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# Rovibrational analysis of the absorption spectrum of HDO between 10 110 and 12 215 cm<sup>-1</sup> †

### O. Naumenko,<sup>ab</sup> S.-M. Hu,<sup>c</sup> S.-G. He<sup>c</sup> and A. Campargue<sup>a</sup>

<sup>a</sup> Laboratoire de Spectrométrie Physique (associated with CNRS, UMR 5588), Université Joseph Fourier de Grenoble, B.P. 87, 38402 Saint-Martin-d'Hères Cedex, France

<sup>b</sup> Institute of Atmospheric Optics, Russian Academy of Science, Tomsk, Russia

<sup>c</sup> Laboratory of Bond-Selective Chemistry, University of Science and Technology of China, Hefei 230026, China

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The weak absorption spectrum of monodeuterated water has been investigated between 10 110 and 12 215 cm<sup>-1</sup> by high resolution Fourier transform absorption spectroscopy with a 105 m absorption pathlength. The spectrum is dominated by the  $4\nu_1$ ,  $3\nu_3$ , and  $\nu_2 + 3\nu_3$  bands at 10 378.95, 10 631.68 and 11 969.76 cm<sup>-1</sup>, respectively. The rovibrational assignment was performed on the basis of the *ab initio* calculations of Schwenke and Partridge and by using the effective Hamiltonian approach. 502 energy levels belonging to a total of 13 vibrational states were determined. The rotational structure of the (003) state is mostly isolated and could be fitted with an rms of 0.006 cm<sup>-1</sup>, slightly larger than the experimental uncertainty. Most of the other levels are connected *via* a complex interaction scheme involving dark states which prevented a satisfactory modeling. The analysis of the rotational structure of the (013), (400) and (112) states is investigated for the first time, while the existing information for the (003) state is considerably enlarged and improved. Transitions reaching the nine other (dark) states, including the highly excited bending states, (051), (061) and (080), were detected in the spectrum as a result of intensity borrowing due to resonance interactions with bright states.

### I. Introduction

The knowledge of the rovibrational energy levels of monodeuterated water is much less complete than that of the main isotopomer,  $H_2^{16}O$ , particularly in terms of line intensities. However, numerous new experimental data have been obtained on the rotational structure of the HDO vibrational states both by high resolution Fourier transform (FT) absorption spectroscopy<sup>1-6</sup> and intracavity laser absorption spectroscopy (ICLAS).<sup>7-14</sup> In addition, great progress has been recently achieved in the calculation of water vapor absorption spectra from high quality *ab initio* potential and dipole moment surfaces.<sup>15,16</sup>

The present contribution is devoted to the analysis of the HDO absorption spectrum investigated between 10110 and 12215 cm<sup>-1</sup> by Fourier transform absorption spectroscopy with a 105 m absorption pathlength (see Fig. 1). The 9625–10100 cm<sup>-1</sup> region, which is just below the region of interest, was very recently analyzed by ICLAS based on vertical external cavity surface emitting lasers (VeCSEL).<sup>13</sup> The high sensitivity of ICLAS, allowed a significant improvement over a previous study performed in that region on the basis of the same FT spectrum presently investigated.<sup>5</sup>

As labeled in Fig. 1, the absorption spectrum of HDO in the considered region, is dominated by the  $4\nu_1$ ,  $3\nu_3$  and  $\nu_2 + 3\nu_3$  bands at 10 378.95, 10 631.68 and 11 969.76 cm<sup>-1</sup> respectively (we use the traditional labeling of the vibrational modes:  $\nu_1(2723.68 \text{ cm}^{-1})$ ,  $\nu_2(1403.48 \text{ cm}^{-1})$  and  $\nu_3(3707.47 \text{ cm}^{-1})$  for the OD stretching, the bending and the OH stretching vibration, respectively). The  $3\nu_3$  band was previously investigated between 10 280 and 10 770 cm<sup>-1</sup> by ICLAS based

† Electronic supplementary information (ESI) available: HDO absorption spectrum and rotational assignments. See http:// www.rsc.org/suppdata/cp/b3/b312514a/ on a  $F_2^+$ :LiF color center laser.<sup>14</sup> The analysis of the spectrum (317 lines) resulted in the derivation and modeling (rms = 0.028 cm<sup>-1</sup>) of 97 observed energy levels of the (003) state. However, the quality of the FT spectrum presently analyzed allowed us to improve this analysis both in term of sensitivity and in term of accuracy. For completeness, we should mention the analysis of the low  $JK_aK_c$  levels of the (003) state



**Fig. 1** Overview of the HDO spectrum between 9500 and 12500 cm<sup>-1</sup>. Upper panel: stick spectrum retrieved from the FT spectrum recorded with a 105 m absorption pathlength at a total pressure of water of 15 hPa. Lower panel: stick spectrum calculated by Schwenke and Partridge (SP 2000) from *ab initio* potential energy and dipole moment surfaces.<sup>16</sup> The normal mode labeling of the upper vibrational state of the dominant transitions is given.

observed by photoacoustic spectroscopy using an injection seeded optical parametric oscillator  $(OPO)^{17}$  which allowed the transition moment alignment in HOD to be studied.<sup>18</sup> To the best of our knowledge, the analysis of the  $4\nu_1$  and  $\nu_2 + 3\nu_3$  bands has not been previously reported.

### **II.** Experimental details

The sample of D<sub>2</sub>O was purchased from Peking Chemical Industry, Ltd (China). The stated deuterium purity is 99.8%. The spectra obtained in Hefei (China) were recorded at room temperature with a Bruker IFS 120 HR Fourier Transform interferometer equipped with a pathlength adjustable multipass cell, a tungsten light source, a CaF<sub>2</sub> beam splitter, and a Si-diode detector. The unapodized resolution was 0.02  $cm^{-1}$  and the absorption pathlength was 105 m. Since in the region under study many lines are due to D<sub>2</sub>O and H<sub>2</sub>O, two samples were prepared in order to discriminate the different isotopomers from the relative line intensities, as illustrated in Fig. 2. The first one was a 1:1 mixture of H<sub>2</sub>O and D<sub>2</sub>O at a pressure of 15 hPa which is assumed to lead to a mixture of H<sub>2</sub>O:HDO:D<sub>2</sub>O in the proportion 1:2:1. In the second one, the D<sub>2</sub>O concentration was much higher. The FT spectrum was calibrated with water lines listed in the HITRAN database. The accuracy of the absolute positions of unblended lines is estimated to be around  $0.002 \text{ cm}^{-1}$ .



Fig. 2 Comparison of a section of the FT spectra obtained with different proportion of water isotopomers (P = 15 hPa, l = 105 m). The lower panel (*b*) corresponds to a mixture of H<sub>2</sub>O:HDO:D<sub>2</sub>O in the proportion 1:2:1 while the upper panel (*a*) was recorded with a large amount of deuterated water. The variation of the relative intensities helps to discriminate the HDO lines: compared to HDO lines, the H<sub>2</sub>O lines are much stronger in (*b*) than in (*a*) while it is the opposite for D<sub>2</sub>O lines. The lines due to D<sub>2</sub>O and H<sub>2</sub>O are indicated.

### III. Spectrum assignment and energy level determination

Our recent high sensitivity ICLAS investigation covered the  $9625-10\,100$  cm<sup>-1</sup> region.<sup>13</sup> We thus limit the present analysis of the FT spectrum to the region just above 10100 cm<sup>-</sup> About 1800 absorption lines, observed between 10110 and 12215 cm<sup>-1</sup>, were attributed to HDO, some of them being blended by H<sub>2</sub>O and D<sub>2</sub>O lines. For identification purposes, approximate intensities were derived for all observed lines from their peak absorption. The spectral assignment was performed relying on the Schwenke and Partridge (SP) database,<sup>15,16</sup> which predicts accurate line positions and intensities for H<sub>2</sub>O and its isotopomers. A computer code developed for automatic rovibrational spectrum assignment was also intensively exploited.<sup>19</sup> This program used the SP synthetic spectrum and experimental line positions and intensities as input data. Of all HDO lines, 1579 were rovibrationally assigned which, taking into account the unresolved multiplets, corresponds to 1827 transitions reaching 13 upper vibrational states.

The experimental information obtained is summarized in Table 1 while the resulting line list, which includes the experimental line positions and SP calculated intensities followed by the rovibrational assignment, is available as electronic supplementary information (ESI).† Note that, as shown in Fig. 1, the analyzed spectral region consists in two well isolated parts,  $10110-11452 \text{ cm}^{-1}$  and  $11750-12215 \text{ cm}^{-1}$ , corresponding mainly to the  $3\nu_3$  and  $\nu_2 + 3\nu_3$  bands, respectively. In the region in between, the HDO absorption lines were too weak to be reliably identified.

Accurate upper energy levels for all observed transitions were derived by adding the ground state rotational term values.<sup>20</sup> The total number of energy levels obtained for every observed vibrational state is presented in Table 1 while their values are listed in Tables 2 and 3, which also include their experimental uncertainty and the deviations from SP predictions. In order to check and confirm SP rovibrational assignment, we attempted to reproduce the observed energy levels with the effective Hamiltonian (EH) approach (see next section). However, in most cases we did not succeed in performing a complete EH fitting due to the lack of sufficient experimental information since the three analyzed "bright" states, (400), (013) and (112), were strongly perturbed by many "dark' states (see next Section). However, the EH modeling, though being not fully satisfactory, allowed the correction of some SP assignments and added to the understanding of the resonance scheme involved in the considered polyads of interacting states. The only successful fit was achieved for the (003) state, which appears to be mostly isolated from the other vibrational states.

Tables 2 and 3 show that the deviations of SP prediction from the observed energy levels can reach a value of 0.43  $cm^{-1}$  for some vibrational states. However, as we have already stressed,<sup>13</sup> these deviations vary regularly with J and  $K_a$ . This feature combined with a reasonable agreement between calculated<sup>16</sup> and experimental intensity values and supported by the systematic use of the ground state combination difference (GSCD) method, provided a complete and reliable assignment of the observed spectrum. Fig. 1 and Fig. 3 show the overall excellent agreement between the observed and calculated<sup>16</sup> spectrum in the whole region and an expanded section, respectively. As a result, 502 energy levels were derived and assigned to 13 vibrational states. The only previous reports relevant to these states are (i) an ICLAS study of the (003) state near 10 500 cm<sup>-1</sup>, which led to the derivation of 97 energy levels;<sup>14</sup> (ii) the photoacoustic spectroscopy of the same (003) state from which 13 low  $JK_aK_c$  levels were derived,<sup>17</sup> and (*iii*) our ICLAS-VeCSEL study in the 9625–10 100 cm<sup>-1</sup> region<sup>13</sup> which allowed us to determine 35 and 11 energy levels from hot bands transitions, for the (032) and (112) states, respectively.

Table 1 Summary of the HDO experimental information obtained in the 10110–12215 cm<sup>-1</sup> spectral region

$E_v/\mathrm{cm}^{-1}$					
SP <sup>15,16</sup>	Observed	$v_1 v_2 v_3$	Integrated intensity <sup>a</sup> /cm <sup>-2</sup> atm <sup>-1</sup>	Number of transitions	Number of levels
10119.37		(080)	$3.45 \times 10^{-5}$	7	2
10 318.55		(051)	$2.21 \times 10^{-4}$	25	12
10 378.64	10378.9511	(400)	$3.58 \times 10^{-3}$	231	89
10403.21	10403.1422	(211)	$1.55 \times 10^{-4}$	34	19
10480.73		(131)	$1.40 \times 10^{-4}$	8	2
10 631.92	10631.6832	(003)	$2.78  imes 10^{-1}$	983	172
10 652.52		(320)	$1.00 \times 10^{-3}$	20	9
11 243.06	11 242.923	(032)	$4.75 \times 10^{-6}$	4	4
11 315.43	11 315.4333	(112)	$3.83  imes 10^{-4}$	81	53
11 533.48		$(061)^{b}$	$4.90  imes 10^{-6}$	3	1
11 804.57		(141)	$4.50 \times 10^{-5}$	10	7
11958.26		$(330)^{b}$	$4.65 \times 10^{-4}$	17	9
11 970.07	11969.7584	(013)	$1.90 \times 10^{-2}$	320	123

 $^{a}$  The integrated intensities were obtained from the sum of the calculated  $^{16}$  pure HDO intensities of all assigned transitions for a given vibrational band.  $^{b}$  Integrated intensities for the (061)–(000) and (330)–(000) bands seem to be distorted (see text).



Fig. 3 The FT absorption spectrum of HDO in the 10703–10720 cm<sup>-1</sup> spectral region corresponding to a section of the  $3\nu_3$  band centered at 10631.683 cm<sup>-1</sup>. The  $J'K'_aK'_c - J''K''_aK''_c$  rotational assignments are given for some of the lines (all of the assignments are listed in the ESI†). The rotational assignments written in bold characters are those of extra lines borrowing their intensities from transitions of the  $3\nu_3$  band (see text). The upper panel is a stick spectrum retrieved from the FT spectrum recorded with a 105 m absorption pathlength at a total pressure of water of 15 hPa. The lower panel is the HDO spectrum calculated by Schwenke and Partridge from *ab initio* potential energy and dipole moment surfaces.<sup>16</sup>

For completeness, we include in Table 3 the energy levels of the (032) state obtained by ICLAS-VeCSEL.<sup>13</sup> Compared to the extensive set of levels obtained for the (102)–(022) interacting dyad,<sup>13</sup> five additional levels, listed in Table 3, could be determined. It is worth noticing that in accordance with the wave function mixing obtained in the fitting of the (112)–(032) dyad, we inverted some vibrational assignments of the (032) and (112) states with respect to those in ref. 13. The values of ref. 14 for the energy levels of the (003) state show deviations of 0.07 cm<sup>-1</sup> on average from our more precise values, with a maximum discrepancy of 0.14 cm<sup>-1</sup>. Five energy levels were incorrect. Thus, the total number of newly derived energy levels is 404, while the determination of the 92 previously determined energy levels of the (003) state is significantly improved.

## IV. Effective Hamiltonian and resonance interactions with dark states

Numerous strong resonance interactions between the analyzed vibrational states were evidenced during the spectral

assignment and modeling process. In many cases, these interactions induce important intensity transfer to weak transitions, which would have been unobservable in absence of such couplings.

The (003) state belongs to the sequence of the  $(00v_3)$  states corresponding to the excitation of the OH stretch, which is expected to be isolated from the other vibrational states. The (004), (005), (006) and  $(007)^{4,21}$  states could also be considered as well isolated. Note, however, that the observation of some line-resonance partners from dark upper states has revealed noticeable perturbations affecting both the (004) and the (005) states. The (003) state with a band origin at  $10\,631.6832$  cm<sup>-1</sup> can also be treated as a nearly isolated state. We could reproduce 160 of the 172 observed energy levels with a rms deviation of  $0.006 \text{ cm}^{-1}$  by using an effective rotational Hamiltonian written using Padé-Borel approximants<sup>22,23</sup> and by adjusting 16 parameters. These parameters are listed in Table 4, while the (obs. - calc.) values of the energy levels are given in Table 2. The comparison included in Table 4 with the parameters previously obtained show a very good agreement (except for the vibrational term) with ref. 14, but significant differences with ref. 17. This disagreement can be explained by the fact that only 13 low  $[JK_aK_c]$  energy levels were obtained and included in the fit of ref. 17 and by the fact that three of these levels, [313], [532] and [524], were found to have errors as large as 1.07 cm<sup>-</sup>

Overall 13 energy levels of dark states were derived through transitions borrowing their intensities *via* resonance interaction with the rotational levels of the bright (003) state. They belong to the following vibrational states: (320) (9 levels), (131) (2 levels), (051) (1 level), and (080) (1 level) (see Table 3). Interestingly the (003) energy level shifts, caused by all these resonance interactions, are very small: all the interacting energy levels could indeed be included in the EH fit, except the (003) [716] level at 11 092.0415 cm<sup>-1</sup> which deviates from its unperturbed position by  $-0.068 \text{ cm}^{-1}$  due to interaction with the (320) [716] level at 11 092.5192 cm<sup>-1</sup>, observed through five lines. Indeed, this interaction induces a strong intensity transfer (up to 50%!) to the extra lines of the (320) upper state.

In contrast to the (003) state, the pure OD stretching (400) state at 10 378.9511 cm<sup>-1</sup> was found to be strongly perturbed and could not be fitted as an isolated state. Already at J = 2, all energy levels with  $K_a = 2$  interact strongly with the corresponding  $K_a = 2$  levels of the (051) state. As J increases, the energy levels with  $K_a = 1$  gradually become involved in the resonance. This interaction results in the observation of eleven energy levels of the (051) state. A further

**Table 2** Rotational energy levels (in  $cm^{-1}$ ) of the (003), (013), (400), and (112) vibrational states of HDO<sup>*a*</sup>

	(003)					(013)				(400)			(112)				
$JK_aK_c$	$E_{\rm obs}/{\rm cm}^{-1}$	σ	Ν	δ	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ
0 0 0	10631.6832	0.5	2	-5	-23	11969.7584		1	-30	10 378.9511		1	31	11 315.4333		1	0
101	10646.9209	0.7	4	-5	-23	11 985.0655	0.2	2	-30	10 393.4742	2.5	2	31	11 330.6251	3.5	2	-1
111	10658.2329	0.2	4	1	-23	11 997.7239	0.3	2	-29	10 408.2096	0.3	2	29	11 345.1810		1	-2
110	10661.1253	0.2	3	2	-23	12 000.8530	0.2	2	-29	10410.7151	1.4	3	30	11 348.2613	4.0	1	-1
202	10676.9016	0.5	4	-5	-24	12015.1657	0.4	2	-30	10 422.1939	0.4	2	31	11 360.5839 I	4.0	2	-2
212	10 685.8082	0.3	5	0	-23	12025.16/1	0.1	2	-24	10434.6930	0.7	1	29	11 3/2.5/45 1	2./	2	-3
211	10 094.4/82	0.3	0	3	-23	12034.3387	0.3	3 2	-27	10 442.1948	0.7	2	12	11 381.7429	6.0	2	-2
221 220	10 728.2709	0.2	4	0	-22	12072.0148	0.5	2	-30	10 486 8462	0.5	2	13	11 425.9575 11 426 3031 I	0.0	1	-1
303	10 720 7236	0.4	4	_7	-22	12075.1205	0.4	3	-29	10464 5207	0.5	3	31	11 404 4839	0.4	2	_3
313	10 726 8779	0.9	7	-/	-24	12 059.1114	0.7	2	-19	10404.3207	0.4	3	29	11 413 0864	1.2	2	-5
312	107441453	0.2	5	4	-23	12,084,6845	0.4	2	-14	10 489 1680	1.0	3	29	11 431 6699	0.6	2	-2
322	10773.9442	0.2	6	5	-23	12118.4943	0.3	3	-30	10 530.8539	0.7	2	7	11 471.6485 I	7.8	2	$-2^{-2}$
321	10776.2778	0.3	7	2	-23	12120.9352	0.8	3	-30	10 532.2193	0.7	2	8	11 473.7583	9.1	2	0
331	10837.3960	0.5	4	4	-21	12189.2570	0.2	2	-27	10 602.5402	0.3	3	22	11 555.6234	7.6	2	-2
330	10837.4472	1.1	6	4	-21	12189.3079	0.0	2	-28	10 602.5574	2.2	2	22	11 555.6035		1	-4
404	10777.4061	0.6	6	-4	-23	12115.8496	0.5	2	-30	10 519.7165	0.2	2	30	11461.3680	1.8	2	-2
414	10781.1586	0.3	6	-1	-23	12120.2967	0.2	2	-20	10 526.6379	2.2	3	30	11467.2098	0.7	2	-2
413	10809.6231	0.4	5	6	-24	12151.2261	0.3	2	-22	10 551.4045	0.4	4	27	11 497.6150	1.5	2	-2
423	10834.3934	0.3	7	-15	-23	12179.2353	0.4	3	-30	10 581.2398	1.0	2	17	11 532.2232 I	4.8	2	-2
422	10840.9191	0.5	7	2	-23	12186.0546	0.3	3	-30	10 585.7546	1.0	2	20	11 538.1657		1	-2
432	10899.0329	0.4	6	4	-22	12251.2185	0.5	3	-28	10 660.5294	0.3	2	23	11617.6196	3.3	2	0
431	10899.3878	0.4	7	4	-21	12251.5744	0.3	2	-27	10 660.6731	1.0	2	22	11617.8379	3.6	2	0
441	10985.9618	3.0	5	-6	-22	12 347.9676	4.4	2	-25	10764.6908	6.0	2	16	11739.5909 I		1	-2
440	10985.9630	2.4	4	-10	-22	12 347.9748	4.6	2	-25	10764.6814	1.5	2	16	11739.5890 I		1	-2
505	10846.2961	0.5	7	-4	-24	12184.6684	0.5	2	-31	10 587.1330	0.4	2	30	11 530.4815	2.5	2	-2
515	10848.3666	0.4	7	-1	-24	12187.1611	0.0	2	-30	10 591.7025	2.2	3	29	11 534.0419 I	9.6	2	-3
514	10890.1461	0.5	8	9	-25	12232.4550	1.0	4	-29	10 628.9840	1.4	3	23	11 578.9381		1	-3
524	10909.3412	0.5	7	-4	-23	12254.4885	1.2	3	-30	10 654.0890	0.9	2	26	11 607.3936 I	3.4	2	-2
523	10922.9291	0.3	8	2	-24	12268.7728	0.2	2	-31	10 663.2086	0.8	3	27	11 620.0915		1	-2
533	10976.1700	0.4	7	4	-22	12328.7706	0.8	4	-29	10733.0702	0.0	2	23	11 695.2210		1	0
532	10977.5350	0.6	4	4	-22	12 330.1801	0.4	2	-26	10733.6328	0.3	2	23				
542	11 062.9136	1.1	5	-1	-19	12425.3531	2.0	2	-26	10 836.8307	2.0	1	18				
541	11 062.9532	0.4	3	-2	-20	12425.3868	0.0	1	-25	10 836.8272	2.0	2	10	11077 2962		1	2
550	11173.4008	1.0	2	2	-18	12 547.4251	0.0	2	-24	10 968.9128		1	ð	119//.2803		1	-2
550	10.027.1046	0.5	7	5	-10	12 347.4237	0.0	2	-24	10 908.9128	0.2	2	20	11 611 4580		1	-2
616	10.927.1940	0.5	8	-3 _4	-23 -24	12 205.5145	0.5	1	-31	10 669 2193	1.1	2	20	11 613 4937	3.0	2	_2
615	10 984 6911	0.5	8	19	-25	12 200.0220	13	2	-30	10 715 8139	1.1	2	28	11 674 7787	5.0	1	_2
625	10 998 3525	0.7	4	_3	-24	12 343 8619	0.4	2	-30	10 739 7832	3.1	3	28	11 696 8059	15	2	_2
624	11 021 9768	0.7	6	1	-25	12 368 7801	0.2	2	-31	107559417	1.0	2	29	117194132	49	2	-3
634	11 068.6870	0.4	6	3	-23	12421.7972	2.1	3	-29	10 820.1242	0.1	2	23	11788.4278		1	1
633	11072.5208	0.3	7	3	-23	12425.6718	1.1	3	-30	10 821.7540	0.6	2	23	11791.1673		1	-2
643	11155.4660	0.8	6	-1	-21	12 518.4365	0.7	2	-24	10923.4723		1	20				
642	11155.6626	0.6	5	0	-21	12518.6461	0.7	3	-24	10923.5267	3.8	2	16				
652	11 265.4731	1.2	6	1	-19	12639.9821	2.4	3	-24	11 055.0393		1	8				
651	11 265.4734	1.1	4	-3	-19	12639.9840	2.3	3	-25	11 055.0462		1	9				
661	11 398.7914	0.4	4	1	-16	12786.6845		1	-21	11 214.5218	0.6	2	0				
660	11 398.7912	0.3	3	1	-16	12786.6846		1	-21	11 214.5218	0.6	2	0				
707	11 020.1461	0.3	4	7	-24	12357.8145	1.1	2	-32	10757.3682	1.0	3	29	11704.2390		1	-3
717	11 020.6581	0.5	6	0	-23	12358.4684	1.7	2	-31	10759.0413	1.4	3	28	11705.3647		1	-3
716	*11 092.0415	0.6	6	-68	-43	12436.3499	0.9	2	-24	10820.8888	0.7	3	31	11784.1163		1	-2
726	11 101.0149	0.5	7	-2	-24	12446.7815	0.8	2	-31	10838.7283	0.5	2	23	11800.1176 I		1	-1
725	11137.3365	0.6	5	-1	-25	12485.3411	1.2	2	-33	10864.0910		1	30	11835.5169		1	-2
735	11176.3410	0.4	5	3	-24	12 530.0501	1.4	3	-30	10921.6150	1.0	2	23				
734	11 184.9815	0.7	7	1	-24	12 538.8208	0.6	3	-31	10925.4328		1	24				
744	11 263.6688	0.6	5	0	-22	12627.2236	•	1	-27	11024.7070	0.9	3	17				
743	11 264.3604	0.7	5	0	-22	12628.0762	2.0	3	-25	11024.8982	2.2	2	17				
153	11 3/3.0665	1./	3	-3	-20	12 /48.1354		1	-23	11 155.5/94		1	/				
152	11 3/3.0913	1./	3	-2	-20	12/48.1652		1	-25	11 155.6092		1	9				
762	11 505.8706	0.8	5 1	2	-16	12894.3257		1	-22	11 314.4874		1	0				
/01 771	11 202.8094	1.8	4	1	-10	12094.323/		1	-22	11 514.48/4		1	0				
//I 770	11 001.1338	0.7	5	-1 1	-12	13064.2464		1	-1/ 17	11 500.0005		1	-/7				
2 N 0	11 1001.1008	0.7	5	-1	-12	13004.2404		1	-1/	11 300.0003	0.5	1	-/ 20	11 800 0021		1	5
0 U 0 8 1 9	11 123.1840	0.4	5 Д	-3 _1	-20 -25	12402.2400	20	2	-31	10 861 1082	0.5	∠ 2	∠ð 27	11 000.0004		1	-3 _1
817	11 125.4517	0.2	+	_13	-25 -25	12 402.3003	2.7	∠ 1	_31	10 937 7600	0.7	∠ 2	30	11 007.3033		1	_4 _1
01/	11211.3/3/	0.0	U	-13	-23	12 330.4370		1	51	10731.1009	0.0	4	50	11,000.9700		1	-4

Table 2(continued)	Fable 2	(continued)
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	(003)					(013) (4			(400)				(112)				
$JK_aK_c$	$E_{\rm obs}/{\rm cm}^{-1}$	σ	Ν	δ	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	Ν	Δ
827	11216.9224	0.7	7	0	-25	12 563.0174	1.6	2	-30	10950.1189	1.5	2	28	11916.9268		1	-1
826	11 268.0330	0.9	8	-3	-27	12 617.4500	0.6	2	-33	10987.3605		1	31	10.000 5000			
836	11 298.7778	0.4	6	4	-24	12 653.1799	0.7	2	-31	11 037.2592	3.6	2	23	12 020.7880		1	-2
833	11 31 5.2098	0.5	6	-2	-20 -23	12 009.9037	0.4	2	-32 -28	11 045.1047	0.8	2	20 18	12033.8027		1	-2
844	11 389.4571	0.5	6	0	-23	12753.2785	3.6	2	-30	11 141.0908	0.0	1	19				
854	11 496.2545	0.9	5	-4	-19	12871.9621	2.6	2	-28	11 270.6271		1	11				
853	11 496.3511	0.9	7	-4	-21	12872.0654	0.8	2	-27	11 270.6340		1	11				
863	11 628.3688	0.5	3	4	-17	13 017.4798		1	-21								
802	11 028.3093	0.9	3		-18 -14	13 017.4793		1	-22 -21								
871	11 783.0788	1.4	4	-4	-14	13 186.8068		1	-21								
881	11959.3797	0.7	2	3	-10												
880	11959.3797	0.7	2	3	-10												
909	11 242.3879	0.3	5	1	-26	12 578.6638		1	-33	10974.4522	0.4	2	28	11925.4476		1	-7
919	11 242.5058	0.6	5 5	5	-26	125/8.814/	37	1	-33	109/5.6009	0.5	2	25 34				
928	11 342.7240	0.9	5	-0 -1	-20 -26	12 692 0701	0.2	2	-30	11004.2217		1	54				
927	11 412.9010	0.1	4	-4	-27	12 763.9101	0.2	1	-34	11 125.2122		1	33	12046.8614		1	$^{-2}$
937	11 435.5609	0.3	6	3	-26	12 790.8031	3.0	2	-27	11 166.7731		1	25				
936	11 462.9235	0.9	6	-1	-27	12818.8776	2.7	2	-33	11 181.0421	0.0	2	28				
946	11 526.8300	0.9	5	2	-24	12 891.6368		1	-30	11 272 2101		1	20				
945	*11.635.0565	1.0	8 2	-2	-25	12 896.0611		1	-30	112/2.3191		1	20				
955	*11635.3744	1.4	3	-30 -42	-21 -21	13 011.8229	6.7	2	-29 -29								
964	11 766.3313	1.2	3	8	-19	13 156.1551	017	1	-23								
963	11766.3395	3.8	2	5	-20	13 156.1556		1	-25								
973	11920.3366	2.2	5	-1	-15	13 324.7837		1	-22								
972	11 920.3371	2.7	5	-1	-15	13 324.7836		1	-22								
982	12 096.0241	0.3	2	-5 -5	-12 -12												
10 0 10	11 371.7562	1.3	5	-3	-12 -26	12 707.0907		1	-34	11 100.5294	1.2	2	27				
10 1 10	11 371.8055	1.7	2	0	-26	12 707.1610		1	-34								
10 1 9	11 485.5208	0.9	4	4	-26	12831.6381	0.9	2	-32	11 205.7091		1	30				
10 2 9	11 487.1040	1.5	3	-2	-27	12833.6803	1.8	2	-33	11211.1364		1	30				
10 2 8	11 570.7061	1.1	4	-2	-27	12 923.4168	0.0	2	-35	11 276.8914		1	33				
10 3 8	11 580.2178	1.0	3	4 _1	-20 -28	12 942.1893		1	-35 -35	11 510.5055		1	27				
10 4 7	11 681.4159	0.9	3	2	-25	13 047.1120		1	-34								
10 4 6	11 691.1102	0.8	4	-4	-27	13 056.8321		1	-31								
10 5 6	11789.5943	0.5	5	8	-24	13 166.8001		1	-32								
10 5 5	11 790.5113	1.2	3	7	-23												
10 6 3	11919./810	2.3	3	_12	-20 -20												
10 7 4	12072.9418	0.3	3	18	-17												
1073	12072.9401	1.6	4	15	-17												
11 0 11	11 513.2824		1	3	-27	12 847.5269		1	-34	11 238.4974		1	27				
11 1 11	11 513.2931	6.0	2	-10	-30	12 847.5714		1	-34	11 235.0284		1	25				
11 1 10	11 640.0636	0.5	4	-3	-27	12 986.512/		1	-34								
11 2 10	11 740.4016	0.4	3	-2	-29	13 094.8466		1	-35								
11 3 9	11 750.2738	2.1	4	3	-28	13 107.1285		1	-32								
11 3 8	11 807.3350	1.9	4	-2	-29	13 166.5269		1	-27								
11 4 8	11 850.8996	0.6	3	3	-26	13 217.5995		1	-35								
1147	11 868.6056	0.8	4	-7	-28	13 235.1242		1	-37								
11 5 7	11959.0805	0.7	2	-11 -18	-25												
11 6 6	12 088.8061	7.3	2	7	-20												
11 6 5	12088.9240	2.9	2	-18	-22												
11 7 5	*12240.9120		1	47	-18												
1174	*12 240.9017		1	31	-18												
1184	12414.9341		1	-2 _2	-14 -14												
12 0 12	11 666.9441	3.3	2	-2	-29	12999.9345		1	-37								
12 1 12	11 666.9605	4.2	2	8	-28	12999.9622		1	-36								
12 1 11	11 806.4885	0.3	5	1	-28	13 153.1689		1	-33								
12 2 11	11 806.9018	0.4	2	5	-28	13 153.7202		1	-35								

Table 2 (c	ontinued)
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	(003)			(013)				(400)				(112)					
$JK_aK_c$	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	δ	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	Ν	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ
12 2 10	11921.4033	0.6	2	-1	-29	13 277.4641		1	-37								
12 3 10	11927.3058	1.1	2	7	-28												
1239	12001.6749	1.6	2	-1	-30												
1249	12034.8363	1.9	2	8	-28												
12 4 8	12063.8227		1	6	-29												
12 5 8	*12145.2318	1.6	2	-37	-26												
12 5 7	*12150.2251	4.4	2	-55	-27												
12 6 7	*12273.3380		1	-34	-23												
13 0 13	11832.7179	0.1	3	2	-29	13164.0834		1	-34								
13 1 13	11832.7179	0.0	2	-2	-29	13164.3547		1	-36								
13 1 12	11984.8520		1	10	-29	13331.3642		1	-36								
13 2 12	11985.0419	0.1	2	-1	-28												
13 2 11	12113.5813		1	-1	-29												
13 3 11	12116.9235		1	-11	-31												
13 3 10	12209.0071	2.1	2	-2	-31												
13 4 10	12232.7217		1	11	-29												
1349	12275.9915		1	11	-30												
14 0 14	12010.5785		1	11	-31	13 340.6434		1	-37								
14 1 14	12010.5545		1	-10	-29	13 340.6674		1	-37								
14 1 13	*12174.6229		1	-528	-23			-									
14 2 13	12175.2537	1.1	2	5	-28												
14 2 12	12317.0492		1	_9	-30												
14 3 12	12318.8821		1	-15	-32												
14 3 11	12428.1375		1	-4	-30												
14 4 11	12444.0427		1	6	-33												
14 4 10	12 503 9766		1	2	-33												
15 0 15	12 200.4484		1	-5	-31												
15 1 15	12,200,4602		1	6	-30												
15 1 14	*12 377.3476		1	-59	-29												
15 2 14	*12 377 4911		1	37	-29												
15 2 13	*12 531 9700		1	-23	-33												
15 3 13	12 532 9819		1	5	-31												
15 3 12	126582709		1	-3	-32												
16 0 16	12.402.3362		1	-6	-32												
16 1 16	12.402.3310		1	-11	-32												
16 2 15	12 591 6144		1	3	-32												
17 0 17	12616 1902		1	3	-32												
17 1 17	12 616 1931		1	6	_34												
1/11/	12010.1991		1	0	<i>J</i> -r												

<sup>*a*</sup> N is the number of lines used for the upper energy level determination and  $\sigma$  denotes the corresponding experimental uncertainty in 10<sup>-3</sup> cm<sup>-1</sup> units.  $\Delta$  is the difference between the experimental energy level and the predictions of ref. 15 in 10<sup>-2</sup> cm<sup>-1</sup> units. For the (003) state, we also give the difference,  $\delta$ , between the experimental and calculated (EH) energy level in 10<sup>-3</sup> cm<sup>-1</sup> units (\* denotes energy level excluded from the EH fitting). A number of energy levels of the (112) state (marked by I) were first determined by ICLAS in ref. 13.

example of strong perturbation concerns the [918] (400) level at  $11064.2217 \text{ cm}^{-1}$  coupled with a level at  $11067.8104 \text{ cm}^{-1}$  (mis)assigned to the same rovibrational state by SP. Using the EH approach, we could assign this level, derived from three weak observed lines, to the [918] level of the highly excited pure bending state (080). The (400) state seems to be also perturbed by other dark states like (150) and (131), which makes accurate EH fitting problematical.

Rotational energy levels of the (032) and (112) states were derived in ref. 13 from hot transitions from the (010) state observed in the 9600–10 100 cm<sup>-1</sup> spectral region. In the considered spectral region, weak transitions belonging mostly to the (112)–(000) band were assigned, and 43 extra levels could be determined. The (032) and (112) states are linked by a strong Fermi-type resonance interaction, leading to a nearly equal mixing of high  $K_a$  sublevels and thus to an ambiguity in the vibrational assignment. An rms of 0.022 cm<sup>-1</sup> could be achieved for 84 energy levels (six levels excluded) with 18 varied parameters (not given), but the influence of other dark states on the (032)–(112) dyad hampered a satisfactory EH fitting.

The (013) state interacts mainly with the (330) and (141) dark states. As the vibrational energies of the (013) and (330) states are close - 11969.76 and 11958.26 cm<sup>-1</sup> respectively - the interaction is already important for small J and  $K_a$  values. This interaction induces relatively strong (up to  $1.1 \times 10^{-4} \text{ cm}^{-2} \text{atm}^{-1}$ ) transitions to the  $K_a = 1$  rotational levels allowing for the retrieval of nine experimental energy levels for the (330) state. We have already stressed the overall good agreement between the SP calculated and experimental intensities, but in the specific case of the (013)-(330) interaction, the calculated intensities of both the (330)-(000) transitions and the corresponding resonance-partners of the (013)-(000) band seem to be strongly distorted. For example, the calculated intensities are 2.5-3 times underestimated for transitions to the [313] (330) upper level, while those for transitions sharing the [312] (330) level seem to be overestimated by a factor of 2.

The resonance interaction between the  $K_a = 4$  energy levels of the (013) state and those of the (141) state at 11 804.57 cm<sup>-1</sup> is very strong. Seven energy levels of the (141) state were derived through transitions borrowing their intensities from

Table 3 Rotational energy levels (cm<sup>-1</sup>) of the vibrational states of HDO derived in the 10110–12215 cm<sup>-1</sup> spectral region<sup>a</sup>

$\overline{J K_a K_c}$	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ	$J K_a K_c$	$E_{\rm obs}/{\rm cm}^{-1}$	σ	N	Δ	$J K_a K_c$	$E_{\rm obs}/{\rm cm}^{-1}$	σ	Ν	Δ
	(032)					(051)					(141)			
0.0.0	11 242.923 I		1	-14	2.2.1	10476.2865		1	4	441	12 344 8847		1	-14
101	11 258.346 I	3	2	-13	$\frac{2}{2}$ $\frac{2}{2}$ $\frac{1}{0}$	10476.5950	5.0	2	4	440	12 344 8828		1	-15
111	11276.3265 F		1	-13	322	10 521 9172		1	11	541	12,423,1458		1	-16
110	11 279.734 I		1	-15	321	10 523.4100	1.1	3	11	633	12 340.9520		1	-14
202	11 288.739 I	1	2	-14	423	10 590 5170		1	0	643	12,516,8737		1	-15
212	11 303.737 I	_	1	-14	422	10 594, 5649	1.0	2	-2	642	12 517.1675		1	-17
2 1 1	11 313.9513 IF		1	-15	514	10 620.4112	0.6	2	14	743	12 627.0978	6.2	2	-23
2 2 0	11 367.001 I	4	2	-14	524	10665.9336	0.8	2	-6		(061)			
303	11 333.250 I	2	2	-15	523	10675.1175	0.0	1	-6	5 4 2	(001)	2.2	2	0
313	11 344.603 I	2	3	-14	615	10724.4236		1	7	342	12 330.0703	2.3	3	0
312	11 364.961 I		1	-14	716	10837.5803		1	6		(080)			
3 2 2	11412.640 I	4	2	-14	735	11019.2934	4.5	2	-2	624	10929.6690	0.6	2	1
3 2 1	114150418 IF	1	2	-15		(220)				918	11067.8104	2.5	3	-9
331	11 503 428 I		1	-15	212	(320)		1	27		(211)			
404	11 390 832 I	2	2	-15	212	10.602.2028		1	27	0.0.0	10/03 1/22		1	7
414	11 398 624 I	2	1	-15	211	10 721 8179		1	26	101	10/17 0828	5.1	3	- /
413	11 432 263 I		1	-14	312	10741 9097	0.3	2	20	111	10433 9840	1.4	2	-0
422	11 479.879 I		1	-14	413	10 807 3968	0.5	1	28	110	10436 8191	1.7	1	-5
515	11465.5211 IF	1	2	-15	514	10 888 1707		1	20	202	10447 3050	35	2	-5
514	11 515.150 I	-	1	-15	615	10 983 4895		1	26	212	104608507	5.5	1	-6
524	11 549.348 I		1	-15	716	11 092 5192	19	5	43	212	10469 3741	12	2	-5
523	11 562.525 I		1	-13	817	11 214 0069	1.5	1	23	2 2 1	10 516 7938	1.2	1	_7
533	11 642.629 I		1	-3		11211.0009				303	104904529	0.1	2	-5
606	11 542 105 I	2	2	-14		(330)				331	10.634.6125	0	1	-8
616	11 545.028 I	1	2	-14	212	12026.5502		1	18	404	10 546.6118	0.7	2	-5
615	11612.735 I	-	1	-13	2 1 1	12036.3374		1	22	414	10 554.0936	0.1	2	-5
625	11 639.292 I	9	2	-15	3 1 3	12065.6916	0.1	2	12	413	10 582.7202	3.1	2	-1
624	11 662.743 I		1	-16	312	12085.6479	0.8	2	10	505	10615.0610	0.8	2	-5
634	11735.026 I		1	5	414	12118.2064	6.5	2	22	515	10620.0143	1.9	2	-6
707	11 635.213 I	3	2	-15	413	12150.3998	0.6	2	16	606	10 695.4058		1	-6
717	11 636.891 I		1	-16	514	12231.0306	2.7	2	21	616	10698.5367	0.5	2	-3
726	11 743.065 I		1	-14	615	12 326.2977		1	21	707	10787.5587		1	-5
725	11 780.050 I	9	2	-13	716	12435.5690	1.5	2	13	717	10789.4442		1	-5
808	11 740.001 I		1	-17		(102)					(131)			
818	11 740.997 I	5	2	-14	070	(102)		1	(	2.2.0	(131)	4.0	•	
909	11 856.590 I	5	2	-15	8/2	11 137.7422		1	-0	330	10 / /2. /864	4.2	2	-11
919	11 857.248 I		1	-11	8/1	11 137.7422		1	-0	431	10835.2885	1.9	6	-11
	(022)		-		881	11 320.0453		1	-3					
	(022)			-	000	11 320.0433		1	-3					
1156	11 399 1247		1	-8										

<sup>*a*</sup> *N* is the number of lines used for the upper energy level determination and  $\sigma$  denotes the corresponding experimental uncertainty in 10<sup>-3</sup> cm<sup>-1</sup>.  $\Lambda$  is the difference between the experimental energy level and the prediction of ref. 15 in 10<sup>-2</sup> cm<sup>-1</sup> units. For the (032) state only, we merge the levels presently determined from FT data (marked with *F*) with those obtained by ICLAS-VeCSEL<sup>13</sup> (marked with *I*). Note in particular that many levels of the (022)–(102) dyad were determined in ref. 13 but are not listed here.

the (013)-(000) lines. Though the (013) energy levels with  $K_a = 3$  are very close to those of the (061) state with  $K_a = 4$ , perturbations were not evident, neither for line positions nor for intensities of the (013)-(000) transitions. However, the resonance link between the [532] (013) level at 12 330.1801  $cm^{-1}$  and the [542] (061) level at 12 330.0703  $cm^{-1}$  induces a remarkable intensity transfer to weak transitions of the (061)-(000) band. The [524] (061) energy level was observed through three strong lines (see Table 5) while SP intensities were calculated more than 10 times weaker than observed. This is the only case where we found such a significant failure in SP intensity calculations for HDO. The resonance interactions of the (013) state with the (330) and (141) states are strongly dominant: we could indeed reproduce 133 energy levels of the (013)–(330)–(141) triad within  $0.021 \text{ cm}^{-1}$  by varying 26 parameters and excluding six levels. Ideally, other interactions with dark states should be taken into account to approach the experimental accuracy of the energy levels.

We have gathered in Table 6 the high-order resonance interactions in HDO which were evident in our work, together with those previously reported. This table shows that these interactions couple mostly vibrational states with a large difference in the bending quantum number  $v_2$ . This characteristic interaction is the result of a strong centrifugal distortion effect in HDO, which is induced by a high amplitude bending motion near the linear configuration. Similar effects were discussed in detail for the main isotopomer H<sub>2</sub>O.<sup>24</sup>

### V. Conclusion

Experimental and theoretical study of the HDO FT absorption spectrum in the wide spectral region 10110–12215 cm<sup>-1</sup> resulted in the derivation and assignment of a large dataset of rovibrational energy levels. The HDO synthetic spectrum of Schwenke and Partridge<sup>15,16</sup> appeared to be accurate enough to provide an overall complete spectral assignment. Calculations using the effective Hamiltonian approach performed in parallel helped us to understand the nature of resonance interactions in HDO and to confirm most SP assignments. The analyzed rovibrational energy levels of HDO were found to undergo strong resonance interactions inducing

 
 Table 4
 Rotational and centrifugal distortion constants of the (003)
vibrational state of HDO (in cm<sup>-1</sup>) and a comparison with previous studies

	This work	Ref. 14	Ref. 17
$E_{v}$	10 631.68900(200)	10 631.6366(83)	10631.681(4)
A	20.381736(320)	20.3814(21)	20.370(4)
В	9.069274(140)	9.06731(89)	9.093(2)
С	6.1700468(990)	6.17030(22)	6.149(1)
$\Delta_k$	$9.5005(150) \times 10^{-3}$	$9.31(11) \times 10^{-3}$	$2.1(9) \times 10^{-3}$
$\Delta_{ik}$	$5.4727(610) \times 10^{-4}$	$6.44(65) \times 10^{-4}$	$-34(2) \times 10^{-4}$
$\Delta_i$	$3.96914(710) \times 10^{-4}$	$3.599(82) \times 10^{-4}$	$9.9(6) \times 10^{-4}$
$\delta_k$	$1.7254(100) \times 10^{-3}$	$1.507(42) \times 10^{-3}$	$1.1(8) \times 10^{-3}$
$\delta_i$	$1.40648(490) \times 10^{-4}$	$1.239(41) \times 10^{-4}$	$2.7(2) \times 10^{-4}$
$H_k$	$3.1672(260) \times 10^{-5}$	$3.39(19) \times 10^{-5}$	
$H_{ki}$	$-1.0015(150) \times 10^{-5}$	$-1.40(13) \times 10^{-5}$	
$H_{ik}$	$2.2330(540) \times 10^{-6}$	$2.4 \times 10^{-6}$	
$H_i$	$4.044(170) \times 10^{-8}$	$3.9 \times 10^{-8}$	
$h_k$	$1.4365(500) \times 10^{-5}$	$1.8 \times 10^{-5}$	
$h_{ki}$	$9.583(450) \times 10^{-7}$	$10.0 \times 10^{-7}$	
$h_j$	$1.863(120) \times 10^{-8}$	$2.1  imes 10^{-8}$	
<sup>a</sup> Valu	ues in parentheses are the	$1\sigma$ statistical error int	ervals.

 
 Table 5
 Comparison of the observed and calculated<sup>16</sup> line intensities
of transitions to the (061) [542] upper state

J K <sub>a</sub> K	2	Observed	Intensity/cr	Linner lavel /			
Upper	Lower	cm <sup>-1</sup> cm <sup>-1</sup>	SP <sup>15,16</sup>	Observed	cm <sup>-1</sup> cm <sup>-1</sup>		
5 4 2 5 4 2 5 4 2	6 3 3 5 3 3 4 3 1	11 860.4100 11 956.4026 12 034.3914	$\begin{array}{c} 1.4 \times 10^{-6} \\ 1.6 \times 10^{-6} \\ 1.9 \times 10^{-6} \end{array}$	$\begin{array}{c} 3.2 \times 10^{-5} \\ 2.0 \times 10^{-5} \\ 2.8 \times 10^{-5} \end{array}$	12 330.0735 12 330.0684 12 330.0689		

Table 6 Summary of the high-order resonance couplings evident in HDO

Resonance	Interacting states	
partner of $v_1 v_2 v_3$	<i>v</i> <sub>1</sub> <i>v</i> <sub>2</sub> <i>v</i> <sub>3</sub> - <i>v</i> ' <sub>1</sub> <i>v</i> ' <sub>2</sub> <i>v</i> ' <sub>3</sub>	Ref.
$v_1 \mp 4 \ v_2 \ v_3 \pm 3$	(411)–(014)	8
$v_1 \mp 3 \ v_2 \pm 1 \ v_3 \pm 2$	(302)–(014)	8
$v_1 \pm 3 \ v_2 \pm 2 \ v_3 \mp 3$	(003)–(320)	This work
	(013)–(330)	This work
$v_1 \pm 5 \ v_2 \mp 2 \ v_3 \mp 3$	(024)–(501)	11
$v_1 \pm 1 \ v_2 \pm 3 \ v_3 \mp 2$	(014)–(142)	8
	(023)–(151)	9
	(003)–(131)	This work
	(013)–(141)	This work
$v_1 \mp 2 \ v_2 \pm 4 \ v_3$	(200)–(040)	25
	(210)–(050)	26,27
$v_1 \mp 4 \ v_2 \pm 5 \ v_1 \pm 1$	(400)–(051)	This work
$v_1 v_2 \pm 5 v_3 \mp 2$	(004)–(052)	7
	(015)-(063)	11
	(013)-(061)	This work
	(003)-(051)	This work
	(014)–(062)	8
$v_1 \pm 2 \ v_2 \pm 6 \ v_3 \mp 4$	(015)-(271)	11
$v_1 \mp 3 \ v_2 \pm 6 \ v_3$	(300)–(060)	28
$v_1 v_2 \pm 6 v_3 \mp 2$	(022)–(080)	13
$v_1 \mp 1 \ v_2 \pm 8 \ v_3 \mp 2$	(142)-(0 12 0)	8
$v_1 v_2 \pm 8 v_3 \mp 3$	(003)–(080)	This work
$v_1 \pm 1 \ v_2 \pm 9 \ v_3 \mp 4$	(024)-(1 11 0)	11

considerable intensity redistribution between line-resonance partners. High-order resonance interactions coupling the  $(v_1, v_2, v_3)$  bright state to the  $(v_1 \pm 1, v_2 \pm 3, v_3 \mp 2)$  and  $(v_1, v_2 \pm 5, v_3 \mp 2)$  dark states were frequently observed in HDO both in the present analysis and in previous investigations.

The analysis of the observed spectrum illustrated the limits of both theoretical approaches used. The effective Hamiltonian approach could not be applied in the case of large sets of interacting states in particular when the experimental energy levels were unknown for most of the coupled states. On the other hand, SP calculations failed to reproduce properly the intensities of some strongly perturbed transitions. We hope that the results obtained in this paper will help for further development of theoretical models.

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### References

- R. A. Toth, J. Mol. Spectrosc., 1997, 186, 66-89. 1
- R. A. Toth, J. Mol. Spectrosc., 1997, 186, 276-292. 2
- 3 R. A. Toth, J. Mol. Spectrosc., 1999, 195, 73-97.
- 4 A. Jenouvrier, M. F. Merienne, M. Carleer, R. Colin, A.-C. Vandaele, P. F. Bernath, O. L. Polyansky and J. Tennyson, J. Mol. Spectrosc., 2001, **209**, 165–168. X. Wang, S. He, S. Hu, J. Zheng and Q. Zhu, Chin. Phys., 2000,
- 5 09 885-891
- S. Hu, S. He, J. Zheng, X. Wang, Y. Ding and Q. Zhu, Chin. 6 Phys., 2001, 10, 1021-1027.
- 7 O. Naumenko, E. Bertseva and A. Campargue, J. Mol. Spectrosc., 1999, 197, 122-132
- O. Naumenko and A. Campargue, J. Mol. Spectrosc., 2000, 199, 8 59-72.
- 9 O. Naumenko, E. Bertseva, A. Campargue and D. Schwenke, J. Mol. Spectrosc., 2000, 201, 297-309.
- 10 E. Bertseva, O. Naumenko and A. Campargue, J. Mol. Spectrosc., 2000, 203, 28-36.
- 11 A. Campargue, E. Bertseva and O. Naumenko, J. Mol. Spectrosc., 2000, 204, 94-105.
- S. Hu, H. Lin, S. He, J. Cheng and Q. Zhu, Phys. Chem. Chem. 12 Phys., 1999, 1, 3727-3730.
- 13 E. Bertseva, O. Naumenko and A. Campargue, J. Mol. Spectrosc., 2003. 221. 38-46.
- 14 A. D. Bykov, V. A. Kapitanov, O. V. Naumenko, T. M. Petrova, V. I. Serdyukov and L. Sinitsa, J. Mol. Spectrosc., 1992, 153, 197 - 207
- 15 H. Partridge and D. W. Schwenke, J. Chem. Phys., 1997, 106, 4618-4639.
- 16 D. W. Schwenke and H. Partridge, J. Chem. Phys., 2000, 113, 6592-6597
- O. Votava, J. R. Fair, D. F. Plusquellic, E. Riedle and D. J. 17 Nesbitt, J. Chem. Phys., 1997, 107, 8854-8865.
- 18 J. R. Fair, O. Votava and D. J. Nesbitt, J. Chem. Phys., 1998, 108, 72-80.
- A. Bykov, O. Naumenko, A. Pshenichnikov, L. Sinitsa and A. 19 Shcherbakov, Optics and Spectroscopy, 2003, 94, 528–537. R. Toth, J. Mol. Spectrosc., 1993, 162, 20–40.
- 20
- 21 S. S. Voronina, Atmos. Ocean. Opt., 2002, 15, 727-729.

- 1997, 183, 300-309.
- 24 A. D. Bykov, O. V. Naumenko, L. N. Sinitsa, B. Voronin, J.-M. Flaud, C. Camy-Peyret and R. Lanquetin, J. Mol. Spectrosc., 2001, 206, 1-8.
- 25 A. D. Bykov, B. A. Voronin, O. V. Naumenko, L. N. Sinitsa and T. M. Petrova, Atmos. Ocean. Opt., 1999, 12, 786-791.
- 26 A. D. Bykov, O. V. Naumenko, L. N. Sinitsa, B. P. Winnewisser, M. Winnewisser, P. S. Ormsby and K. Narahari Rao, SPIE, 1993, 2205, 248-252.
- O. N. Ulenikov, S. Hu, E. S. Bekhtereva, G. A. Onopenko, X. Wang, S. He, J. Zheng and Q. Zhu, *J. Mol. Spectrosc.*, 2001, 27 208, 224-235.
- 28 S. Hu, O. N. Ulenikov, G. A. Onopenko, E. S. Bekhtereva, S. He, X. Wang, H. Lin and Q. Zhu, J. Mol. Spectrosc., 2000, 203, 228-234.