110 | 13,587.3321 |
| 013 | 716 | 13,220.7965 | 070 | 981 | 14,152.8365 | 211 | 1129 | 13,759.9917 |
| 013 | 734 | 13,365.1943 | 070 | 1266 | 14,023.7719 | 211 | 1139 | 13,764.7181 |
| 013 | 744 | 13,444.6627 | 080 | 432 | 12,732.3341 | 211 | 1138 | 13,881.6545 |
| 013 | 743 | 13,448.5002 | 080 | 533 | 12,850.7892 | 211 | 1147 | 13,975.2421 |
| 013 | 753 | 13,558.1037 | 080 | 734 | 13,155.6245 | 211 | 1157 | 14,065.7315 |
| 013 | 752 | 13,558.4318 | 080 | 1029 | 13,385.7607 | 211 | 1156 | 14,079.3573 |
| 013 | 762 | 13,704.8293 | 090 | 505 | 12,854.6668 | 211 | 1166 | 14,236.1432 |
| 013 | 761 | 13,704.8269 | 103 | 303 | 14,424.4111 | 211 | 1165 | 14,237.3018 |
| 013 | 771 | 13,869.7096 | 112 | 000 | 12,389.0982 | 211 | 12111 | 13,824.0075 |
| 013 | 770 | 13,869.7098 | 112 | 111 | 12,425.3523 | 211 | 12211 | 13,823.9559 |
| 013 | 827 | 13,391.1882 | 112 | 202 | 12,456.1885 | 211 | 1239 | 14,159.8274 |
| 013 | 826 | 13,498.4842 | 112 | 212 | 12,465.4581 | 211 | 1258 | 14,343.7127 |
| 013 | 836 | 13,517.6054 | 112 | 211 | 12,482.1374 | 211 | 1257 | 14,369.6720 |
| 013 | 845 | 13,635.8799 | 112 | 220 | 12,523.7783 | 211 | 1266 | 14,520.2780 |
| 013 | 844 | 13,645.9383 | 112 | 321 | 12,597.3163 | 211 | 1276 | 14,647.6944 |
| 013 | 854 | 13,747.5917 | 112 | 413 | 12,656.4227 | 211 | 13211 | 14,292.5280 |
| 013 | 853 | 13,748.7996 | 112 | 431 | 12,766.3773 ${ }^{\text {a }}$ | 211 | 13311 | 14,291.7728 |
| 013 | 863 | 13,894.3227 | 112 | 441 | 12,871.8545 | 211 | 13310 | 14,465.3710 |
| 013 | 862 | 13,894.3506 | 112 | 440 | 12,871.8553 | 211 | 1349 | 14,584.6735 |
| 013 | 872 | 14,059.8794 | 112 | 515 | 12,698.5162 | 211 | 14311 | 14,774.9798 |
| 013 | 871 | 14,059.8793 | 112 | 524 | 12,791.8746 | 211 | 14410 | 14,915.1957 |
| 013 | 909 | 13,412.9724 | 112 | 533 | 12,881.9022 | 230 | 550 | 12,689.6183 |
| 013 | 918 | 13,578.3877 | 112 | 551 | 13,120.0065 | 230 | 642 | 12,637.0821 |
| 013 | 927 | 13,707.7747 | 112 | 550 | 13,119.8881 | 230 | 651 | 12,832.2488 |
| 013 | 937 | 13,723.7340 | 112 | 616 | 12,811.9659 | 230 | 661 | 13,054.9417 |
| 013 | 936 | 13,796.1451 | 112 | 625 | 12,922.4329 | 230 | 660 | 13,054.9380 |
| 013 | 946 | 13,849.6136 | 112 | 624 | 12,976.6054 | 230 | 743 | 13,297.6905 ${ }^{\text {a }}$ |
| 013 | 945 | 13,871.3425 | 112 | 634 | 13,021.6023 | 230 | 753 | 12,998.7548 |
| 013 | 955 | 13,960.5051 | 112 | 633 | 13,035.2924 | 230 | 752 | 12,998.9286 |
| 013 | 954 | 13,963.9208 | 112 | 643 | 13,130.4071 | 230 | 762 | 13,221.9353 |
| 013 | 964 | 14,108.4659 | 112 | 651 | 13,259.7445 | 230 | 761 | 13,221.9400 |
| 013 | 963 | 14,108.7150 | 112 | 661 | 13,425.0044 | 230 | 771 | 13,543.7650 |
| 013 | 973 | 14,273.7493 | 112 | 660 | 13,424.9994 | 230 | 770 | 13,543.7652 |
| 013 | 972 | 14,273.7594 | 112 | 725 | 13,149.4985 ${ }^{\text {a }}$ | 230 | 854 | 13,188.7802 |
| 013 | 10110 | 13,596.4001 | 112 | 734 | 13,209.5239 ${ }^{\text {a }}$ | 230 | 853 | 13,189.4620 |
| 013 | 1019 | 13,780.6601 | 112 | 744 | 13,295.0249 | 230 | 863 | 13,412.5491 |
| 013 | 1029 | 13,780.9301 | 112 | 753 | 13,425.1928 | 230 | 862 | 13,412.5741 |
| 013 | 1028 | 13,932.6650 | 112 | 752 | 13,423.7712 | 230 | 872 | 13,737.2494 |
| 013 | 1038 | 13,934.4728 | 112 | 762 | 13,592.2851 | 230 | 871 | 13,737.2448 |
| 013 | 1037 | 14,042.7902 | 112 | 761 | 13,592.2634 | 230 | 881 | 14,075.0817 |
| 013 | 1047 | 14,085.7348 | 112 | 771 | 13,773.2519 | 230 | 880 | 14,075.0819 |
| 013 | 1046 | 14,122.6118 | 112 | 770 | 13,773.2565 | 230 | 954 | 13,404.0162 |
| 013 | 1056 | 14,196.4166 | 112 | 826 | 13,073.6352 | 230 | 963 | 13,626.3939 |
| 013 | 1055 | 14,204.1651 | 112 | 836 | 13,099.6240 ${ }^{\text {a }}$ | 230 | 972 | 13,954.2475 |
| 013 | 1065 | 14,345.5452 | 112 | 835 | 13,142.5327 | 230 | 1028 | 13,222.1291 |
| 013 | 1064 | 14,346.4382 | 112 | 845 | 13,221.7778 | 230 | 1047 | 13,438.6147 |
| 013 | 1073 | 14,510.9924 | 112 | 844 | 13,229.5617 | 230 | 1056 | 13,638.0210 |
| 013 | 11011 | 13,797.0816 | 112 | 854 | 13,351.7746 | 230 | 1055 | 13,644.7059 |
| 013 | 11111 | 13,796.7472 | 112 | 853 | 13,355.2241 | 230 | 1065 | 13,972.6397 |
| 013 | 11110 | 14,001.1183 | 112 | 862 | 13,524.6673 | 230 | 1074 | 14,194.3666 |

Table 4 (continued)

| Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | Term value | Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | Term value | Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | Term value |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 013 | 11210 | 14,001.2429 | 112 | 872 | 13,654.9494 | 230 | 1073 | 14,194.3979 |
| 013 | 1129 | 14,173.3546 | 112 | 871 | 13,654.9448 | 230 | 1083 | 14,394.4022 |
| 013 | 1139 | 14,174.7398 | 112 | 881 | 13,846.3029 | 230 | 1129 | 13,478.4030 |
| 013 | 1138 | 14,306.4902 | 112 | 880 | 13,846.3031 | 230 | 1138 | 13,596.8133 |
| 013 | 1148 | 14,316.0702 | 112 | 927 | 13,285.0776 | 230 | 1148 | 13,693.0144 |
| 013 | 1147 | 14,398.6690 | 112 | 945 | 13,449.0326 | 230 | 1156 | 13,908.6228 |
| 013 | 1156 | 14,501.8885 | 112 | 955 | 13,573.5747 | 230 | 1257 | 14,197.3815 |
| 013 | 12012 | 14,014.4246 | 112 | 954 | 13,567.8762 | 310 | 431 | 12,501.1754 |
| 013 | 12112 | 14,017.7220 | 112 | 963 | 13,738.6028 | 310 | 523 | 12,557.0953 |
| 013 | 12111 | 14,237.4463 | 112 | 973 | 13,869.2758 | 310 | 542 | 12,726.5996 |
| 013 | 12211 | 14,238.7048 | 112 | 972 | 13,869.7983 | 310 | 624 | 12,707.9009 |
| 013 | 12210 | 14,430.6443 | 112 | 982 | 14,060.1030 | 310 | 633 | 12,768.5133 |
| 013 | 12310 | 14,431.0632 | 112 | 981 | 14,060.1029 | 310 | 642 | 12,868.9496 |
| 013 | 1239 | 14,585.6370 | 112 | 1019 | 13,358.1987 | 310 | 651 | 13,001.9964 |
| 013 | 1248 | 14,696.2065 | 112 | 1028 | 13,511.4812 | 310 | 735 | 12,916.3286 |
| 013 | 1257 | 14,793.8583 | 112 | 1038 | 13,517.8719 | 310 | 734 | 12,942.6388 |
| 013 | 13113 | 14,247.9783 | 112 | 1047 | 13,662.0118 | 310 | 744 | 13,032.9191 |
| 013 | 13112 | 14,492.2750 | 112 | 1046 | 13,700.1557 | 310 | 743 | 13,036.2446 |
| 013 | 13212 | 14,493.6606 | 112 | 1056 | 13,805.2033 | 310 | 762 | 13,334.4951 |
| 013 | 13311 | 14,704.1405 | 112 | 1055 | 13,805.2109 | 310 | 761 | 13,334.1878 |
| 013 | 14014 | 14,499.9287 | 112 | 1064 | 13,978.2343 | 310 | 771 | 13,463.4491 |
| 013 | 14114 | 14,499.9659 | 112 | 1073 | 14,107.4283 | 310 | 770 | 13,463.4493 |
| 013 | 14113 | 14,759.2145 | 112 | 1083 | 14,296.6101 | 310 | 818 | 13,089.5121 |
| 013 | 14213 | 14,761.0111 | 112 | 1082 | 14,296.6052 | 310 | 817 | 13,236.8084 |
| 013 | 15115 | 14,768.3668 | 112 | 1129 | 13,755.0959 | 310 | 827 | 13,240.0704 |
| 013 | 16016 | 15,053.5327 | 112 | 1139 | 13,746.1355 | 310 | 826 | 13,342.3347 |
| 023 | 101 | 14,062.5460 | 112 | 1148 | 13,905.5702 | 310 | 836 | 13,366.9024 |
| 023 | 111 | 14,077.5312 | 112 | 1147 | 14,222.1738 | 310 | 835 | 13,414.4384 |
| 023 | 202 | 14,107.4037 | 112 | 1174 | 14,367.6347 | 310 | 845 | 13,482.4090 |
| 023 | 211 | 14,135.9807 | 112 | 1183 | 14,555.5513 | 310 | 844 | 13,489.9219 |
| 023 | 221 | 14,180.2653 | 112 | 12211 | 13,812.0145 | 310 | 854 | 13,614.4032 |
| 023 | 220 | 14,181.7524 | 112 | 12210 | 14,012.9422 | 310 | 853 | 13,611.7311 |
| 023 | 303 | 14,171.4394 | 112 | 12310 | 14,005.4680 | 310 | 863 | 13,783.2448 |
| 023 | 312 | 14,213.3712 | 112 | 1249 | 14,176.2383 | 310 | 862 | 13,783.1209 |
| 023 | 322 | 14,250.0033 | 112 | 1258 | 14,337.7970 | 310 | 872 | 13,963.9670 |
| 023 | 330 | 14,337.7281 | 112 | 13410 | 14,453.8447 | 310 | 871 | 13,963.9624 |
| 023 | 413 | 14,314.0216 | 131 | 541 | 12,521.9155 | 310 | 881 | 14,175.5434 |
| 023 | 422 | 14,359.4724 | 131 | 550 | 12,711.0583 | 310 | 880 | 14,175.5436 |
| 023 | 515 | 14,352.3014 | 131 | 642 | 12,665.2303 | 310 | 919 | 13,252.9844 |
| 023 | 514 | 14,435.4494 | 131 | 652 | 12,855.2955 | 310 | 918 | 13,422.5818 |
| 023 | 524 | 14,454.2088 | 131 | 651 | 12,855.1645 | 310 | 928 | 13,431.1750 |
| 023 | 533 | 14,553.5952 | 131 | 661 | 13,067.7143 | 310 | 927 | 13,552.1868 |
| 023 | 606 | 14,464.8871 | 131 | 660 | 13,067.7209 | 310 | 937 | 13,557.4993 |
| 023 | 615 | 14,575.0068 | 131 | 743 | 12,832.9951 | 310 | 936 | 13,638.4798 |
| 023 | 624 | 14,643.3381 | 131 | 753 | 13,023.4806 | 310 | 946 | 13,691.7157 |
| 023 | 633 | 14,709.5615 | 131 | 752 | 13,022.7544 | 310 | 945 | 13,705.7840 |
| 023 | 717 | 14,596.3898 | 131 | 762 | 13,235.8002 | 310 | 955 | 13,827.1263 |
| 023 | 726 | 14,742.3701 | 131 | 761 | 13,235.7962 | 310 | 954 | 13,823.5385 |
| 023 | 725 | 14,820.1109 | 131 | 845 | 13,023.2156 | 310 | 964 | 13,998.0490 |
| 023 | 808 | 14,744.3129 | 131 | 853 | 13,212.9771 | 310 | 963 | 13,997.5102 |
| 023 | 818 | 14,744.5861 | 131 | 863 | 13,427.4009 | 310 | 973 | 14,177.9134 |
| 032 | 541 | 12,700.6302 | 131 | 862 | 13,427.3664 | 310 | 972 | 14,177.9079 |
| 032 | 551 | 12,884.5978 | 131 | 872 | 13,738.1580 | 310 | 982 | 14,393.2515 |
| 032 | 550 | 12,884.2926 | 131 | 871 | 13,738.1619 | 310 | 981 | 14,393.2514 |
| 032 | 633 | 12,705.0766 | 131 | 927 | 13,029.9664 | 310 | 991 | 14,514.1374 |
| 032 | 643 | 12,843.7340 | 131 | 946 | 13,236.3497 | 310 | 990 | 14,514.1374 |
| 032 | 642 | 12,844.3443 | 131 | 955 | 13,421.7749 | 310 | 10010 | 13,436.5247 |
| 032 | 652 | 13,029.0329 | 131 | 954 | 13,422.2579 | 310 | 10110 | 13,436.8854 |
| 032 | 651 | 13,028.5674 | 131 | 964 | 13,642.7339 | 310 | 1019 | 13,625.5693 |
| 032 | 661 | 13,235.5217 | 131 | 1028 | 13,266.9745 | 310 | 1029 | 13,623.7259 |
| 032 | 660 | 13,235.5180 | 131 | 1047 | 13,471.5419 | 310 | 1028 | 13,777.9427 |
| 032 | 725 | 12,793.9487 | 131 | 1065 | 13,879.5412 | 310 | 1038 | 13,779.6750 |
| 032 | 744 | 13,010.8320 | 131 | 1139 | 13,535.2669 | 310 | 1037 | 13,886.1255 |
| 032 | 752 | 13,196.8493 | 131 | 1138 | 13,645.1719 | 310 | 1047 | 13,922.0005 |
| 032 | 762 | 13,404.4220 | 131 | 1148 | 13,727.8601 | 310 | 1046 | 13,959.7291 |
| 032 | 761 | 13,404.4398 | 131 | 1147 | 13,776.5673 | 310 | 1056 | 14,062.4649 |
| 032 | 771 | 13,632.2362 | 131 | 1157 | 13,920.6223 | 310 | 1055 | 14,058.8641 |
| 032 | 770 | 13,632.2386 | 131 | 1175 | 14,455.6488 | 310 | 1065 | 14,237.1919 |

Table 4 (continued)

| Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | Term value | Vib | $J K_{\mathrm{a}} K_{\mathrm{c}}$ | Term value | Vib | $J K_{\mathrm{a}} K_{\text {c }}$ | Term value |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 032 | 826 | 13,002.6834 | 131 | 1257 | 14,221.7103 | 310 | 1074 | 14,414.8877 |
| 032 | 836 | 13,048.5739 | 131 | 1267 | 14,517.4420 | 310 | 1073 | 14,414.8699 |
| 032 | 845 | 13,201.2690 | 131 | 1266 | 14,425.8768 | 310 | 1083 | 14,634.2104 |
| 032 | 844 | 13,216.2475 | 150 | 982 | 14,294.5736 | 310 | 1082 | 14,634.2055 |
| 032 | 853 | 13,389.0166 | 150 | 981 | 14,294.5735 | 310 | 1092 | 14,755.7875 |
| 032 | 863 | 13,596.9427 | 150 | 1083 | 14,536.6212 | 310 | 1091 | 14,755.7875 |
| 032 | 862 | 13,596.9579 | 150 | 1082 | 14,536.6163 | 310 | 11110 | 13,837.6346 |
| 032 | 872 | 13,826.6696 | 150 | 1174 | 14,457.3056 | 310 | 11210 | 13,842.6968 |
| 032 | 871 | 13,826.6684 | 160 | 404 | 12,593.0429 | 310 | 1129 | 14,019.9026 |
| 032 | 936 | 13,318.0300 | 160 | 505 | 12,710.7098 | 310 | 1139 | 14,018.3035 |
| 032 | 946 | 13,412.0763 | 160 | 616 | 12,934.8835 | 310 | 1138 | 14,151.4816 |
| 032 | 954 | 13,605.2537 | 160 | 11011 | 13,798.0690 | 310 | 1148 | 14,172.3122 |
| 032 | 963 | 13,813.3222 | 160 | 12211 | 14,240.5007 | 310 | 1147 | 14,245.7451 |
| 032 | 972 | 14,043.5606 | 211 | 532 | 12,630.5888 | 310 | 1156 | 14,345.5455 |
| 032 | 1038 | 13,479.2969 | 211 | 845 | 13,225.7343 | 310 | 1165 | 14,496.2902 |
| 032 | 1047 | 13,672.8565 | 211 | 854 | 13,362.9598 | 310 | 1174 | 14,674.6264 |
| 032 | 1046 | 13,681.1811 | 211 | 863 | 13,529.1005 | 310 | 12211 | 14,079.1811 |
| 032 | 1074 | 14,283.1184 | 211 | 872 | 13,654.2506 | 310 | 12310 | 14,275.7107 |
| 032 | 1138 | 13,852.5076 | 211 | 871 | 13,654.2505 | 310 | 1257 | 14,631.3305 |
| 032 | 1148 | 13,919.4432 | 211 | 881 | 13,895.6291 | 310 | 1266 | 14,780.1871 |
| 032 | 1147 | 13,955.6084 | 211 | 880 | 13,895.6374 | 310 | 13112 | 14,332.2084 |
| 032 | 12310 | 13,989.3529 | 211 | 936 | 13,376.2645 | 310 | 13310 | 14,438.1281 |
| 032 | 1249 | 14,198.0829 | 211 | 945 | 13,456.5804 | 320 | 321 | 14,110.6224 |
| 032 | 14312 | 14,585.8217 | 211 | 946 | 13,435.9009 | 320 | 423 | 14,194.9757 |
| 051 | 881 | 14,005.2230 | 211 | 954 | 13,575.7981 |  |  |  |

Notes: Vib $J K_{\mathrm{a}} K_{\mathrm{c}}$ - vibration and rotation assignment of the energy level; term value - energy level obtained in this study $\left(\mathrm{cm}^{-1}\right)$.
Energy term values corrected compare to corresponding value given by IUPAC-TG (10) and Ref. [23] are given in italic.
${ }^{\text {a }}$ In Ref. [10], this levels assigned to the (310) vibration state (see text).

## 5. Conclusion

Absorption spectrum of ${ }^{17} \mathrm{O}$ enriched water vapor was recorded and analyzed between 12,277 and $12,894 \mathrm{~cm}^{-1}$. More than 3400 transitions, corresponding to five water isotopologues $\left(\mathrm{H}_{2}{ }^{17} \mathrm{O}\right.$, $\mathrm{H}_{2}{ }^{16} \mathrm{O}, \mathrm{H}_{2}{ }^{18} \mathrm{O}, \mathrm{HD}^{16} \mathrm{O}$, and $\mathrm{HD}^{17} \mathrm{O}$ ), were observed in this work. About 3000 of them were assigned using the calculations of Partridge and Schwenke [31,32] in combination with the known empirical energy levels [9-11]. In which more than 1100 transitions of $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ as well as 51 lines of the $5 v_{1}$ band of $\mathrm{HD}^{17} \mathrm{O}$ were observed for the first time. Quite a few disagreements between recorded spectrum and the HITRAN2016 data for the $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ molecules were observed.

About 420 highly excited vibration-rotation energy levels were newly obtained through this work which could make significant improvements of 193 energies reported by IUPAC-TG [9-11] and Refs. $[23,34]$ in the studied region. Rotational energies of the (023) and (230) states of $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ have been assigned for the first time.

The reported data could be useful for the future improvements of the spectroscopic databases HITRAN [24] and GEISA [41] as well as for the updates of the energy level sets [9-11].

## Acknowledgments

This work was supported in part by NSFC (21473172, 91436209 and 21688102) and RFBR (Russia, Grant nos. 14-05-91150 and 16-55-16017). SNM gratefully acknowledge the University of Science and Technology of China for visiting professorship. SNM activity was also partly supported by Ural Federal University in the frame of Act 211, contract no. 02.A03.21.0006. The authors are grateful to Dr. A.V. Nikitin (Tomsk, Russia) whose computer code SpectraPlot was used for the line list preparation and to determine the line parameters. We are also grateful to Dr. O.V. Naumenko (Tomsk, Russia) for helpful discussions on the vibrational labeling of water lines.

## Supplementary materials

Supplementary material associated with this article can be found, in the online version, at doi:10.1016/j.jqsrt.2018.10.009.

## References

[1] Tan Y, Wang J, Zhao X-Q, Liu A-W, Hu S-M. Cavity ring-down spectroscopy of the fifth overtone of CO. J Quant Spectrosc Radiat Transf 2017;187:274-9. doi:10.1016/J.JQSRT.2016.10.003.
[2] Tan Y, Zhao X-Q Liu A-W, Hu S-M, Lyulin OM, Tashkun SA, et al. Cavity ringdown spectroscopy of $\mathrm{CO}_{2}$ overtone bands near 830 nm . J Quant Spectrosc Radiat Transf 2015;165:22-7. doi:10.1016/J.JQSRT.2015.06.010.
[3] Tan Y, Wang J, Cheng C-F, Zhao X-Q, Liu A-W, Hu S-M. Cavity ring-down spectroscopy of the electric quadrupole transitions of $\mathrm{H}_{2}$ in the $784-852 \mathrm{~nm}$ region. J Mol Spectrosc 2014;300:60-4. doi:10.1016/J.JMS.2014.03.010.
[4] Liu A-W, Li X-F, Wang J, Lu Y, Cheng C-F, Sun YR, et al. The $4 v$ CH overtone of ${ }^{12} \mathrm{C}_{2} \mathrm{H}_{2}$ : sub-MHz precision spectrum reveals perturbations. J Chem Phys 2013;138:014312. doi:10.1063/1.4773473.
[5] Hu S-M, Pan H, Cheng C-F, Sun YR, Li X-F, Wang J, et al. The $v=3 \leftarrow 0$ S (0)-S (3) electric quadrupole transitions of $\mathrm{H}_{2}$ near $0.8 \mu \mathrm{~m}$. Astrophys J 2012;749:76. doi:10.1088/0004-637X/749/1/76.
[6] Lu Y, Liu A-W, Pan H, Li X-F, Perevalov VI, Tashkun SA, et al. High sensitivity cavity ring down spectroscopy of ${ }^{13} \mathrm{C}^{16} \mathrm{O}_{2}$ overtone bands near 806 nm . J Quant Spectrosc Radiat Transf 2012;113:2197-204. doi:10.1016/J.JQSRT.2012.07.020.
[7] Song K-F, Lu Y, Tan Y, Gao B, Liu A-W, Hu S-M. High sensitivity cavity ring down spectroscopy of $\mathrm{CO}_{2}$ overtone bands near 790 nm . J Quant Spectrosc Radiat Transf 2011;112:761-8. doi:10.1016/J.JQSRT.2010.11.006.
[8] Song K-F, Gao B, Liu A-W, Perevalov VI, Tashkun SA, Hu S-M. Cavity ring-down spectroscopy of the $6 \nu_{3}$ bands of ${ }^{15} \mathrm{~N}$ substituted $\mathrm{N}_{2} \mathrm{O}$. J Quant Spectrosc Radiat Transf 2010;111:2370-81. doi:10.1016/J.JQSRT.2010.05.022.
[9] Tennyson J, Bernath PF, Brown LR, Campargue A, Császár AG, Daumont L, et al. IUPAC critical evaluation of the rotational-vibrational spectra of water vapor. Part III: energy levels and transition wavenumbers for $\mathrm{H}_{2}{ }^{16} \mathrm{O}$. J Quant Spectrosc Radiat Transf 2013;117:29-58. doi:10.1016/j.jqsrt.2012.10.002.
[10] Tennyson J, Bernath PF, Brown LR, Campargue A, Carleer MR, Császár AG, et al. IUPAC critical evaluation of the rotational-vibrational spectra of water vapor. Part I: energy levels and transition wavenumbers for $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{18} \mathrm{O}$. J Quant Spectrosc Radiat Transf 2009;110:573-96. doi:10.1016/j.jqsrt.2009.02.014.
[11] Bernath PF, Brown LR, Campargue A, Császár AG, Daumont L, Gamache RR, et al. IUPAC critical evaluation of the rotational-vibrational spectra of water vapor. Part II: energy levels and transition wavenumbers for $\mathrm{HD}^{16} \mathrm{O}, \mathrm{HD}^{17} \mathrm{O}$, and $\mathrm{HD}^{18} \mathrm{O}$. J Quant Spectrosc Radiat Transf 2010;111:2160-84. doi:10.1016/J. JQSRT.2010.06.012.
[12] Toth RA. Measurements of $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ line positions and strengths: 11,610 to 12,861 $\mathrm{cm}^{-1}$. J Mol Spectrosc 1994;166:176-83. doi:10.1006/JMSP.1994.1183.
[13] Flaud J-M, Camy-Peyret C, Bykov A, Naumenko O, Petrova T, Scherbakov A, et al. The high-resolution spectrum of water vapor between 11,600 and 12,750 $\mathrm{cm}^{-1}$. J Mol Spectrosc 1997;183:300-9. doi:10.1006/JMSP.1997.7275.
[14] Tolchenov RN, Naumenko O, Zobov NF, Shirin SV, Polyansky OL, Tennyson J, et al. Water vapour line assignments in the $9250-26,000 \mathrm{~cm}^{-1}$ frequency range. J Mol Spectrosc 2005;233:68-76. doi:10.1016/J.JMS.2005.05.015.
[15] Tolchenov R, Tennyson J. Water line parameters from refitted spectra constrained by empirical upper state levels: study of the $9500-14,500 \mathrm{~cm}^{-1}$ region. J Quant Spectrosc Radiat Transf 2008;109:559-68. doi:10.1016/J.JQSRT. 2007.08.001.
[16] Tolchenov RN, Tennyson J, Brault JW, Canas AAD, Schermaul R. Weak line water vapor spectrum in the $11,787-13,554 \mathrm{~cm}^{-1}$ region. J Mol Spectrosc 2002;215:269-74. doi:10.1006/JMSP.2002.8653.
[17] Campargue A, Mikhailenko S, Liu AW. ICLAS of water in the 770 nm transparency window ( $12,746-13,558 \mathrm{~cm}^{-1}$ ). Comparison with current experimental and calculated databases. J Quant Spectrosc Radiat Transf 2008;109:2832-45. doi:10.1016/J.JQSRT.2008.07.003.
[18] Bykov A, Naumenko O, Petrova T, Scherbakov A, Sinitsa L, Mandin JY, et al. The second decade of $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ : line positions and energy levels. J Mol Spectrosc 1995;172:243-53. doi:10.1006/JMSP.1995.1172.
[19] Tanaka M, Naumenko O, Brault JW, Tennyson J. Fourier transform absorption spectra of $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ in the $3 v+\delta$ and $4 v$ polyad region. J Mol Spectrosc 2005;234:1-9. doi:10.1016/j.jms.2005.07.007.
[20] Mazzotti F, Tolchenov RN, Campargue A. High sensitivity ICLAS of $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ in the region of the second decade (11,520-12,810 $\mathrm{cm}^{-1}$ ). J Mol Spectrosc 2007;243:78-89. doi:10.1016/JJJMS.2007.02.020.
[21] Leshchishina OM, Naumenko OV. Campargue a. high sensitivity ICLAS of $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ in the $12,580-13,550 \mathrm{~cm}^{-1}$ transparency window. J Quant Spectrosc Radiat Transf 2011;112:913-24. doi:10.1016/J.JQSRT.2010.11.012.
[22] Mondelain D, Mikhailenko SN, Karlovets EV, Béguier S, Kassi S, Campargue A. Comb-assisted cavity ring down spectroscopy of ${ }^{17} \mathrm{O}$ enriched water between 7443 and $7921 \mathrm{~cm}^{-1}$. J Quant Spectrosc Radiat Transf 2017;203:206-12. doi:10. 1016/j.jqsrt.2017.03.029.
[23] Lodi L, Tennyson J. Line lists for $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{17} \mathrm{O}$ based on empirical line positions and ab initio intensities. J Quant Spectrosc Radiat Transf 2012;113:850-8. doi:10.1016/J.JQSRT.2012.02.023.
[24] Gordon IE, Rothman LS, Hill C, Kochanov RV, Tan Y, Bernath PF, et al. The HITRAN2016 molecular spectroscopic database. J Quant Spectrosc Radiat Transf 2017;203:3-69. doi:10.1016/j.jqstt.2017.06.038.
[25] Lazarev VV, Petrova TM, Sinitsa LN, Zhu Q Han J, Hao L. Absorption spectrum of $\mathrm{HD}^{16} \mathrm{O}$ in $0.7 \mu \mathrm{~m}$ region. Atmos Oceanic Opt 1998;11:809-12. http://ao.iao. ru/en/content/vol.11-1998/iss.09/4.
[26] Hu S, Lin H, He S, Cheng J, Zhu Q. Fourier-transform intra-cavity laser absorption spectroscopy of HOD $\nu_{O D}=5$ overtone. Phys Chem Chem Phys 1999;1:3727-30. doi:10.1039/a903593a.
[27] Campargue A, Vasilenko I, Naumenko O. Intracavity laser absorption spectroscopy of HDO between 11,645 and $12,330 \mathrm{~cm}^{-1}$. J Mol Spectrosc 2005;234:216-27. doi:10.1016/JJMS.2005.09.007.
[28] Naumenko OV, Voronin BA, Mazzotti F, Tennyson J, Campargue A. Intracavity laser absorption spectroscopy of HDO between 12,145 and $13,160 \mathrm{~cm}^{-1}$. J Mol Spectrosc 2008;248:122-33. doi:10.1016/JJMS.2007.12.005.
[29] Voronin BA, Naumenko OV, Carleer M, Coheur P-F, Fally S, Jenouvrier A, et al. HDO absorption spectrum above $11,500 \mathrm{~cm}^{-1}$ : assignment and dynamics. J Mol Spectrosc 2007;244:87-101. doi:10.1016/J.JMS.2007.03.008.
[30] Nikitin AV, Kochanov RV. Visualization and identification of spectra by the spectraplot program. Atmos Oceanic Opt 2011;24:936-41. http://ao.iao.ru/en/ content/vol.24-2011/iss.11/2.
[31] Partridge H, Schwenke DW. The determination of an accurate isotope dependent potential energy surface for water from extensive $a b$ initio calculations and experimental data. J Chem Phys 1997;106:4618-39. doi:10.1063/1.473987.
[32] Schwenke DW, Partridge H. Convergence testing of the analytic representation of an ab initio dipole moment function for water: Improved fitting yields improved intensities. J Chem Phys 2000;113:6592. doi:10.1063/1.1311392.
[33] Mikhailenko SN, Babikov YuL, Golovko VF. Information-calculating system spectroscopy of atmospheric gases. The structure and main functions. Atmos Oceanic Opt 2005;18:685-95.
[34] Bubukina II, Zobov NF, Polyansky OL, Shirin SV, Yurchenko SN. Optimized semiempirical potential energy surface for $\mathrm{H}_{2}{ }^{16} \mathrm{O}$ up to $26,000 \mathrm{~cm}^{-1}$. Opt Spectrosc 2011;110:160-6. http://ao.iao.ru/en/content/vol.18-2005/iss.09/4. doi:10. 1134/S0030400X11020032.
[35] Shirin SV, Zobov NF, Ovsyannikov RI, Polyansky OL, Tennyson J. Water line lists close to experimental accuracy using a spectroscopically determined potential energy surface for $\mathrm{H}_{2}{ }^{16} \mathrm{O}, \mathrm{H}_{2}{ }^{17} \mathrm{O}$, and $\mathrm{H}_{2}{ }^{18} \mathrm{O}$. J Chem Phys 2008;128:224306. doi:10.1063/1.2927903.
[36] Campargue A, Leshchishina OM, Naumenko OV. $\mathrm{D}_{2}{ }^{16} \mathrm{O}$ : ICLAS between 13,600 and $14,020 \mathrm{~cm}^{-1}$ and normal mode labeling of the vibrational states. J Mol Spectrosc 2009;254:1-9. doi:10.1016/JJMSS.2008.12.002.
[37] Mikhailenko S, Kassi S, Wang L, Campargue A. The absorption spectrum of water in the $1.25 \mu \mathrm{~m}$ transparency window ( $7408-7920 \mathrm{~cm}^{-1}$ ). J Mol Spectrosc 2011;269:92-103. doi:10.1016/JJMS.2011.05.005.
[38] Lavrentieva NN, Voronin BA, Naumenko OV, Bykov AD, Fedorova AA. Linelist of $\mathrm{HD}^{16} \mathrm{O}$ for study of atmosphere of terrestrial planets (Earth, Venus and Mars). Icarus 2014;236:38-47. doi:10.1016/J.ICARUS.2014.03.037.
[39] Regalia L, Oudot C, Mikhailenko S, Wang L, Thomas X, Jenouvrier A, et al. Water vapor line parameters from 6450 to $9400 \mathrm{~cm}^{-1}$. J Quant Spectrosc Radiat Transf 2014;136:119-36. doi:10.1016/J.JQSRT.2013.11.019.
[40] Polyansky OL, Kyuberis AA, Lodi L, Tennyson J, Yurchenko SN, Ovsyannikov RI, Zobov NF. ExoMol molecular line lists XIX: high-accuracy computed line lists for $\mathrm{H}_{2}{ }^{18} \mathrm{O}$ and $\mathrm{H}_{2}{ }^{17} \mathrm{O}$. Mon Not R Astron Soc 2017;466:1363-71. doi:10.1093/ mnras/stw3125.
[41] Jacquinet-Husson N, Armante R, Scott NA, Chédin A, Crépeau L, Boutammine C, et al. The 2015 edition of the GEISA spectroscopic database. J Mol Spectrosc 2016;327:31-72. doi:10.1016/J.JMS.2016.06.007.


[^0]:    * Corresponding author.

    E-mail addresses: ty7788@mail.ustc.edu.cn (Y. Tan), semen@iao.ru (S.N. Mikhailenko), jinwang@ustc.edu.cn (J. Wang), awliu@ustc.edu.cn (A.-W. Liu), smhu@ustc.edu.cn (X.-Q. Zhao), xiaoqinz@mail.ustc.edu.cn (G.-L. Liu), lgl@mail.ustc.edu.cn (S.-M. Hu).

